



兰州大学
LANZHOU UNIVERSITY

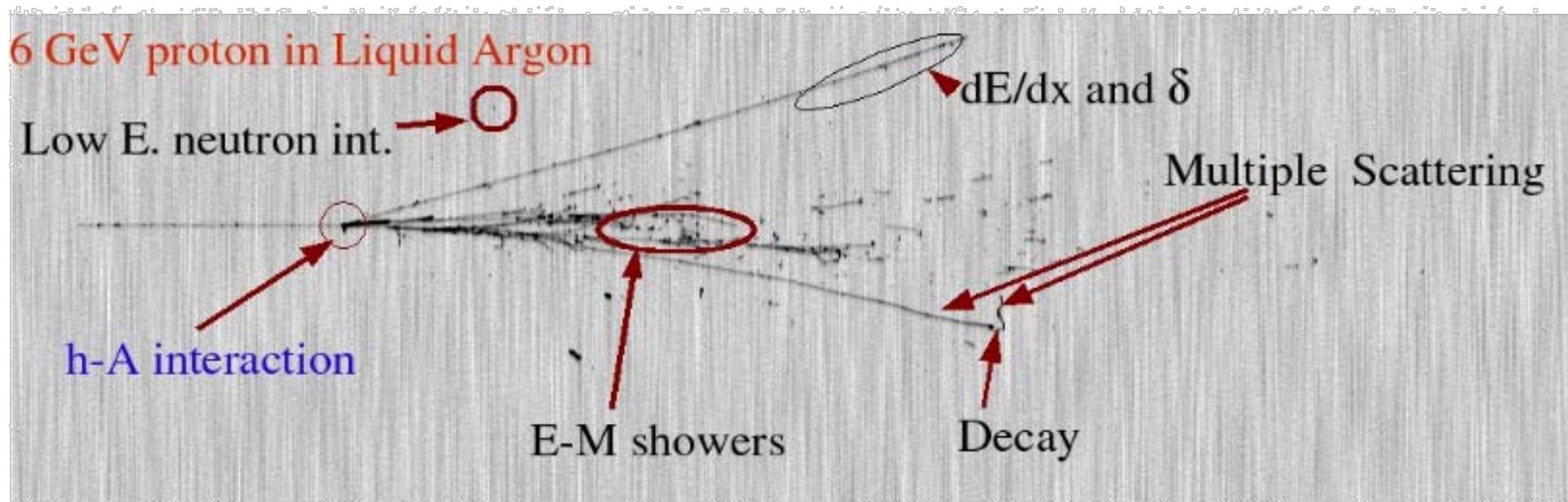
An Introduction to Fluka: a Multipurpose Particle Interaction and Transport MC code

23rd FLUKA Beginner's Course
Lanzhou University
Lanzhou, China
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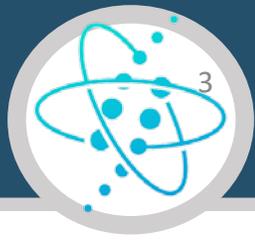


Developed and maintained by the FLUKA Collaboration

>10000 registered users

<http://www.fluka.org>

The FLUKA International Collaboration



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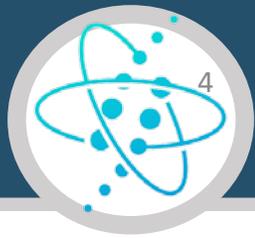
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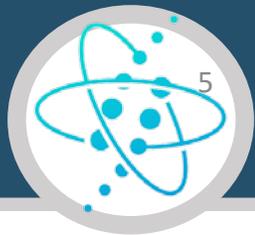


- FLUKA^{1,2} is a general purpose tool for calculations of particle **transport** and **interactions** with matter
- All **Hadrons** (p, n, π , K, pbar, nbar, (anti)hyperons...) [0-100 EeV, 10²⁰ eV]
- **Electromagnetic** (γ , e^{+/-}) and **μ and ν** [100 eV/1 keV – 100 EeV]
- **Nucleus-nucleus** [0-100 EeV/n]
- **Low energy neutrons** (0-20 MeV, multigroup, pointwise, ENDF...)
- Transport in **electric** and **magnetic field**
- Combinatorial (boolean) and Voxel **geometries**
- Double capability to run either fully **analogue and/or biased** calculations
- On-line **evolution** of induced **radioactivity** and **dose**
- **Radiation damage** predictions (NIEL, DPA)
- User-friendly **GUI** interface thanks to the **Flair³** interface

¹ F. Ballarini et al, "FLUKA: status and perspectives"
NEA (2024), SATIF 15 - Sessions 2 & 3: Code Status, Advances, & Model Converters, OECD Publishing, Paris , page 89
²A. Ferrari et al., "FLUKA: a multi-particle transport code",
CERN 2005-10 (2005), INFN/TC_05/11, SLAC-R-773

<http://www.fluka.org>

³V. Vlachoudis, Proc. Int. Conf. on Mathematics, Computational Methods & Reactor Physics (M&C 2009), Saratoga Springs, New York, 2009



Particles transported by FLUKA:

FLUKA name	FLUKA number	Symbol	Common name	Standard PDG number (Particle Data Group) [142]
4-HELIUM ⁽¹⁾	-6	α	Alpha	—
3-HELIUM ⁽¹⁾	-5	${}^3\text{He}$	Helium 3	—
TRITON ⁽¹⁾	-4	${}^3\text{H}$	Triton	—
DEUTERON ⁽¹⁾	-3	${}^2\text{H}$	Deuteron	—
HEAVYION ⁽¹⁾	-2	—	Generic Heavy Ion with $Z > 2$ (see command HI-PROPE)	—
OPTIPHOT	-1	—	Optical Photon	—
RAY ⁽²⁾	0	—	Pseudoparticle	—
PROTON	1	p	Proton	2212
APROTON	2	\bar{p}	Antiproton	-2212
ELECTRON	3	e^-	Electron	11
POSITRON	4	e^+	Positron	-11
NEUTRIE	5	ν_e	Electron Neutrino	12
ANEUTRIE	6	$\bar{\nu}_e$	Electron Antineutrino	-12
PHOTON	7	γ	Photon	22
NEUTRON	8	n	Neutron	2112
ANEUTRON	9	\bar{n}	Antineutron	-2112
MUON+	10	μ^+	Positive Muon	-13
MUON-	11	μ^-	Negative Muon	13
KAONLONG	12	K_L^0	Kaon-zero long	130
PION+	13	π^+	Positive Pion	211
PION-	14	π^-	Negative Pion	-211
KAON+	15	K^+	Positive Kaon	321
KAON-	16	K^-	Negative Kaon	-321
LAMBDA	17	Λ	Lambda	3122
ALAMBDA	18	$\bar{\Lambda}$	Antilambda	-3122
KAONSHRT	19	K_S^0	Kaon-zero short	310
SIGMA-	20	Σ^-	Negative Sigma	3112
SIGMA+	21	Σ^+	Positive Sigma	3222
SIGMAZER	22	Σ^0	Sigma-zero	3212
PIZERO	23	π^0	Pion-zero	111
KAONZERO	24	K^0	Kaon-zero	311
AKAONZER	25	\bar{K}^0	Antikaon-zero	-311
Reserved	26	—	—	—
NEUTRIM	27	ν_μ	Muon Neutrino	14
ANEUTRIM	28	$\bar{\nu}_\mu$	Muon Antineutrino	-14
Blank	29	—	—	—

table continues

FLUKA name	FLUKA number	Symbol	Common name	Standard PDG (Particle Data
Reserved	30	—	—	—
ASIGMA-	31	$\bar{\Sigma}^-$	Antisigma-minus	-3222
ASIGMAZE	32	$\bar{\Sigma}^0$	Antisigma-zero	-3212
ASIGMA+	33	$\bar{\Sigma}^+$	Antisigma-plus	-3112
XSIZEZERO	34	Ξ^0	Xi-zero	3322
AXSIZEZERO	35	$\bar{\Xi}^0$	Antixi-zero	-3322
XSI-	36	Ξ^-	Negative Xi	3312
AXSI+	37	$\bar{\Xi}^+$	Positive Xi	-3312
OMEGA-	38	Ω^-	Omega-minus	3334
AOMEGA+	39	Ω^+	Antionega	-3334
Reserved	40	—	—	—
TAU+	41	τ^+	Positive Tau	-15
TAU-	42	τ^-	Negative Tau	15
NEUTRIT	43	ν_τ	Tau Neutrino	16
ANEUTRIT	44	$\bar{\nu}_\tau$	Tau Antineutrino	-16
D+	45	D^+	D-plus	411
D-	46	D^-	D-minus	-411
DO	47	D^0	D-zero	421
DOBAR	48	\bar{D}^0	AntiD-zero	-421
DS+	49	D_s^+	D_s -plus	431
DS-	50	D_s^-	D_s -minus	-431
LAMBDAc+	51	Λ_c^+	Lambda _c -plus	4122
XSIc+	52	Ξ_c^+	Xi _c -plus	4232
XSIc0	53	Ξ_c^0	Xi _c -zero	4132
XSIc+	54	$\Xi_c^{'+}$	Xi' _c -plus	4322
XSIc0	55	$\Xi_c^{'0}$	Xi' _c -zero	4312
OMEGAc0	56	Ω_c^0	Omega _c -zero	4332
ALAMBDC-	57	$\bar{\Lambda}_c^-$	Antilambda _c -minus	-4122
AXSIc-	58	$\bar{\Xi}_c^-$	AntiXi _c -minus	-4232
AXSIc0	59	$\bar{\Xi}_c^0$	AntiXi _c -zero	-4132
AXSIc-	60	$\Xi_c^{\prime-}$	AntiXi' _c -minus	-4322
AXSIc0	61	$\Xi_c^{\prime0}$	AntiXi' _c -zero	-4312
AOMEGAc0	62	$\bar{\Omega}_c^0$	AntiOmega _c -zero	-4332
Reserved	63	—	—	—
Reserved	64	—	—	—



- Based, as far as possible, on **original** and well-tested **microscopic models**
- **Full cross-talk** between all components:
 - hadronic,
 - electromagnetic,
 - neutrons,
 - muons,
 - heavy ions
- It is a “condensed history” MC code, however with the possibility to use single instead of multiple Coulomb scattering
 - FLUKA is NOT a toolkit! Its physical models are fully integrated**
 - **The user does not need to choose a “physics list”**
 - **The user has, however, the possibility to optimize CPU vs accuracy**
- Fluka provides powerful built-in **scoring**, tested and suited for most applications
 - **The user does not need to write external code to get results and statistics**

What can be done with FLUKA?



