



兰州大学  
LANZHOU UNIVERSITY

# Estimators and Scoring

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- It is often said that **Monte Carlo** (MC) is a “**mathematical experiment**”, so the MC equivalent of the result of a real experiment (*i.e.*, of a **measurement**) is called an **estimator**.
- Just as a real measurement, an estimator is obtained by sampling from a statistical distribution and has a **statistical error** (and in general also a **systematic one**).
- There are often several different techniques to measure the same physical quantity: in the same way the **same quantity can be calculated using different kinds of estimators**.
- FLUKA offers **numerous different estimators**, *i.e.*, directly from the input file the users can request scoring the respective quantities they are interested in.
- As the latter is implemented in a very complete way, users are strongly encouraged to **preferably use the built-in estimators** with respect to user-defined scoring.
- For additional requirements **FLUKA user routines** are provided.



- Several **pre-defined estimators** can be activated in FLUKA.
- One usually refers to these estimators as “**scoring**” capabilities.
- Users have also the possibility to build their own scoring through **user routines**,  
HOWEVER:
  - **Built-In scoring** covers most of the **common needs**
  - **Built-In scoring** has been **extensively tested**
  - **Built-In scoring** takes **BIASING weights automatically into account**
  - **Built-In scoring** has **refined algorithms** for track subdivision
  - **Built-In scoring** comes with **utility programs** that allow to evaluate statistical errors
- Standard scoring can be weighted by means of **simple user routines** (fluscw, comscw).



**What?** Energy deposition and derivatives (dose), fluence or current versus energy/angle/other kinematical variables, time, DPA,...

**Where?** In regions, on boundaries, on region independent grids

**When?**

At the end of each cycle

At each event

**Output?** Files `inpnnn_fort.##` where ## is logical unit number chosen by user

**Results?**

Postprocessing utilities (in `$FLUPRO/flutil`)  
merge cycles, calculate average and rms,  
provide data files for plotting

User code needed

Results normalized  
per primary

Data merging and plotting available in **FLAIR**



The FLUKA output consists of:

- A **main (standard) output**, written on logical output unit **LUNOUT** (predefined as 11 by default) **[\*.out]**
- A file with the last random number seeds, unit **LUNRAN** (2 by default) **[ran\*]**
- A file of error messages, unit **LUNERR** (15 by default) **[.err]**
- Any number (including zero) of **estimator output files**. Their logical unit number is defined by the user **[\_fort.xx]**
- The available range of logical output numbers is: **21-512**
- Generally, the user can choose between **formatted (ASCII)** and **unformatted (binary)** scoring (positive or negative sign in the logical unit number). **Unformatted scoring is mandatory for the use of provided post-processing utilities!**
- Several estimators can be output on the same file (same logical unit) **provided they are of the same type.**
- Possible **additional output generated by the user** in any user routine



## Definitions:

- $\lambda[\text{cm}]$ : **mean free path**, the average distance travelled by a particle in a material before an interaction. It depends on the material, particle type and energy.
- $\Sigma[\text{cm}^{-1}] = 1/\lambda[\text{cm}]$ : **macroscopic cross section**, the probability of interaction per unit distance. It depends on the material, particle type and energy.
- $N_0$ : **number of atoms in volume [atoms/(cm barn)]**
- $\sigma = \frac{\Sigma}{N_0}$  [barn =  $10^{-24}\text{cm}^2$ ]: **microscopic cross section** (atom effective area)
  - i) the **area of an atom weighted with the probability of interaction** (hence the name “cross section”)
  - ii) the **probability of interaction per unit length, with the length measured in atoms/cm<sup>2</sup>** (the number of atoms contained in a cylinder with a base of  $1\text{cm}^2$ )

Microscopic and macroscopic cross sections have a similar physical meaning of “probability of interaction per unit length”, with length measured in different units. Thus, the number of interaction can be obtained by multiplying both by the corresponding particle track-length.



More definitions:

- $R = N l \Sigma$ : The number of **reactions** for  $N$  identical particles along a total distance  $l$ .
- $\dot{R} = N \frac{dl}{dt} \Sigma = N v \Sigma$  [1/s]: **Reaction rate**, where  $v$  is the **average particle velocity**.
- $\frac{d\dot{R}}{dV} = \frac{dN}{dV} v \Sigma = n(\mathbf{r}, v) v \Sigma$  [1/(cm<sup>3</sup>s)]: **Reaction rate** inside a **volume element**  $dV$ , with  $n(\mathbf{r}, v)$  the density of particles with velocity  $v$  at position  $\mathbf{r}$ .
- $\dot{\Phi}(\mathbf{r}, v) = n(\mathbf{r}, v) v$  [1/(cm<sup>2</sup>s)]: **Fluence rate** or **Flux density**.
- $\Phi = \int \dot{\Phi}(\mathbf{r}, v) dt = n(\mathbf{r}, v) l$  [1/(cm<sup>2</sup>)]: **Fluence**, the time integral of the flux density.

Fluence is measured in **particles per cm<sup>2</sup>** but in reality it describes the **density of particle tracks**.



There are several ways to estimate the fluence:

- **Track length** estimation

$$\dot{\Phi} dt = n(v) v dt = \frac{dN(v)}{dV} \frac{dl(v)}{dt} = \lim_{\Delta V \rightarrow 0} \frac{\sum_i l_i(v)}{\Delta V}$$

Sum the length of all particle tracks inside a volume

**USRTRACK**

- **Collision density** estimation

$$\dot{\Phi}(v) = \frac{\dot{R}(v)}{\sigma(v) N_0 \Delta V} = \frac{\dot{R}(v)}{\Sigma(v) \Delta V} = \frac{\dot{R}(v) \lambda(v)}{\Delta V}$$

Count the collisions (weighted by the mean free path) inside a volume

**USRCOLL**





- **Surface crossing** estimation

- Imagine a surface having an infinitesimal thickness  $dt$ .

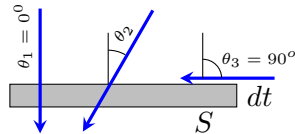
A particle incident with an angle  $\theta$  with respect to the normal of the surface  $S$  will travel a segment  $dt / \cos \theta$ .