FLUKA

Cosmic ray physics

Hadron therapy

Space radiation

Accelerator design

Neutrino physics

Detector simulation

Shielding design

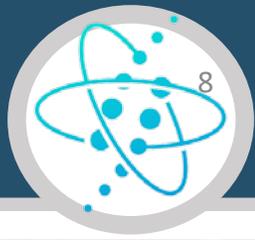
ADS systems, waste transmutation

Dosimetry and radioprotection

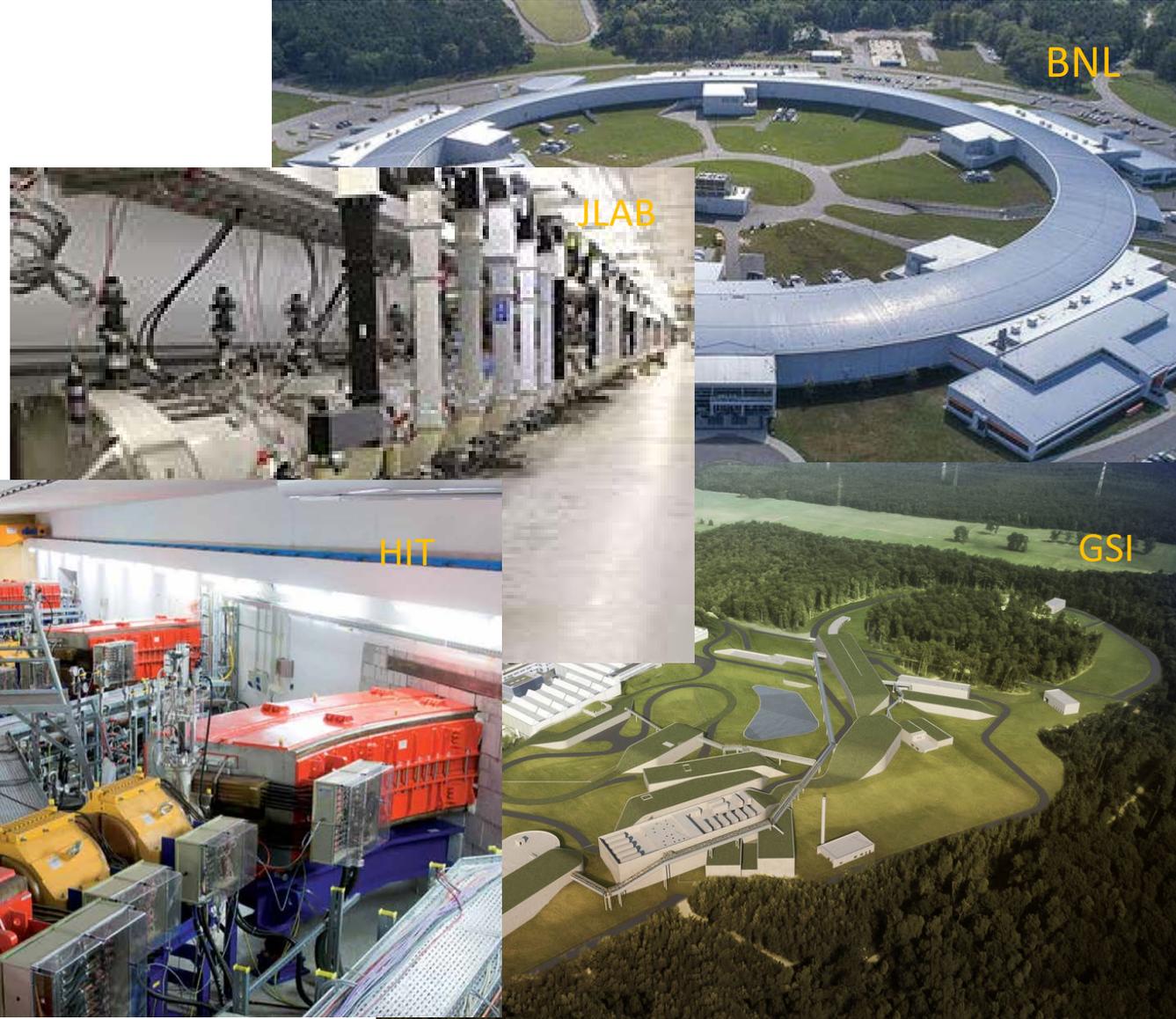
Radiation damage

Neutronics

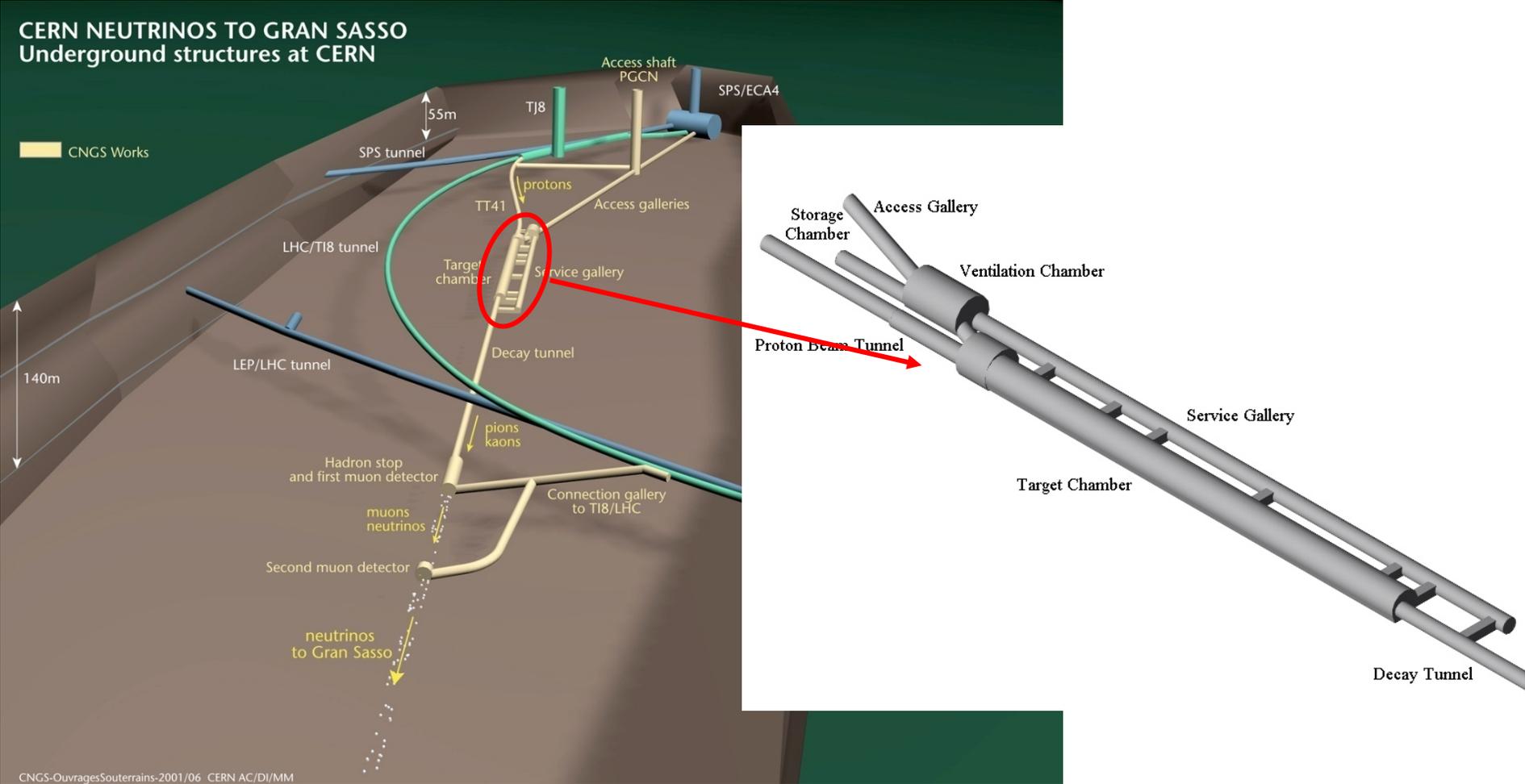
Accelerators:



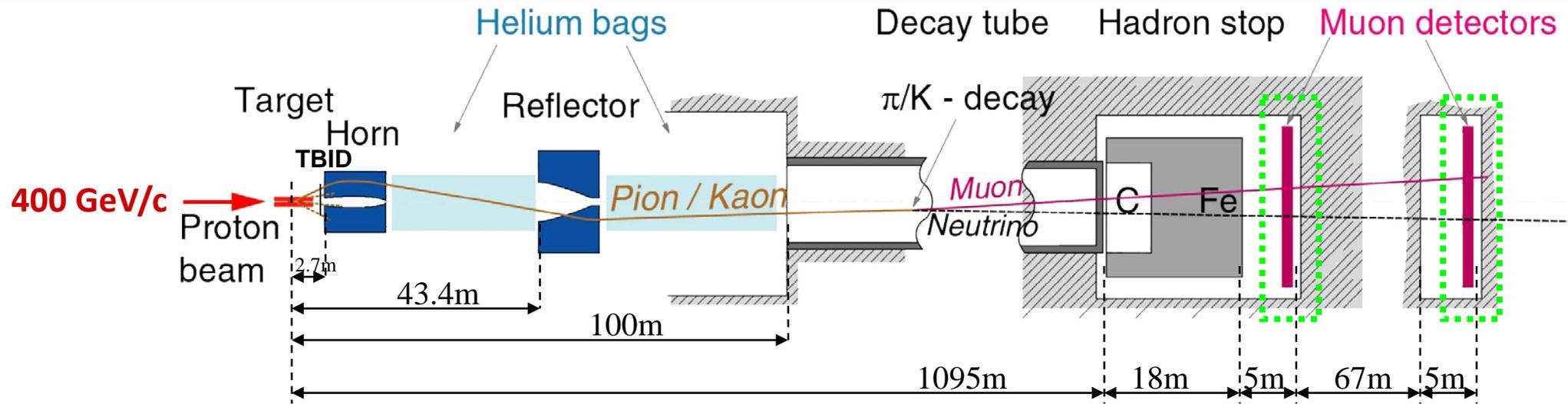
- ❑ Neutrino experiments
- ❑ Detector simulation
- ❑ Shielding, residual dose rates
- ❑ Energy deposition (quenching and damage)
- ❑ Radiation damage (electronics, insulation)
- ❑ Activation, waste disposal
- ❑ Shielding design
- ❑ Spallation sources
- ❑ Secondary beams



Accelerator applications - CNGS neutrino beam



CNGS: ν beam and muon monitors



Flight path to Gran Sasso : 732 km.

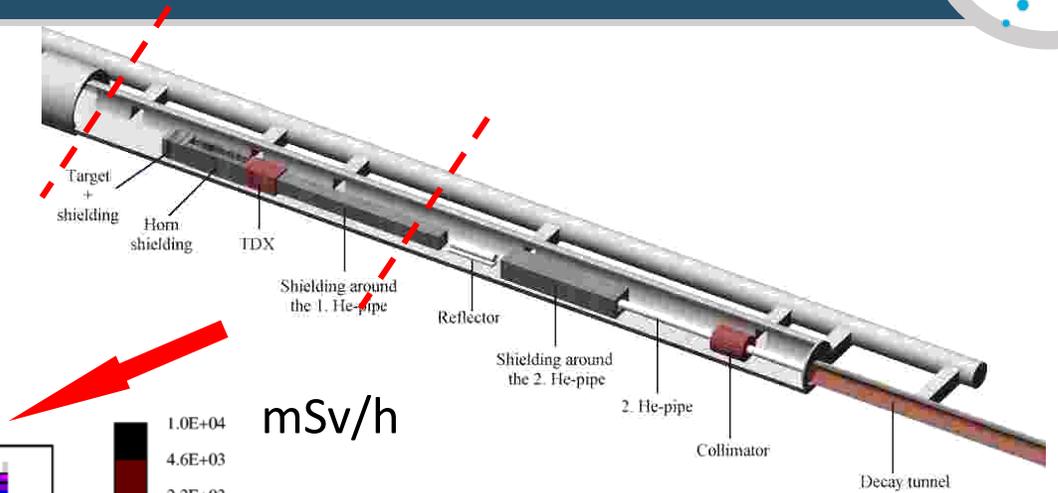
CNGS neutrino beam line designed and optimized with **FLUKA**

Muon monitors: a check of neutrino production

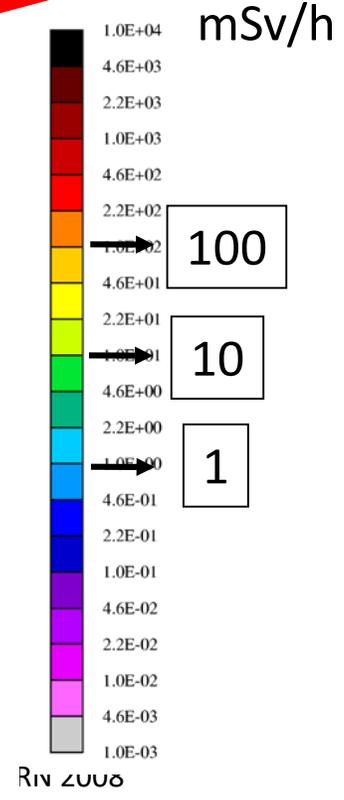
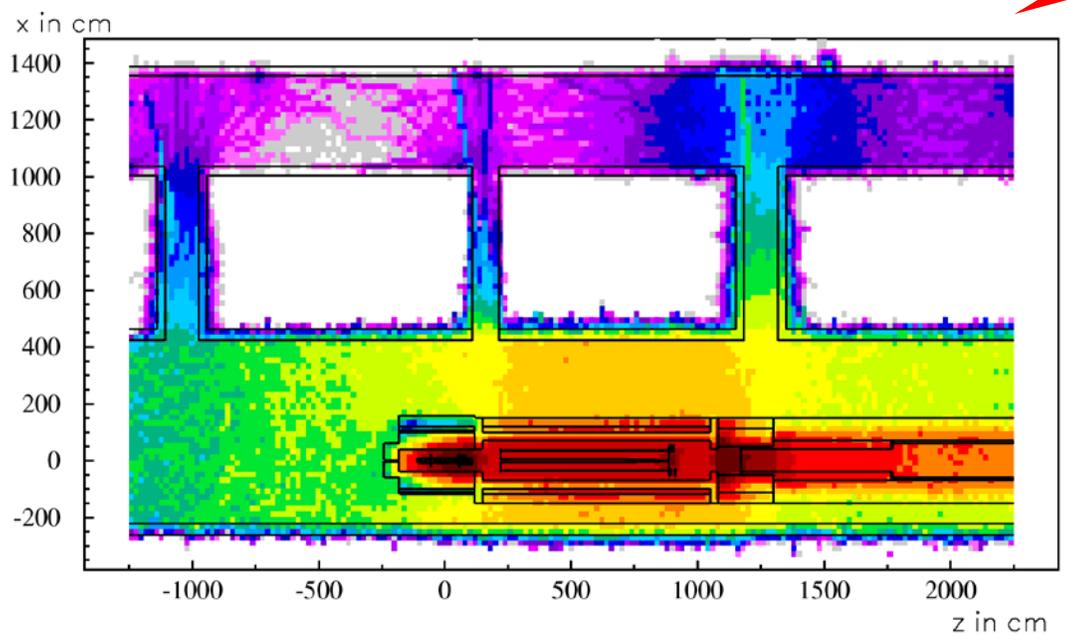
Applications - CNGS

Residual Dose Equivalent Rate (mSv/h):

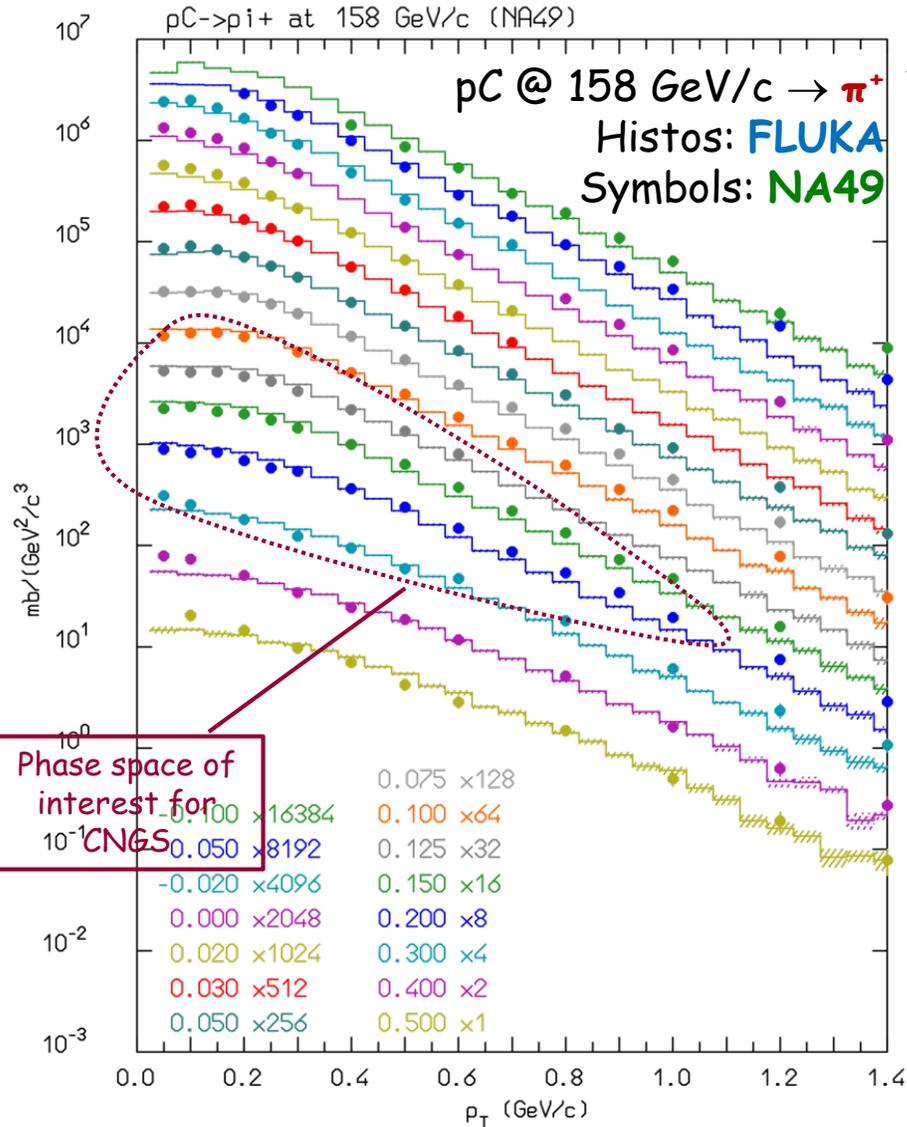
- 200 day long Irradiation;
- 1 day cooling time
- 8×10^{12} 400 GeV/c protons/s