- Therefore, we can calculate an average surface fluence by adding  $dt / \cos \theta$  for each particle crossing the surface and dividing by the volume  $S dt$ :

$$\Phi = \lim_{dt \rightarrow 0} \frac{\sum_i \frac{dt}{\cos \theta_i}}{S dt}$$

- While the **current**  $J$  counts the number of particles crossing the surface divided by the surface:

$$J = \frac{dN}{dS}$$



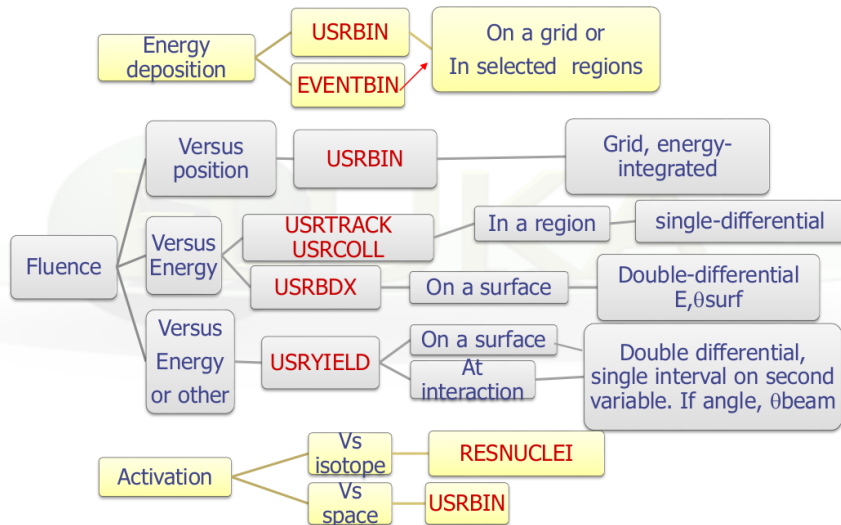


The **fluence is independent** from the orientation of surface  $S$   
while the **current is NOT!**

In an isotropic radiation field it can be easily seen that on a flat surface

$$J = \Phi/2$$

Current is *meaningful* if particles are only counted (e.g. signal trigger).  
But if someone wants to estimate **dose, activation, radiation damage or  
instrument response** the quantity to be used is **fluence**.



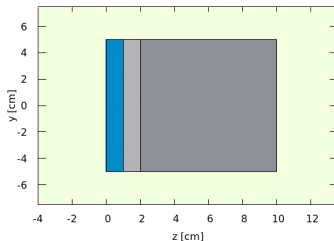


- **USRTRACK, USRCOLL** score average  $d\Phi/dE$  (differential fluence) of a given type or family of particles in a **given region**;
- **USRBDX** scores average  $d^2\Phi/dEd\Omega$  (double-differential fluence or current) of a given type or family of particles on a **given surface**;
- **USRBIN** scores the spatial distribution of deposited energy, total fluence, DPAs, activation, star density, momentum transfer or dosimetric quantities related to cell survival in a **regular mesh** (cylindrical, Cartesian oder per region) described by the user
- **USRYIELD** scores double differential yield of particles escaping from a **surface**. Distribution can be with respect to energy and angle (but also other quantities are possible);
- **SCORE** scores energy deposited (or star density) in all regions;
  - The output of SCORE will be printed in the **main (standard) output**, written on **main (standard) output**, written o **logical output unit LUNOUT** (pre-defined as **11** by default)
- Scoring detectors write their result into **logical output units assigned by the user** (unit numbers must be **>20**)



- **RESNUCLEI** scores residual nuclei in a given region (more details in lecture on activation);
- **DETECT** scores energy deposition in one or more regions in coincidence or anti-coincidence with a trigger separately for each “event” (primary history) to logical output unit 17;
- **EVENTBIN** works like **USRBIN**, but prints the binning output after each event (instead of averaging over event histories);
- **USERDUMP** defines the events to be written to a “collision tape” file;
- **AUXSCORE** defines filters (for particle types) and conversion coefficients (e.g. fluence-to-dose conversion coefficients);
- **DCYSCORE** assigns irradiation and cooling times (see lecture on activation);
- **ROTPRBIN** sets storage precision to single or double precision and assigns rotations/translations to **USRBIN** or **EVENTBIN** scorers (useful in case of LATTICES);
- **TCQUENCH** sets scoring time cut-offs and/or Birks quenching parameters for **USRBIN** or **EVENTBIN** binnings.

- Result of Geometry exercise: Cylinder along Z, filled by water-aluminum-lead and surrounded by CO<sub>2</sub>



- The **USRBIN** command allows to define a **3-D grid**, either cartesian (X-Y-Z), cylindrical (R-Z- $\Phi$ ) or by region.
- On this grid, one can score energy deposition, particle fluence (total or by particle type), as well as the density of interactions;
- There is an equivalent **EVENTBIN** command, that outputs the same quantities event-by-event;
- Using **USERWEIG** the results can be weighted by the **comscw.f** or **fluscw.f** external routines (Advanced!).

**WHAT(1):** code indicating the type of binning selected:

Example:  $\text{WHAT}(1) = 11.0 \rightarrow$  Mesh: R-Z or R- $\Phi$ -Z, no symmetry

**WHAT(2):** particle (or particle family) type to be scored

Examples: NEUTRON for neutrons, HAD-CHAR for charged hadrons

**WHAT(3):** logical output unit, ( $|\text{WHAT}(3)| > 21$  recommended)

$> 0$  formatted (human readable) output,

$< 0$  unformatted (machine readable) output

The first three fields have similar meanings for all estimators. The other **WHATs** contain the limits for regions or boundaries, and the number of intervals for cells/histograms (see manual).

For USRBIN with R- $\Phi$ -Z binning:

**WHAT(4):**  $R_{\max}$

**WHAT(5):** y coordinate of binning axis

**WHAT(6):**  $Z_{\max}$

**SDUM:** estimator name (optional)

Continuation card:

**WHAT(1,3):**  $R_{\min}, Z_{\min}$

**WHAT(2):** x coordinate of binning axis

**WHAT(4-6):** Number of bins in R,  $\Phi$ , Z

**SDUM:** & (continuation character)

- Start from the solution of **Geometry\_exercise** (either copy your `.inp` and `.flair` files and rename them to `example_score`, or download/copy the file `Geometry_exercise/geometry_solution.inp` and rename it):

```
mkdir example_scoring
cp Geometry_exercise/geometry_solution.inp example_score/ex_Score.inp
cd example_scoring
```

- Open in FLAIR or with your preferred editor
- Add **USRBIN** scoring to:
  - 1) score **Energy** on a **CYLINDRICAL GRID** ( $R, \Phi, Z$  covering the target and surroundings:  $0 < r < 10$  cm,  $-5 < z < 15$  cm, with cells having  $\Delta r = \Delta z = 1$  mm,  $\Delta \Phi = 2\pi$ , Output unit = 40 BIN
  - 2) score **Neutron Fluence** on the same grid, Output unit = 41 BIN
  - 3) score **Charged Hadron Fluence** on the same grid, Output unit = 42 BIN
- Run 5 cycles, 1000 primaries each



- 1) Add USRBIN to score **ENERGY** on a **CYLINDRICAL GRID** ( $R, \Phi, Z$ ) covering the target and the surroundings:  $0 < r < 10$  cm,  $-5 < z < 15$  cm, with cells having  $\Delta r = \Delta z = 1$  mm,  $\Delta\Phi = 2\pi$

\* Energy deposition [GeV/cm<sup>3</sup>]

\* ..+....1....+....2....+....3....+....4....+....5....+....6....+....7...

USRBIN	11.	ENERGY	-40.	10.0	15.0	TargEne
USRBIN	0.0		-5.0	100.	200.	&

 **USRBIN**

Type: R- $\Phi$ -Z ▼

Part: ENERGY ▼

Rmin: 0.0

X:

Zmin: -5.0

Unit: 40 BIN ▼

Rmax: 10.0

Y:

Zmax: 15.0

Name: TargEne


NR: 100.