Example:
 $t_{cool} = 1$ day



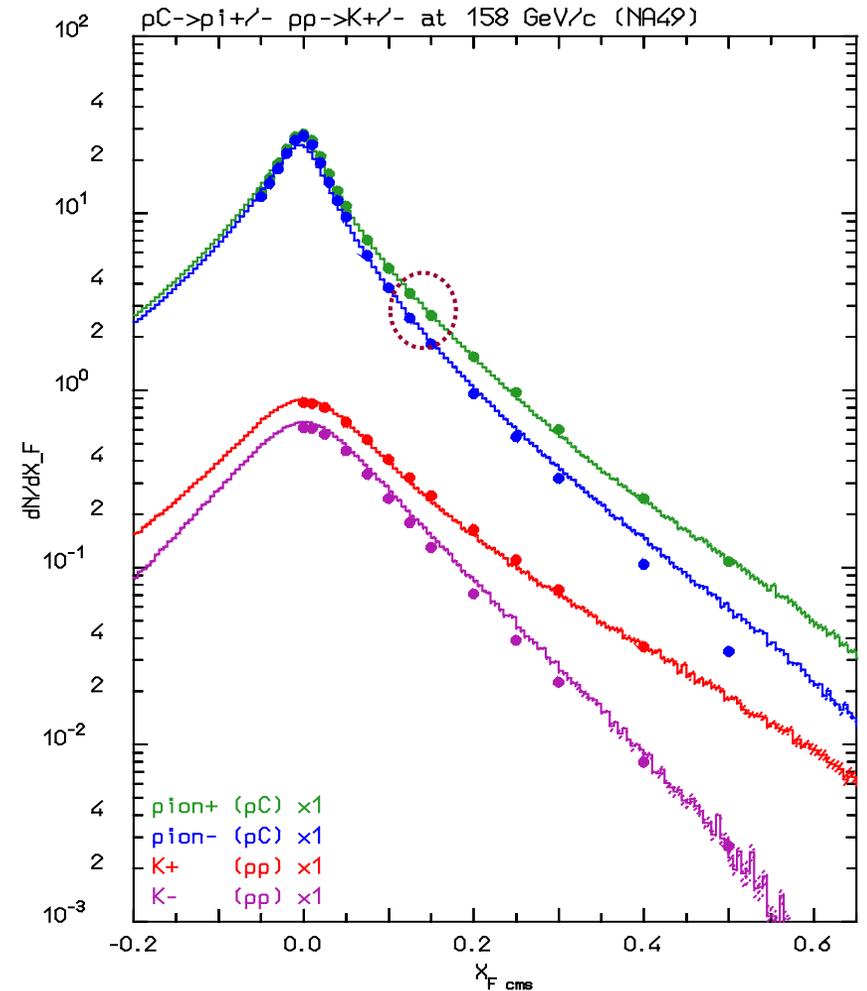
Pion and Kaon production data (ν beams...)



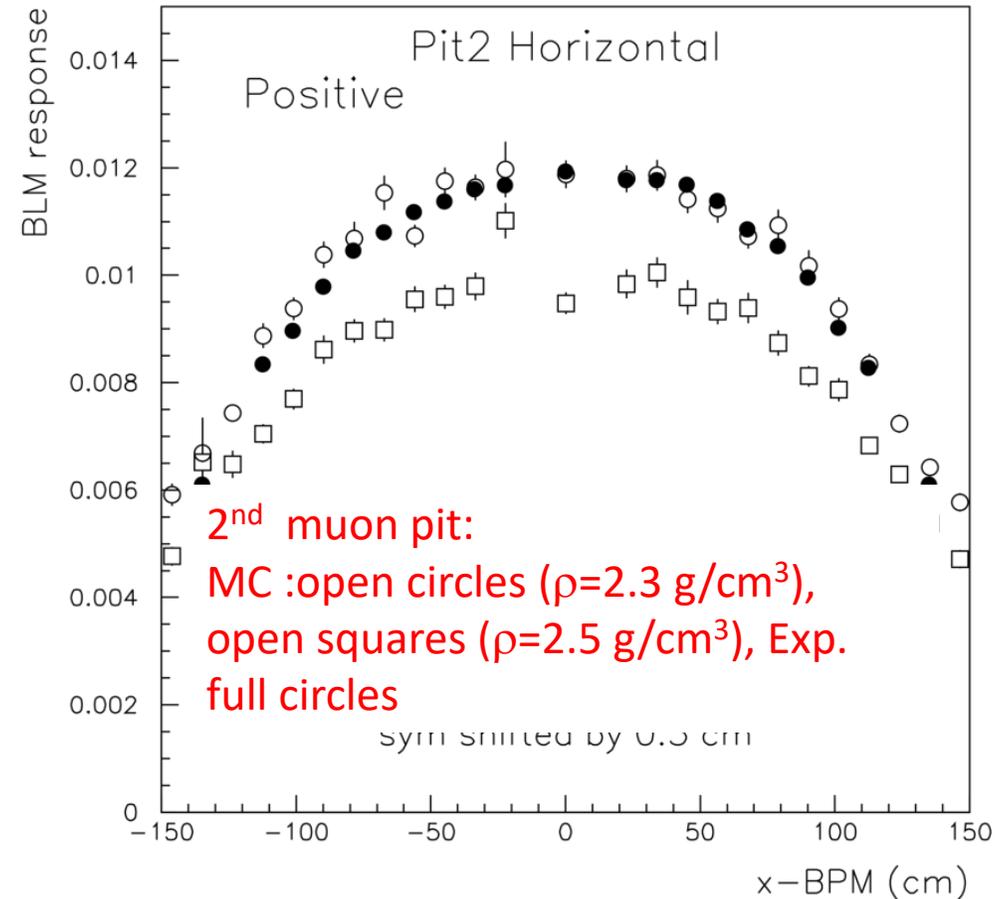
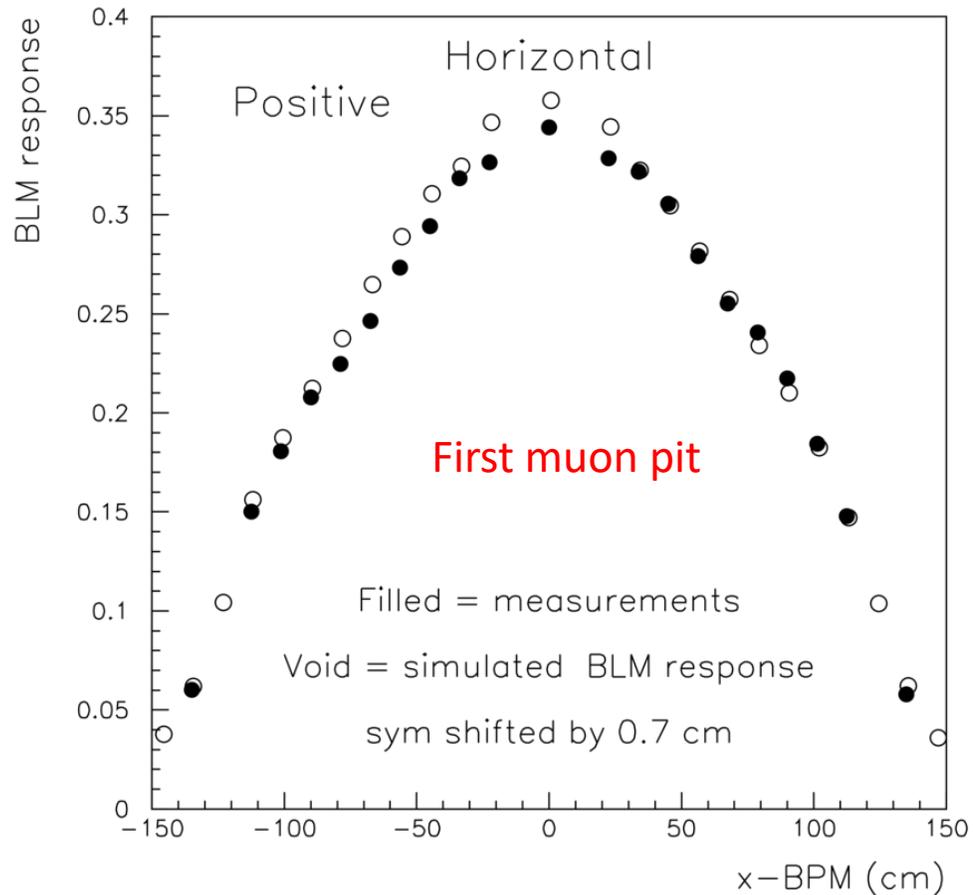
π^+ (left) yield as a function of p_T for different X_F bins for 158 GeV/c p on C

Angle integrated distributions are the most relevant for judging the reliability of ν predictions, at least for the bulk of the spectrum. \longrightarrow