N $\Phi$ :

NZ: 200.


- This is an R- $\Phi$ -Z binning (WHAT(1) = 11.), scoring energy density (generalized particle **ENERGY**, or FLUKA PID 208), writing unformatted output to unit 40, spanning  $0 < R < 10$  cm in 100 bins,  $0 < \Phi < 2\pi$  in 1 bin (default),  $-5 < Z < 15$  cm in 200 bins.

2) Add USRBIN to score **Neutron Fluence** on the same grid as before:

```
* Neutron fluence [1/cm^2]
* ..+....1....+....2....+....3....+....4....+....5....+....6....+....7...
USRBIN          11.  NEUTRON      -41.      10.0      15.0TargNeu
USRBIN          0.0              -5.0      100.      200. &
 USRBIN
Type: R-Φ-Z ▼      Rmin: 0.0      Unit: 41 BIN ▼      Name: TargNeu
Part: NEUTRON ▼      X:              Rmax: 10.0      NR: 100.
                        Y:              NZ: 200.
                        Zmin: -5.0      Zmax: 15.0
```

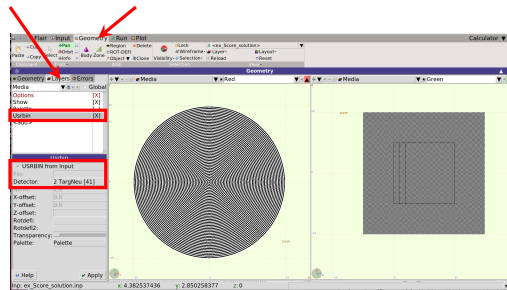
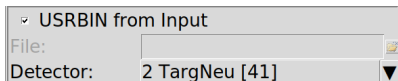
- This is an R- $\Phi$ -Z binning (WHAT(1) = 11.), scoring neutron fluence (particle NEUTRON, or FLUKA PID 8), writing unformatted output to unit 41, spanning  $0 < R < 10$  cm in 100 bins,  $0 < \Phi < 2\pi$  in 1 bin (default),  $-5 < Z < 15$  cm in 200 bins.

3) Add USRBIN to score **Charged Hadron Fluence** (HAD-CHAR) on the same grid as before, writing to output unit = 42 BIN.

- You can use the  Clone button in FLAIR to clone the previous scorer, then modify Unit: and Part: accordingly
- Or simply the copy the two lines in your favorite text editor...

You can display the USBIN-scorer you just defined in FLAIR's "Geometry" menu:

- In the "Geometry"-menu of FLAIR, chose the "Layers"-tab and select the desired layer (e.g. "media")
  - Click "<add>", then mark "USBIN"
- Usrbn [X]
- Mark "USBIN from Input" and select the desired scorer from the drop-down menu

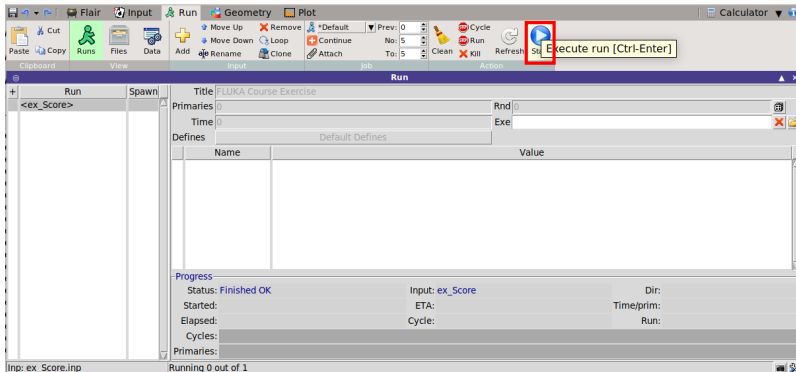


After you ran the example (next slide), deselect "USBIN from Input", and instead select the \*.bnn-file and plot the results inside the FLAIR Geometry-Menu.

# Exercise: USBIN



Run the input file by clicking the  button in the “Run”-menu of FLAIR:



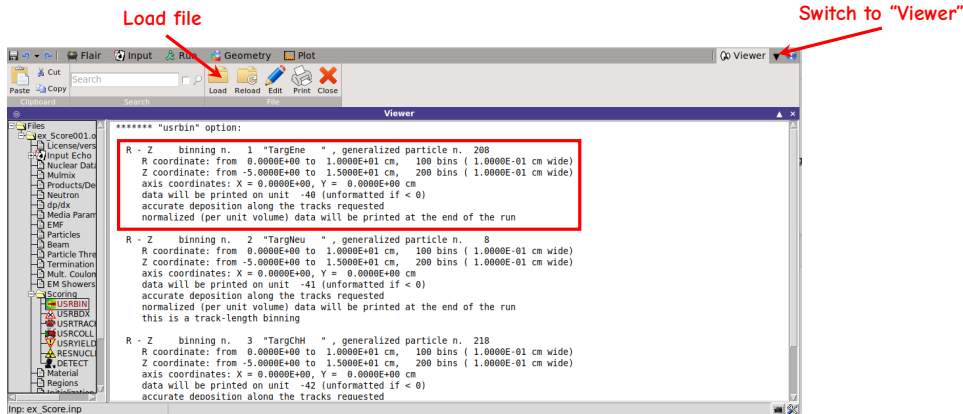
Or execute FLUKA on the commandline:

```
$FLUPRO/flutil/rfluka -e $FLUPRO/flukahp -N0 -M5 ./ex_Score
```

# Exercise: USRBIN



Load the .out-file in the "Viewer" of FLAIR - it gives a complete description of each requested estimator:




Alternatively, check the .out-files on the command line by doing

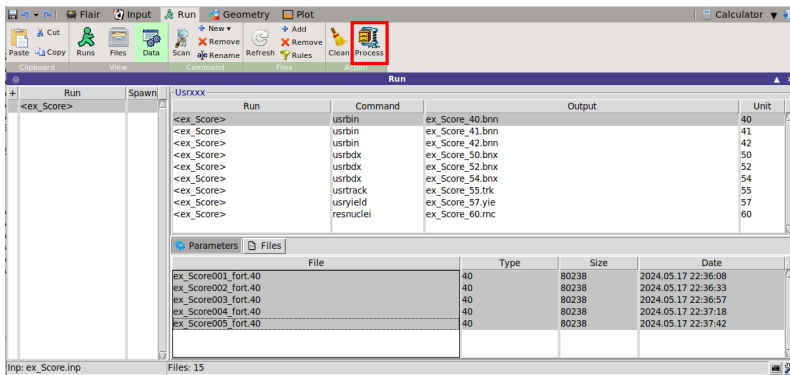
**less ex\_Score001.out**

# Exercise: USBIN - Processing of the data



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Select the desired scorer in the “Run” → “Data” tab, add the corresponding \*\_fort.xx-files and click the  button:



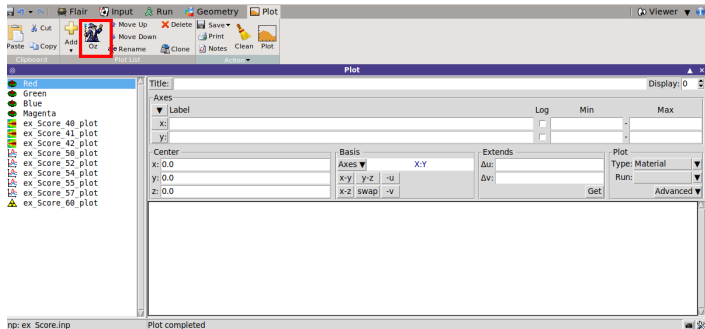
Or execute the post-processing script on the commandline:

**\$FLUPRO/flutil/usbsuw** (then follow the instructions)

# Exercise: USBIN - Plotting of the data



Plots can be created in the “Plot” menu - add new plots or clone from existing ones:



- The “Wizard” button scans the input and creates automatically a plot for each processed unit
- Set a unique filename for each plot
  - This filename will be used for auxiliary files that the plot needs

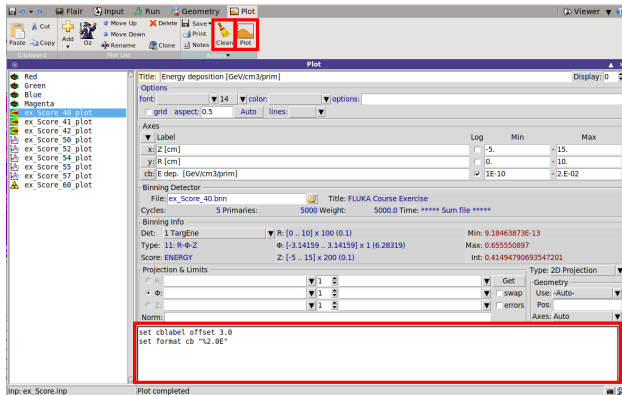
Plot types:



- |            |  |            |   |
|------------|--|------------|---|
| • Geometry | For geometry plots   | • USBR-2D  | To plot double differential distributions from <b>USRBDX</b>                        |
| • USBIN    | For plotting the output of USBIN   | • RESNUCLE | To plot 1d or 2d distributions from <b>RESNUCLEI</b>                                |
| • USBR-1D  | To plot single differential quantities from <b>USRBDX</b> , <b>USRTRACK</b> , <b>USRCOLL</b> , <b>USRYIELD</b> | • USERDUMP | To plot the output of <b>USERDUMP</b><br>Useful for visualizing source distribution |

# Exercise: USBIN - Plotting of the data



FLAIR uses **GNUPLLOT** to create the plots:

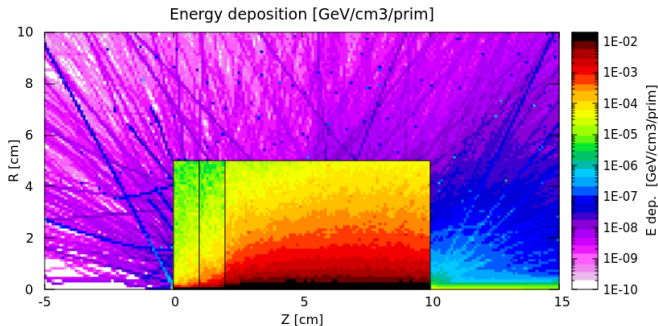


- For all plots one can specify: Title + options, filename, axis labels, legends, Gnuplot commands
- The  button (Ctrl-Enter) will generate all the necessary files (if they don't exist yet) and produce the plot
- The  button will remove all files generated by FLAIR during plotting (useful when the plot name was changed)

- Additional GNUPLLOT commands can be specified in the white field, e.g.:
  - Change colorband label offset
  - Change format of colorband (cb) palette values to “%2.0E”



**WHAT(2) = ENERGY:** Energy deposition from a 3.5 GeV proton beam hitting at [0., 0., 0.] directed along z. Results are normalized to **GeV/cm<sup>3</sup> per primary**

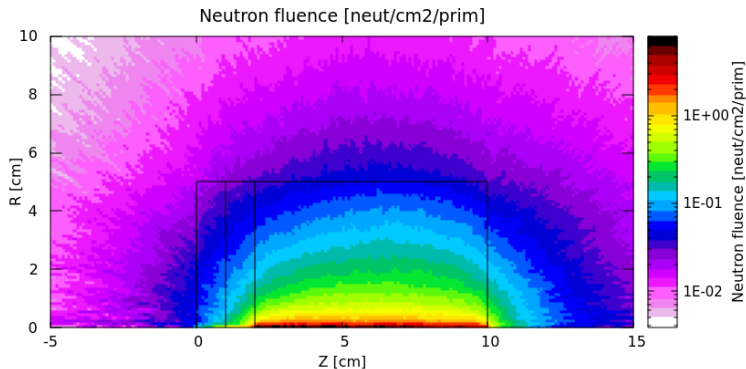


This plot is a 2D projection of a 3D structure → the result is averaged over the 3<sup>rd</sup> coordinate. Projection limits can be set in FLAIR.

## Exercise: USRBIN → The result



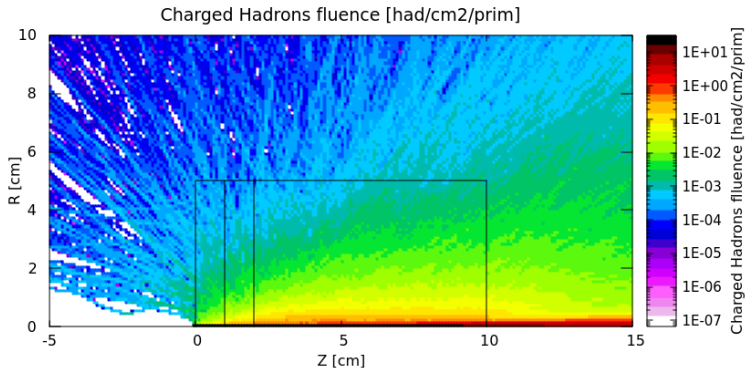
Same for **WHAT(2) = NEUTRON**: Neutron fluence from a 3.5 GeV proton beam hitting at [0., 0., 0.] directed along z. Results are normalized to **neutrons/cm<sup>2</sup> per primary**



## Exercise: USBIN → The result



Same for **WHAT(2) = HAD-CHAR**: Charged hadron fluence from a 3.5 GeV proton beam hitting at [0., 0., 0.] directed along z. Results are normalized to **charged hadrons/cm<sup>2</sup> per primary**





**USRBIN** can score particle fluence as well as “generalized particles”, either fluence-like or energy-like, for instance with **WHAT(2)** given as

- **DOSE** Energy/unit mass [GeV/g]
- **DPA-SCO** Displacements per atom
- **DOSE-EQ** Dose equivalent [pSv]  
(needs **AUXSCORE** card to set fluence-to-dose conversion coefficients)
- **X-MOMENT** x-component of momentum transfer [GeV/c]
- **ACTIVITY** activity per unit volume [Bq/cm<sup>3</sup>]  
(needs **DCYTIMES** and **DCYSCORE** cards to specify decay times)
- ...and more (see FLUKA manual)

**USRBDX** (“User Boundary Crossing” scoring) scores double differential (energy and angle) particle distributions across a boundary surface. The **angle** is with respect to the normal of the surface. The angular distributions must be intended as distributions in solid angle  $\Omega = 2\pi(1 - \cos \theta)$ , where  $\theta$  is the angle between the particle trajectory and the normal to the boundary at the point of crossing.

Input card:

**WHAT(1):** distribution can be fluence or current, one- or two-way, according to the value of WHAT(1) (see manual)  
**WHAT(2):** particle type to be scored  
**WHAT(3):** logical output unit  
**WHAT(4):** first region defining the boundary  
**WHAT(5):** second region defining the boundary  
**WHAT(6):** area of the detector in  $\text{cm}^2$   
**SDUM:** detector identifier

Continuation card:

**WHAT(1):** maximum kinetic energy for scoring [GeV]  
**WHAT(2):** minimum kinetic energy for scoring [GeV]  
**WHAT(3):** number of energy intervals  
**WHAT(4):** maximum solid angle in steradians  
**WHAT(5):** minimum solid angle (linear angular binning) or solid angle of first bin (logarithmic angular binning)  
**WHAT(6):** number of angular bins  
**SDUM:** & (continuation character)

- Score **CHARGED HADRONS** at the outer surface of the lead segment (from TARGS3 to INAIR). **WHAT(1)**=99 means: fluence scoring, one-way only, logarithmic intervals in energy from 1 MeV to 10 MeV in 40 intervals, and *one angular interval (default)*. **WHAT(6)** is a normalization factor: setting it equal to the surface area provides results normalized to  $\text{cm}^{-2}$  (unit of fluence)  $\text{GeV}^{-1} \text{sr}^{-1}$  per primary particle. Write the output to unformatted unit 50:

\* charged hadron fluence exiting lead target

USRBDX	99.	HAD-CHAR	-50.	TARGS3	INAIR	329.87	Sp3ChH
USRBDX	10.	0.001	40.				&

charged hadron fluence exiting lead target

⚠ USRBDX

Type:  $\Phi 1, \text{LogE}, \text{Lin}\Omega$  ▼

Part: HAD-CHAR ▼

Reg: TARGS3 ▼

Emin: 0.001

$\Omega$ min:

Unit: 50 BIN ▼

to Reg: INAIR ▼

Emax: 10.

$\Omega$ max:

Name: Sp3ChH

Area: 329.87

Ebins: 40.

$\Omega$ bins:

Repeat the same between **TARGS1** and **TARGS2**, and between **TARGS2** and **TARGS3** (take care to use the correct normalization factor!).

2) Score **CHARGED HADRONS** at the surface between 2<sup>nd</sup> and 3<sup>rd</sup> section, but in 3 angular bins:

\* double-differential charged hadron fluence entering lead target

USRBDX	99.	HAD-CHAR	-54.	TARGS2	TARGS3	78.5398	Sp2ChHA
USRBDX	10.	0.001	40.				3. &

double-differential charged hadron fluence entering lead target

**USRBDX**

Type:  $\Phi$ 1,LogE,Lin $\Omega$  ▼

Part: HAD-CHAR ▼

Reg: TARGS2 ▼

Emin: 0.001

Qmin:

Unit: 54 BIN ▼

to Reg: TARGS3 ▼

Emax: 10.

Qmax:

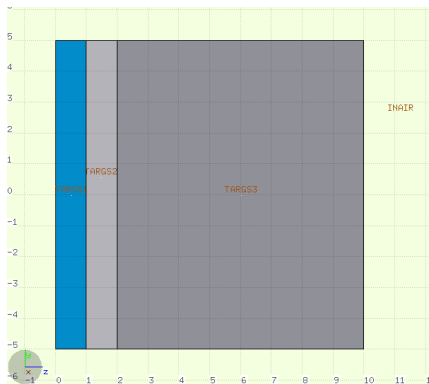
Name: Sp2ChHA

Area: 78.5398

Ebins: 40.

Qbins: 3.

# Exercise: USRBDX - Normalization



$$R_{\text{TARG}} = 5 \text{ cm}$$

$$Z_{\text{TARGS1}} = Z_{\text{TARGS2}} = 1 \text{ cm}$$

$$Z_{\text{TARGS1}} = 8 \text{ cm}$$


Area between TARGS2 and TARGS3:

$$\pi R_{\text{TARG}}^2 = 78.5398 \text{ cm}^2$$

Area between TARGS3 and INAIR:

$$\pi R_{\text{TARG}} Z_{\text{TARGS3}} + \pi R_{\text{TARG}}^2 = 329.87 \text{ cm}^2$$



The FLUKA postprocessing program for USBDX, `$FLUPRO/flutil/usxsuw` generates files with extension `_sum.lis` (human readable) and `_tab.lis` (slightly more machine readable). Invoke it by clicking the  button in FLAIR, or execute the post-processing script on the commandline: `$FLUPRO/flutil/usxsuw` (then follow the instructions)

Example of `_sum.lis` from USBDX:

```
Total primaries run:          5000
Total weight of the primaries run: 5000.00000
Detector n:          1 (      1 ) Sp1ChH
(Area:              78.5398026 cmq,
distr. scored:      218 , Charged Hadrons
from reg.          2 to      3 ,
one way scoring,
fluence scoring scoring)
```

User-provided area in WHAT(6),  
thus normalization in  $\text{cm}^{-2}$

```
Tot. resp. (Part/cmq/pr) 1.3232824E-02 +/- 0.3747608 %
( --> (Part/pr)        1.039303 +/- 0.3747608 % )
```

```
**** Different. Fluxes as a function of energy **** (integrated over solid angle) ****
(...list of data points...)
**** Cumulative Fluxes as a function of energy **** (integrated over solid angle) ****
(...list of data points...)
**** Double diff. Fluxes as a function of energy ****
(...list of data points...)
```

- Example of `_tab.lis` for USRBDX:

```
# Detector n:          1   Sp1ChH      (integrated over solid angle)
# N. of energy intervals          40

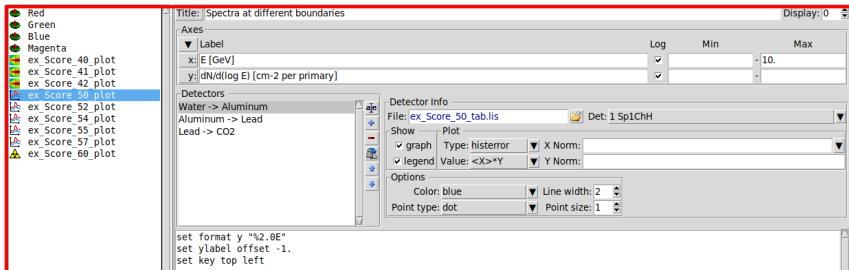
  E_min      E_max      Result      Error (%)
1.000E-02   1.259E-02   1.781E-03   9.900E+01
1.259E-02   1.585E-02   0.000E+00   0.000E+00
1.585E-02   1.995E-02   6.536E-04   9.900E+01
1.995E-02   2.512E-02   7.069E-03   7.818E+01
2.512E-02   3.162E-02   1.320E-03   6.309E+01

# double differential distributions
# number of solid angle intervals          1
# Block n:          1      0.00000000      :   6.28318548
  1.000E-01   1.259E-01   3.060E-04   2.570E+01
  1.259E-01   1.585E-01   3.144E-04   2.570E+01
  1.585E-01   1.995E-01   7.765E-05   2.369E+01
  1.995E-01   2.512E-01   1.959E-04   2.605E+01
  2.512E-01   3.162E-01   1.078E-04   3.540E+01
```