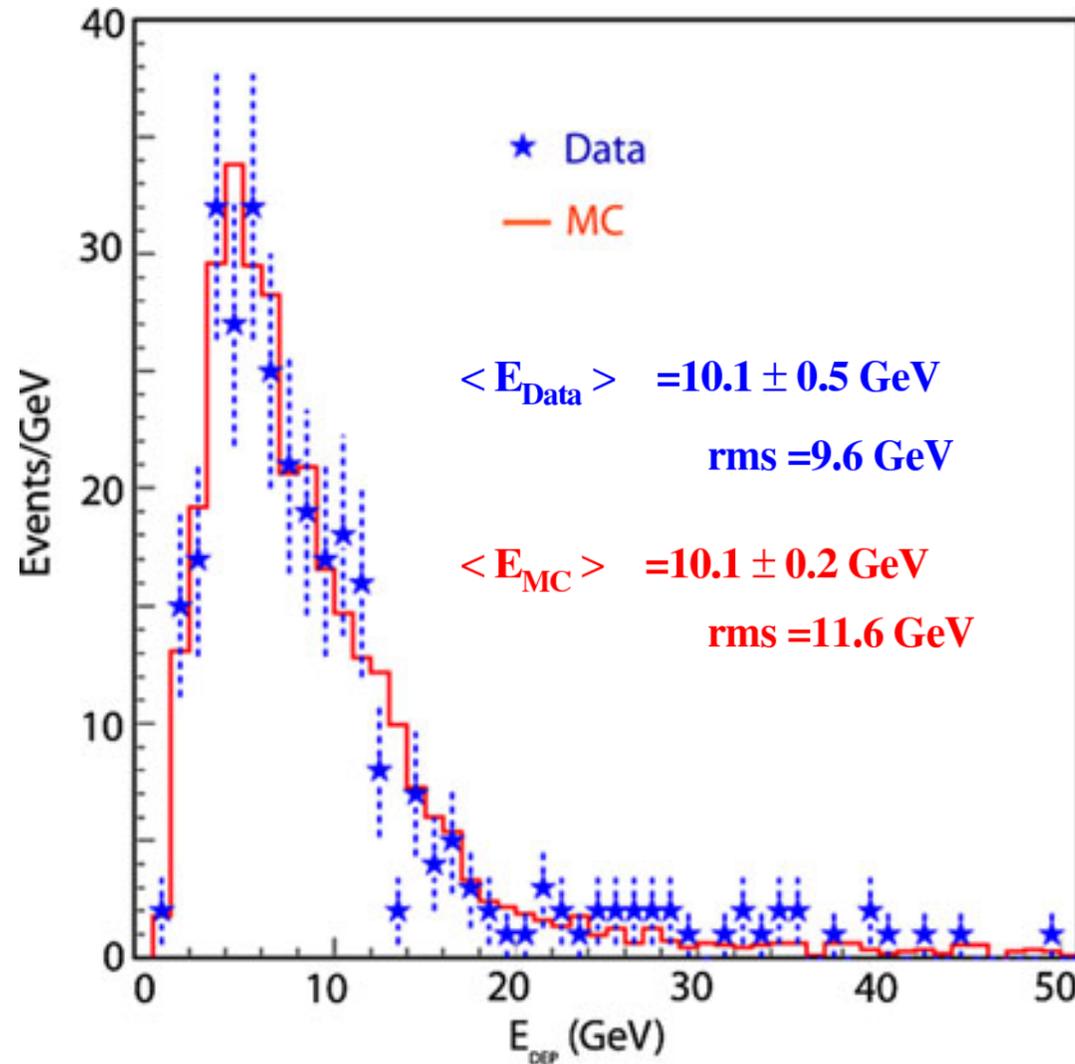
The p_T integrated distributions for $pC \rightarrow \pi^+$, π^- and $pp \rightarrow K^+$, K^- are shown in the right Figure as a function of Feynman X_F (dots exp. data, **NA49**, lines **FLUKA** predictions), together with the "focused" zone for **CNGS**



CNGS Muon pits (~110 m underground): data vs MC



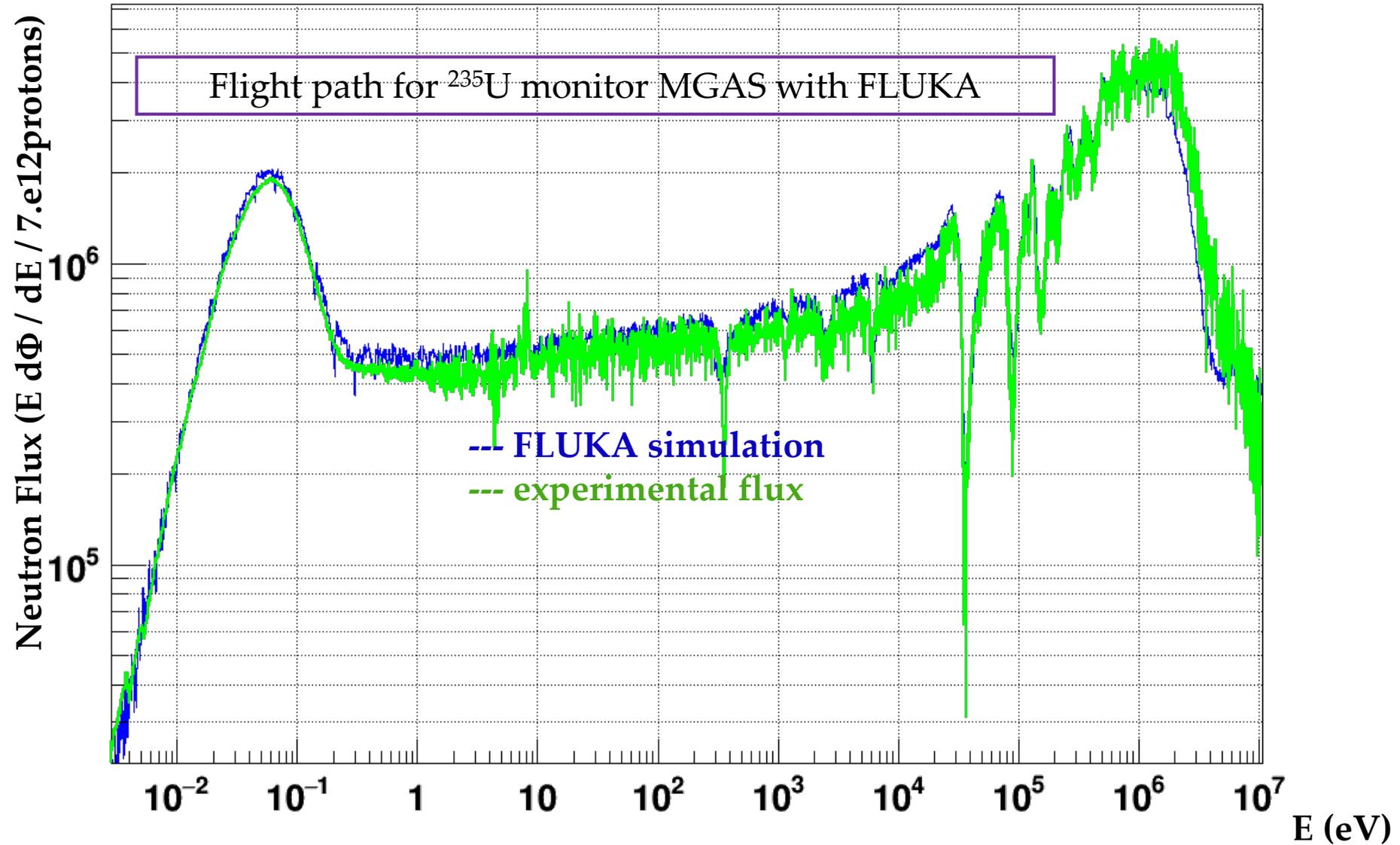
Absolute comparison! Included in MC: effect of **earth magnetic field** (in the 1 km long decay tunnel). Experimental uncertainties: detector calibration, density of the rock in between the two pits (67 m)



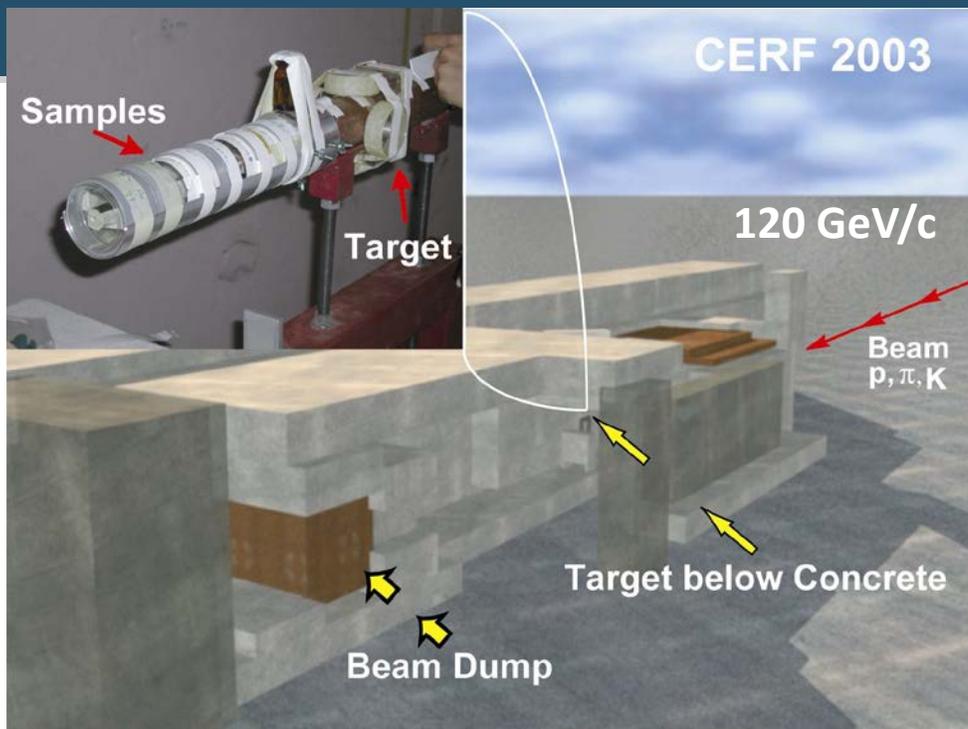
← Distribution of total deposited energy in the ICARUS T600 detector