# Exercise: USRBDX - USR-1D Plotting



Plot → Add → USR-1D (or using the “Wizard”  button)



- USR-1D plots the 1D single differential information from USRBDX, USRCOLL, USRTRACK and USRYIELD cards
- File type should be \_tab.lis (generated by FLUKA data merging tool **usxsuw**)
- Many scoring outputs can be superimposed on the same plot
- Plot error bars using “histerror” or “xyerrorbars”
- Use a **lethargy** plot  to plot the spectrum

The result shows the evolution of charged hadron spectra at the different surfaces. post-processed results are normalized to  $\text{GeV}^{-1} \text{ cm}^{-2}$  per primary (only if surface area is explicitly given).

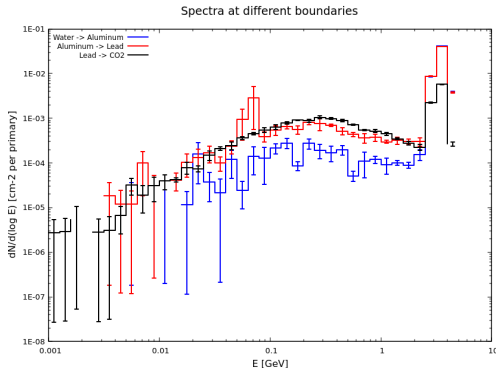
- **Lethargy plots** are used to display spectra where the energy extends over many orders of magnitude

$$y = \frac{dn}{d(\log E)} = E \frac{dN}{dE}$$

Value: <X>\*Y ▼

- In this way, the area of each bin is proportional to the corresponding integral flux, giving an immediate feeling which energy bin or region contributes more or less particles.

From post-processing, we get **single** differential spectra since we asked for one angular bin only

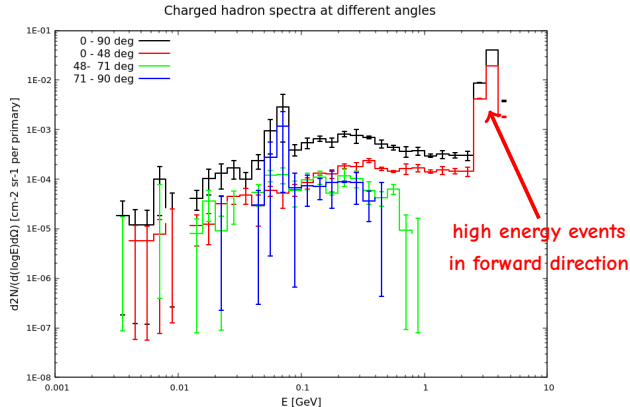


# Exercise: USRBDX - Results



Double-differential charged hadron spectra for 3 consecutive solid angle bins results are normalized to  **$\text{GeV}^{-1} \text{sr}^{-1} \text{cm}^{-2}$  per primary** (only if surface area is explicitly given)

From post-processing, we get **double** differential spectra,  
since we asked for more than one angular bin,  
but the angle-integrated spectrum is provided as well on top



USRTRK calculates **differential fluence** as a function of **energy** by scoring track-length in a given region. Results are normalized to  $\text{GeV}^{-1} \text{cm}^{-2}$  per primary particle if the region volume is provided (otherwise should be intended as  $\text{GeV}^{-1} \text{cm}$  per primary particle, i.e. differential track-length). For all **SDUMs** except EN-NUCL and ENERGY:

- WHAT(1): linear (=1) or logarithmic (= -1) energy binning
- WHAT(2): particle type to be scored
- WHAT(3): logical output unit
- WHAT(4): region defining the detector (if >0) or all regions (if = -1)
- WHAT(5): volume of the detector in  $\text{cm}^3$
- WHAT(6): number of energy bins
- SDUM: track-length detector identifier

Continuation card:

- WHAT(1): maximum kinetic energy for scoring [GeV]
- WHAT(2): minimum kinetic energy for scoring [GeV]
- SDUM: & (continuation character)

For **SDUM=EN-NUCL**: The energy scale will be changed from energy to energy per nucleon.

For **SDUM= ENERGY**: the energy scale will be changed to the default, that is total kinetic energy.

# Exercise: USRTRACK



- 1) Score the track-length of **CHARGED HADRONS** in TARGS3, with logarithmic binning in energy (**WHAT(1)=-1**) using 40 bins between 1 MeV and 10 GeV. Normalize with the region volume in order to have the results in  $\text{GeV}^{-1} \text{cm}^{-2}$  per primary particle. Write the output to unformatted unit 55:

\* charged hadron fluence in lead target

\* ..+....1....+....2....+....3....+....4....+....5....+....6....+....7...

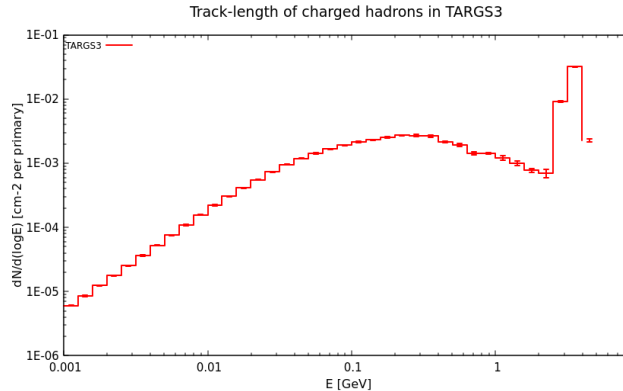
USRTRACK	-1.	HAD-CHAR	-55.	TARGS3	628.31	40.	TrChH
USRTRACK	10.	0.001					&

charged hadron fluence in lead target

USRTRACK			Unit: 55 BIN ▼	Name: TrChH
Type: LogE,Groupwise ▼	Reg: TARGS3 ▼			Vol: 628.31
Part: HAD-CHAR ▼	Emin: 0.001	Emax: 10.		Bins: 40.

**Remember:** USRTRACK scores differential fluence in a **region**, USRBDX scores fluence or current on a **surface**, and USRBIN scores e.g. fluence in **volumes** and gives no differential information.