- CNGS numuCC events ($\sim 20 \text{ GeV } E_\nu$ peak)
- Same reconstruction in MC (FLUKA) and Data
- Neutrino fluxes from FLUKA CNGS simulations
- Absolute agreement on neutrino rate within 6%

Eur. Phys. J. C (2013) 73:2345
Phys. Lett. B (2014)



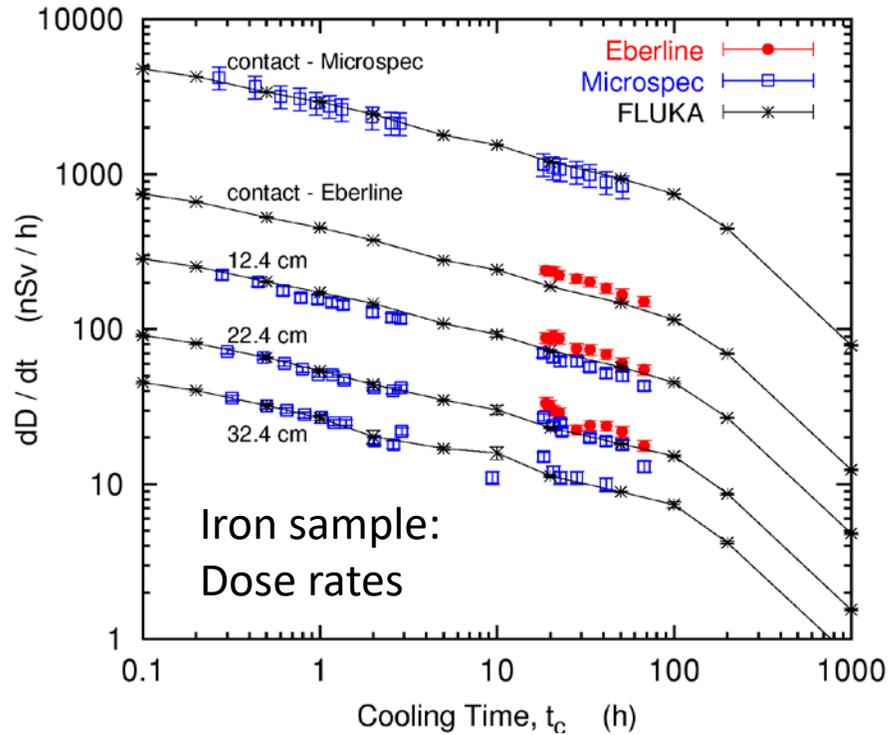
CERN-EU High-Energy Reference Field facility (CERF)



Thermo-Eberline dose-meter FHZ 672



Iron



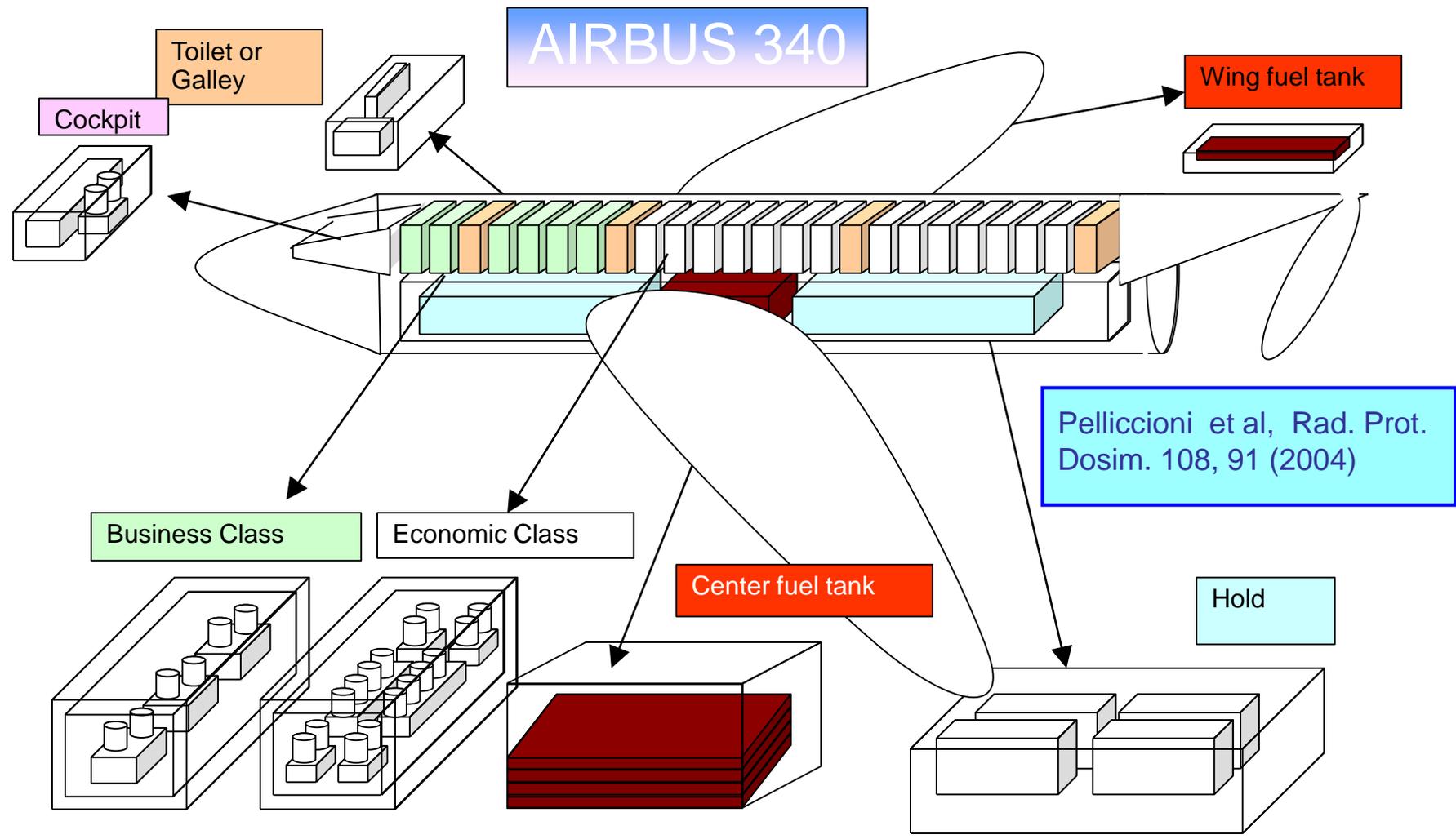
ACTIVATION of various samples in contact with a 50 cm long, 7 cm diameter copper target, centred on the beam axis, and irradiated with a 120 GeV/c beam



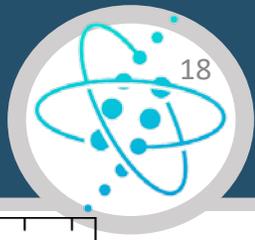
Microspec



Dosimetry applications: doses to aircrew and passengers



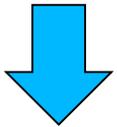
Commercial flight doses: (Pelliccioni et al. RPD93)