Track-length based fluence of **charged hadrons** in region **TARGS3**, plotted as a lethargy plot:





- Scores a double differential particle yield across a boundary surface or at interaction points. The angles are defined respect to the **beam direction**. Only 1 interval is used in the second quantity (use several USRYIELD cards for more intervals in second quantity).
- “Energy-like” quantities:

Kinetic energy , total momentum , total energy , longitudinal momentum in the lab frame ,  
longitudinal momentum in the c.m.s. frame LET

- “Angle-like” quantities:

Rapidity in the lab frame , rapidity in the c.m.s. frame , pseudorapidity in the lab frame ,  
pseudorapidity in the c.m.s. frame , Feynman-x in the lab frame ,  
Feynman-x in the c.m.s. frame , transverse momentum , transverse mass ,  
polar angle (\*) in the lab frame , polar angle (\*) in the c.m.s. frame ,  
square transverse momentum , charge , weighted angle in the lab frame ,  
weighted transverse momentum

- While **USRBDX** scoring calculates angular distributions **with respect to the normal** to the boundary at the point of crossing, **USRYIELD**'s distribution are calculated **with respect to the beam direction**, as defined by the **BEAMPOS** card (use an additional **USRYIELD** card with **SDUM=BEAMDEF** to define a different direction).

## Input card:

**WHAT(1):** definition of two physical quantities  
**WHAT(2):** particle type to be scored  
**WHAT(3):** logical output unit  
**WHAT(4):** first region defining the boundary  
**WHAT(5):** second region defining the boundary  
**WHAT(6):** normalization factor  
**SDUM:** detector name

## Continuation card:

**WHAT(1):** upper limit for the first quantity  
**WHAT(2):** lower limit for the first quantity  
**WHAT(3):** number of scoring intervals (bins) for first quantity  
**WHAT(4):** upper limit for the second quantity  
**WHAT(5):** lower limit for the second quantity  
**WHAT(6):** kind of yield or cross section, and/or target material (see manual)  
**SDUM:** & (continuation character)

# Exercise: USRYIELD



- 1) Score plain double-differential yield (continuation card WHAT(6)=3) for pions, with the first quantity polar angle (degree) and second quantity kinetic energy (WHAT(1)=124), between TARGS3 and INAIR, between 0 and 180 degrees in 18 bins and between 0 and 10 GeV:

\* charged pion angular distribution exiting lead target

USRYIELD	124.	PIONS+-	-57.	TARGS3	INAIR	1.YieAng
USRYIELD	180.	0.0	18.	10.	0.0	3. &

charged pion angular distribution exiting lead target

▼ USRYIELD

ie: Polar  $\theta$  lab deg ▼

Norm: 1.

to Reg: INAIR ▼

Type: Yield ▼

ia: Ekin GeV ▼

Part: PIONS+- ▼

Min1: 0.0

Min2: 0.0

Unit: 57 BIN ▼

ip: Groupwise ▼

Yield: ▼

Max1: 180.

Max2: 10.

Name: YieAng

Log: Linear ▼

Reg: TARGS3 ▼

Nbins1: 18.

Kind: d2N/dx1dx2 ▼

**Remember:** Only one interval is possible for the second variable, but results are normalized as double-differential quantities (in this case, charged pions yield in  $\text{GeV}^{-1} \text{sr}^{-1}$  per primary).

Use WHAT(6) = 3 for plain double differential yield, DEFAULT is plain double-differential cross section!

The FLUKA postprocessing program for USRYIELD scorers, \$FLUPRO/flutil/usysuw, generates files with extension **\_sum.lis** and **\_tab.lis**.

Example of **\_sum.lis** from USRYIELD:

```
Detector n:      1  YieAng
(User norm:     1.00000000
sigma:         1.00000000      mb
distr. scored:   209 PIONS+-
from reg.       4  to reg.     5  )
linear 1st variable (x1) binning from 0.0000E+00 to 3.1416E+00 18 bins ( 1.7453E-01 wide)
2nd variable (x2) ranges from 0.0000E+00 to 1.0000E+01
1st variable (x1) is: Laboratory Angle (radians)
2nd variable (x2) is: Laboratory Kinetic Energy
The scored double differential yield is (normalized per primary particle):
  plain d2 N / d x1 d x2 where x1, x2 are
    the first and second variables

1st variable (x1) differential is a solid angle one

Tot. response (integrated over x1) 2.6879998E-02 +/- 2.035619 %
( --> Tot. response without normaliz. 2.6879998E-02 +/- 2.035619 % )
```

automatic conversion from degrees

**Warning! The total response is not integrated over the second quantity !!**

in this case it turns out to be *particles/GeV per primary*

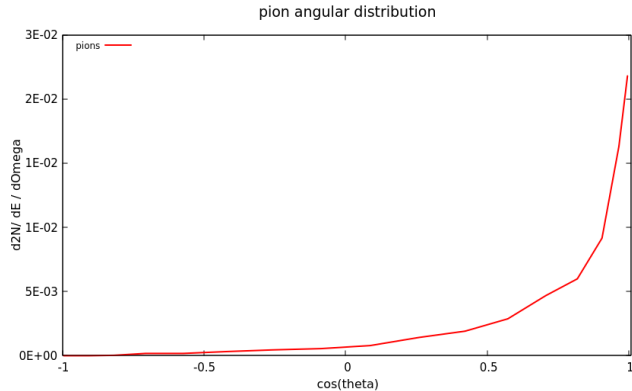
(to be multiplied by the energy width interval of 10 GeV)

Example of `_tab.lis` from USRYIELD:

```
# Detector n:          1   YieAng
# N. of x1 intervals    18

  Thetamin      Thetamax      Result      Error (%)
0.000000      0.1745329      2.1790212E-02      5.133210
0.1745329      0.3490658      1.6298208E-02      5.450077
0.3490658      0.5235988      9.1603426E-03      5.757809
0.5235988      0.6981317      5.9853652E-03      1.356122
0.6981317      0.8726646      4.6743122E-03      7.381429
0.8726646      1.047198      2.8757381E-03      3.952729
1.047198      1.221730      1.9141326E-03      12.11670
1.221730      1.396263      1.4178869E-03      8.692270
1.396263      1.570796      7.8822160E-04      18.23766
1.570796      1.745329      5.4992206E-04      20.41241
1.745329      1.919862      4.5372383E-04      23.19902
1.919862      2.094395      3.2238022E-04      18.22172
2.094395      2.268928      1.7834034E-04      54.48624
2.268928      2.443461      1.8077450E-04      28.57143
2.443461      2.617994      3.1837051E-05      99.00000
2.617994      2.792527      0.000000      0.000000
2.792527      2.967060      0.000000      0.000000
2.967060      3.141593      0.000000      0.000000
```

Pion angular distribution:



Use `gnuplot` commands to plot with FLAIR:

```
set key top left
set format y "%2.0E"
set ylabel offset -1.

plot 'ex_Score_57_tab.lis' ind 0 us (cos(($1+$2)/2.0)):3 w lines lw 2 lc rgb 'red' t 'pions'
```

Plot completed



- Results from **USRBDX**, **USRYIELD**, **USRTRACK**, **USRCOLL** are given as **differential** distributions of fluence in energy, in units of  $\text{cm}^{-2}\text{GeV}^{-1}$  per primary. Thus
  - 1) results are **independent** on the chosen binning
  - 2) to obtain **integral** results (fluence in  $\text{cm}^{-2}$  per primary for that energy interval) one needs to multiply each value by the width of the energy interval or bin (even for logarithmic binning):

$$N = \int \frac{dN}{dE} dE$$

- When scoring **neutrons** with groupwise treatment, energy bin limits below 20 MeV are automatically set to the neutron transport group limits (see lecture on low-energy neutrons), overwriting the user-defined binning
- Normalization is per  $\text{cm}^{-2}$  only if the user provides the surface area or region volume - FLUKA in general is not able to calculate areas or volumes by itself
- The same holds if USRBIN scoring is used with region binning
- Only for USRBIN scoring on grids is the normalization automatically in  $\text{cm}^{-2}$  (or  $\text{cm}^{-3}$  for deposited energy etc.)

There is the possibility to **filter** the estimators, restricting the scoring to a selected subset of particles.

For instance, to score the yield of 56-Iron ions (there is no separate name for each ion specie, except the light ones like 4-HELIUM, 3-HELIUM, TRITON, DEUTERON. HEAVYION scores all isotopes heavier than alpha-particles together)

```
*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...
USRYIELD      124.  ALL-PART      -58.  TARGS3      INAIR      1.Fe56
USRYIELD      180.      0.0      18.      10.      0.0      3. &
AUXSCORE      USRYIELD -5602600.      Fe56      Fe56
```

<b>USRYIELD</b> ie: Polar θ lab deg ▼ Norm: 1. to Reg: INAIR ▼	Type: Yield ▼ ia: Ekin GeV ▼ Part: ALL-PART ▼ Min1: 0.0 Min2: 0.0	Unit: 58 BIN ▼ ip: Groupwise ▼ Yield: ▼ Max1: 180. Max2: 10.	Name: Fe56 Log: Linear ▼ Reg: TARGS3 ▼ Nbins1: 18. Kind: d2N/dx1dx2 ▼
<b>AUXSCORE</b> Delta Ray: ▼	Type: USRYIELD ▼ Z: 26 Det: Fe56 ▼	Part: ▼ A: 56 to Det: Fe56 ▼	Set: ▼ Isomer: 0 Step:

The requested ion is coded in WHAT(2) according to its **A**, **Z** and (optionally) isomeric state **m**:

WHAT(2) = - (100\***Z** + 100 000\***A** + 100 000 000 \* **m**)

A=m=0 means all isotopes, e.g. -2600 = all Iron isotopes



- To be used with care, or better NOT used at all for energy deposition
- **In real world, energy is eventually deposited by electrons only!**
- In Monte Carlo calculations, part of the energy is deposited “by other particles” as continuous energy deposition or point energy deposition, depending on delta-ray threshold, production threshold, transport threshold (see lecture on Transport)
- **Any filtering done with AUXSCORE on energy deposition will depend on the adopted thresholds and settings!**



- The Monte Carlo simulation does **NOT** know the intensity of your beam. It only knows how many event histories were run
- Normalization in all FLUKA intrinsic scoring is “per primary event”, or better, “per unit primary weight” if the source particles are biased
- The normalization to experimental conditions has to be done by the user
- Exception : **Activation scoring**, because there the beam intensity is provided by the user in the **IRRPROFI** card (see lecture on activation).
- For scoring which involves volumes (**USRTRACK**) or areas (**USRBDX**), quantities are normalized per unit volume ( $1/\text{cm}^3$ ) or unit area ( $1/\text{cm}^2$ ) **only for regular meshes (e.g. USRBIN)** - for all other scorers, correct area or volume values have to be given by the user in the scoring card.
- Normalization can also be done at the plotting stage using the **Norm:**  field in FLAIR's **PLOT** menu.



- Can be calculated for **single histories** (not in FLUKA), or for **batches** (cycles or runs) of several histories
- Distribution of scoring contributions by **single histories** can be very asymmetric (many event histories contribute little or zero)
- Scoring distribution **from batches** tends to Gaussian for  $N \rightarrow \infty$ , provided  $\sigma^2 \neq \infty$  (thanks to **Central Limit Theorem!**)
- The standard deviation of an estimator calculated from batches or from single histories is **an estimate of the standard deviation of the actual distribution** ("error of the mean")
- The goodness of such an estimate depends on the type of estimator and on the particular problem (but it converges to the true value for  $N \rightarrow \infty$ )

- The **variance of the mean** of an estimated quantity  $x$  (e.g. *fluence*), calculated in  $N$  batches, is:

$$\sigma_{<x>}^2 = \frac{1}{N-1} \left[ \underbrace{\frac{\sum_1^N n_i x_i^2}{n} - \left( \frac{\sum_1^N n_i x_i}{n} \right)^2}_{\text{mean of squares - square of means}} \right]$$

$N - 1$

where:

$n_i$  = number of histories in the  $N$  batches

$n = \sum n_i$  = total number of histories in the  $N$  batches

$x_i$  = average of  $x$  in the  $i^{\text{th}}$  batch:  $x_i = \sum_{j=1}^{n_i} \frac{x_{ij}}{n_i}$


$x_{ij}$  is the contribution to  $x$  of the  $j^{\text{th}}$  history in the  $i^{\text{th}}$  batch

In the limit  $N = n$ ,  $n_i = 1$ , the formula applies to single history statistics

The FLUKA user routines combining several scoring files will use **this** formula!

## Practical tips:

- Use always at least **5-10 batches** of comparable size (it is not at all mandatory that they be of equal size)
- Never forget that **the variance itself is a stochastic variable** subject to fluctuations
- Plot 2D and 3D distributions! Looking at them the eye is the best tool in judging the quality of the result
- Use FLAIR's option to plot errors in the PLOT menu
- As a rule of thumb, the statistical uncertainty of your estimators in the regions you are interested in should be better than 5% (use variance reduction methods, see lecture on biasing)

- **Post-processing utilities** are provided in the FLUKA distribution. They
  - Sum the results from different cycles
  - Calculate statistical uncertainties
  - Provide a summed file, that contains averages and statistics, and can be re-used to sum other cycles
  - Provide human-readable as well as gnuplot-readable output
- These utilities are **used by flair** in the "Data process"-button 
- Can be (also) used directly from command line
- All of them are in the **\$FLUPRO/flutil/-**directory
- Compiled automatically during installation of FLUKA

- `flutil/detsuw.f` - DETECT: sum and provide `tab_lis`-file
- `flutil/gplevbin.f` - USRBIN: prepare 2D or 1D for plot
- `flutil/usbsuw.f` - USRBIN: sum and `.bnn` binary file
- `flutil/usbrea.f` - USRBIN: convert binary output to ASCII
- `flutil/usrsuwev.f` - RESNUC: offline evolution
- `flutil/usrsuw.f` - RESNUC: sum and `tab_lis`-files
- `flutil/ustsuw.f` - USRTRACK: sum and `tab_lis`-files
- `flutil/usxrea.f` - USRBDX: read summed file
- `flutil/usxsuw.f` - USRBDX: sum and provide `tab_lis`-files
- `flutil/usysuw.f` - USRYIELD: sum and `tab_lis`-files

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