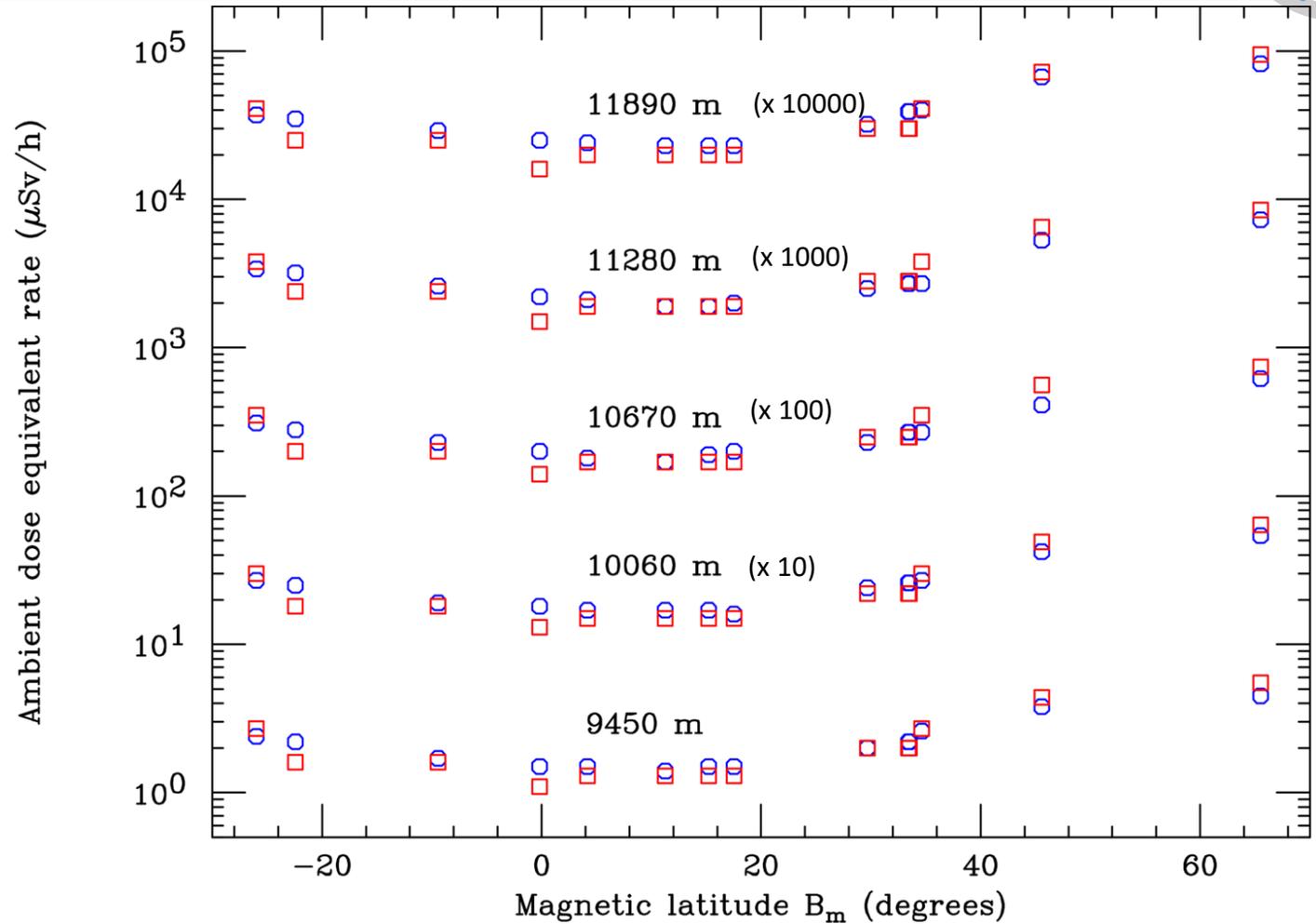
Complete FLUKA simulation of **cosmic rays** interactions in the atmosphere

- Dedicated “cosmic” package available to users
- Ready to use GCR spectra and geomagnetic cut-offs

Model of airplane geometry
Response of dosimeters



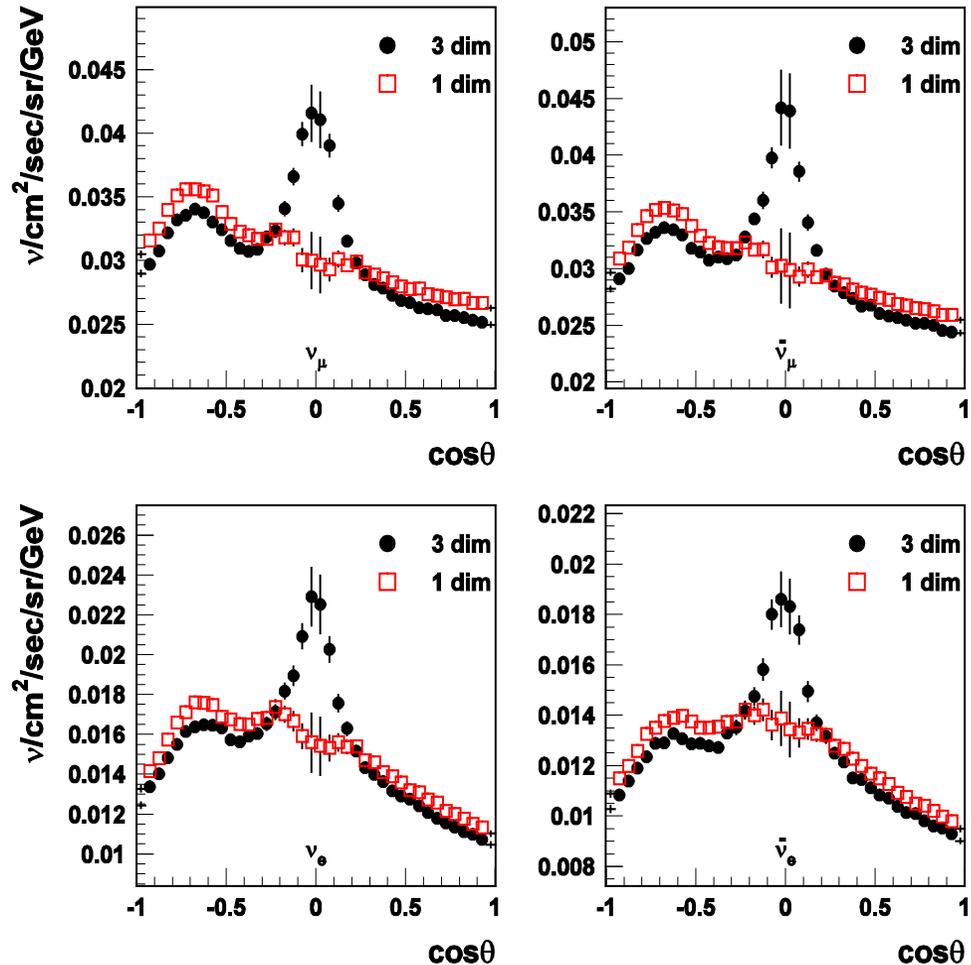
Dose to aircrew on commercial flights, depending on route



Simulated (**FLUKA, red**) and measured (**blue**, NIMA422, 621, 1999) ambient dose equivalent for various altitudes (scaled by one decade) and geomagnetic cut-off's

(3D) Calculation of Atmospheric $\bar{\nu}$ Flux

Sub-GeV flux at Kamioka



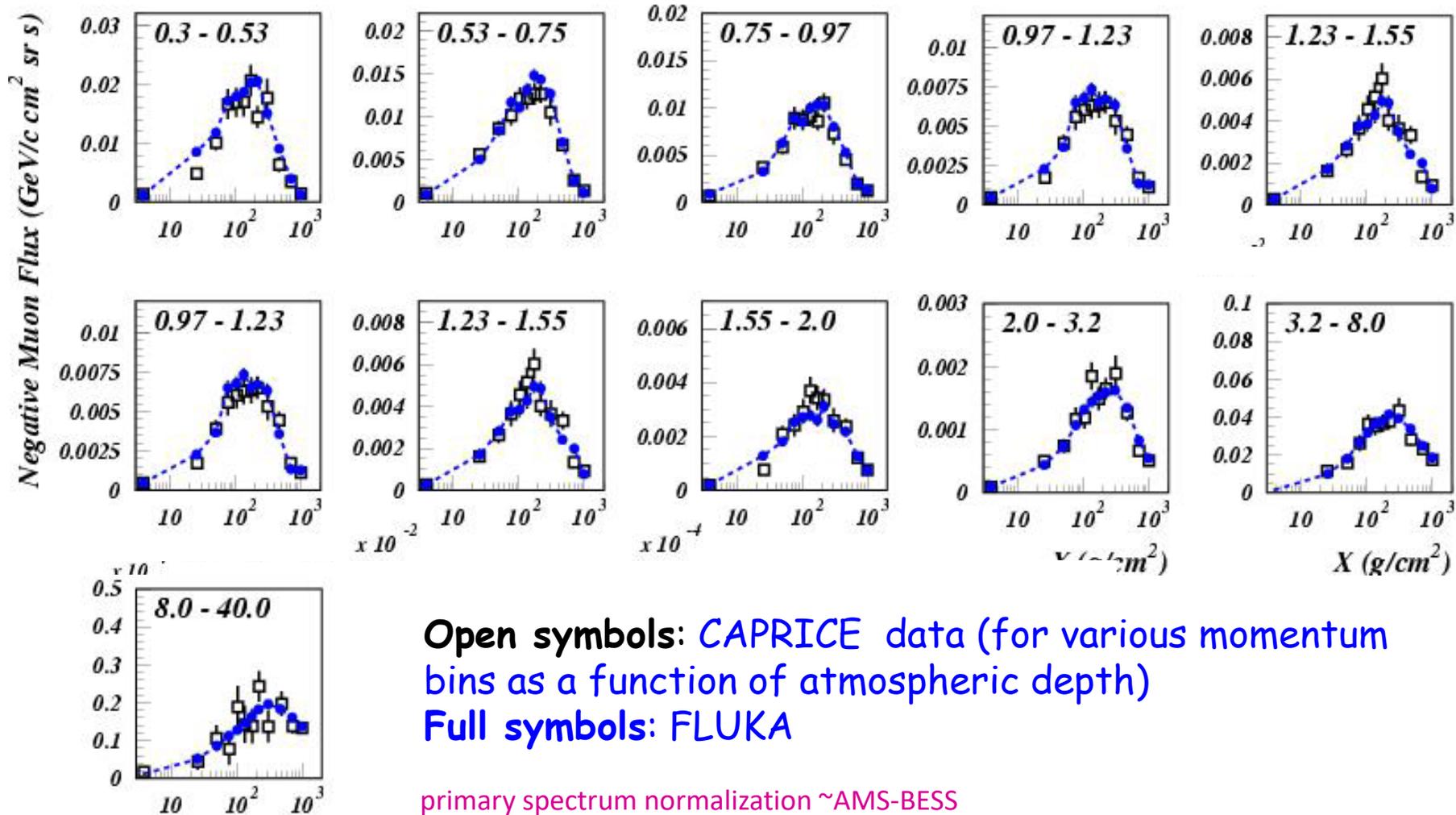
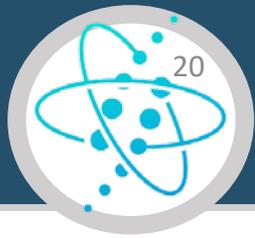
The first 3-D calculation of atmospheric neutrinos was done with FLUKA.

The enhancement in the horizontal direction, which cannot be predicted by a 1-D calculation, was fully unexpected, but is now generally acknowledged.

In the figure: angular distribution of ν_μ , $\bar{\nu}_\mu$, ν_e , $\bar{\nu}_e$

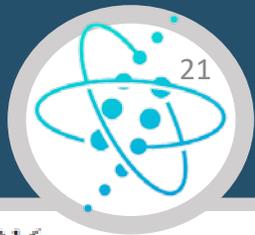
In red: 1-D calculation

COSMIC RAYS: Negative muons at floating altitudes: CAPRICE94



Open symbols: CAPRICE data (for various momentum bins as a function of atmospheric depth)
Full symbols: FLUKA

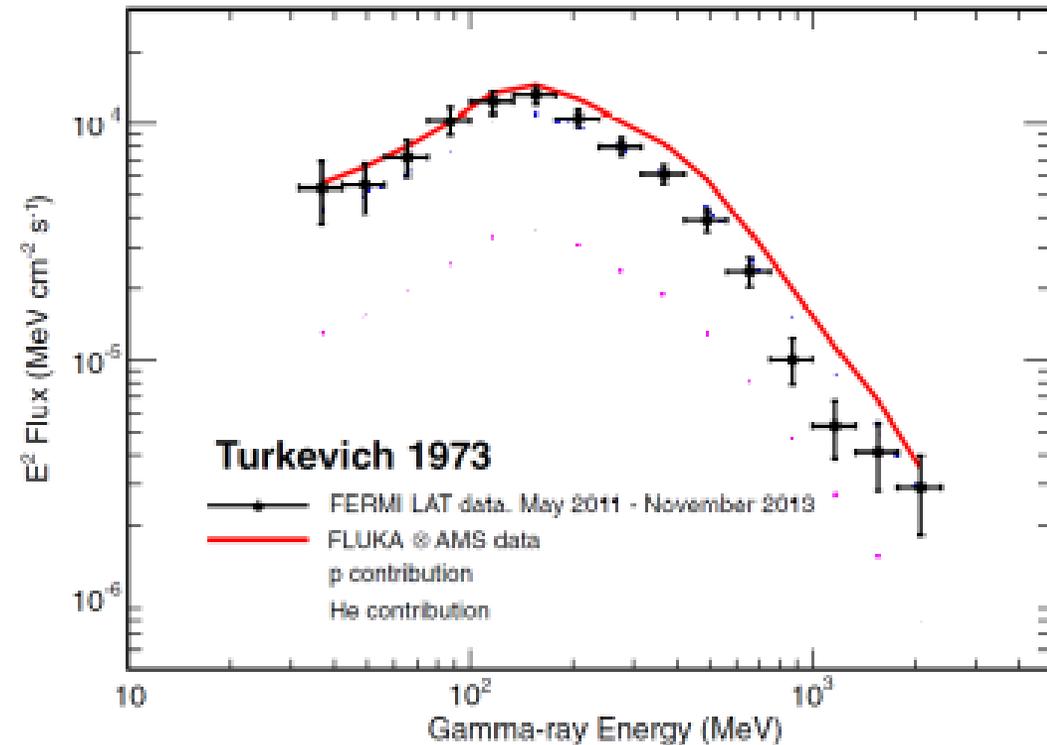
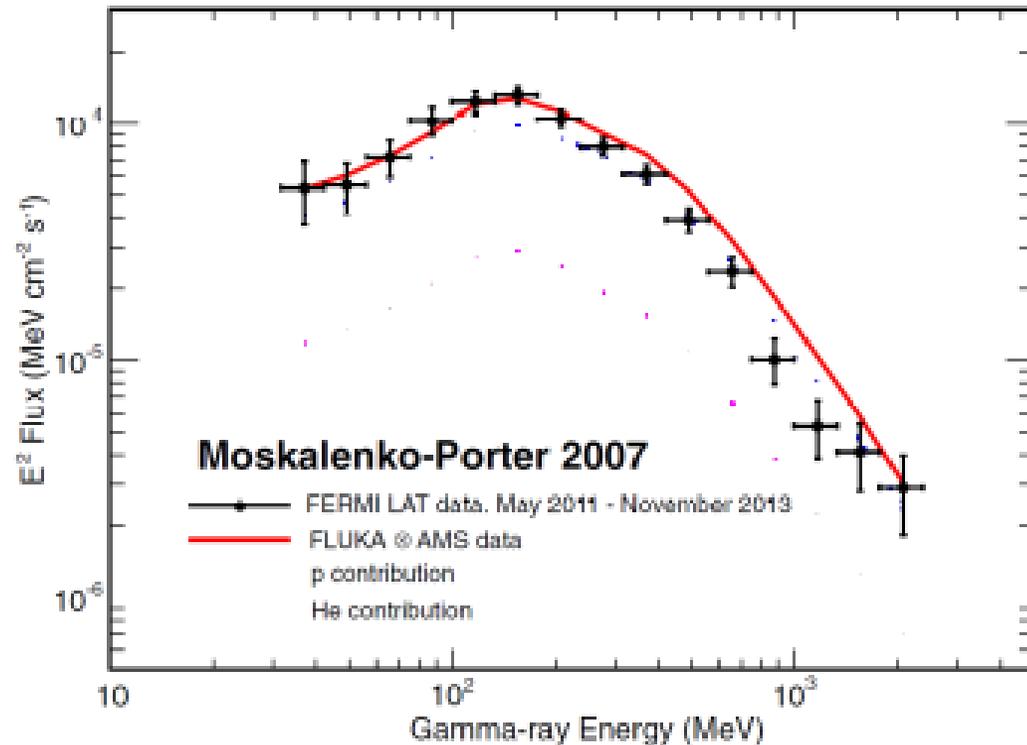
primary spectrum normalization ~AMS-BESS
Astropart. Phys., Vol. 17, No. 4 (2002) p. 477



Gamma rays from GCR interactions with the moon:

M. ACKERMANN *et al.*

PHYSICAL REVIEW D **93**, 082001 (2016)

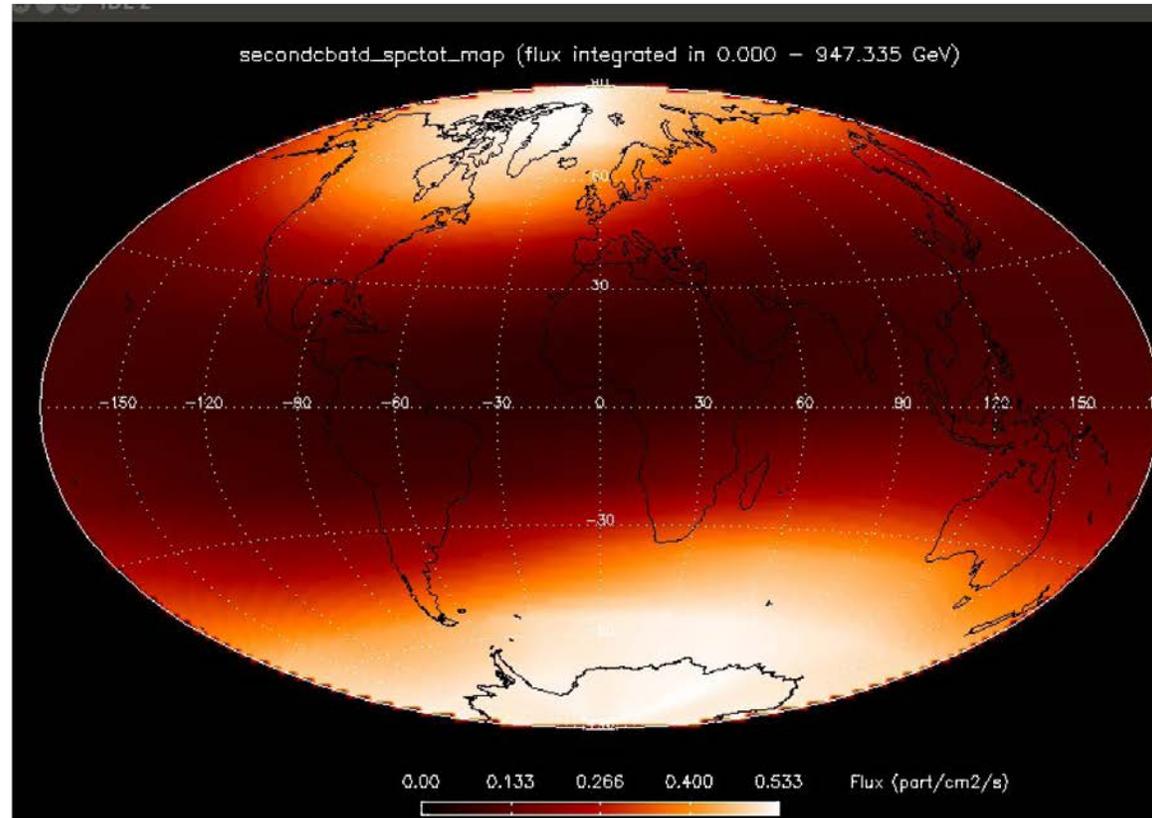


Gamma-ray flux from the Moon in the period May 2011 –November 2013, measured (**FERMI-LAT**) and computed (**FLUKA**) for two different Lunar surface composition models (courtesy of M.Mazziotta, INFN Bari). Primary CR spectra from AMS-02

The neutron albedo from GCR's at 400 km altitude*

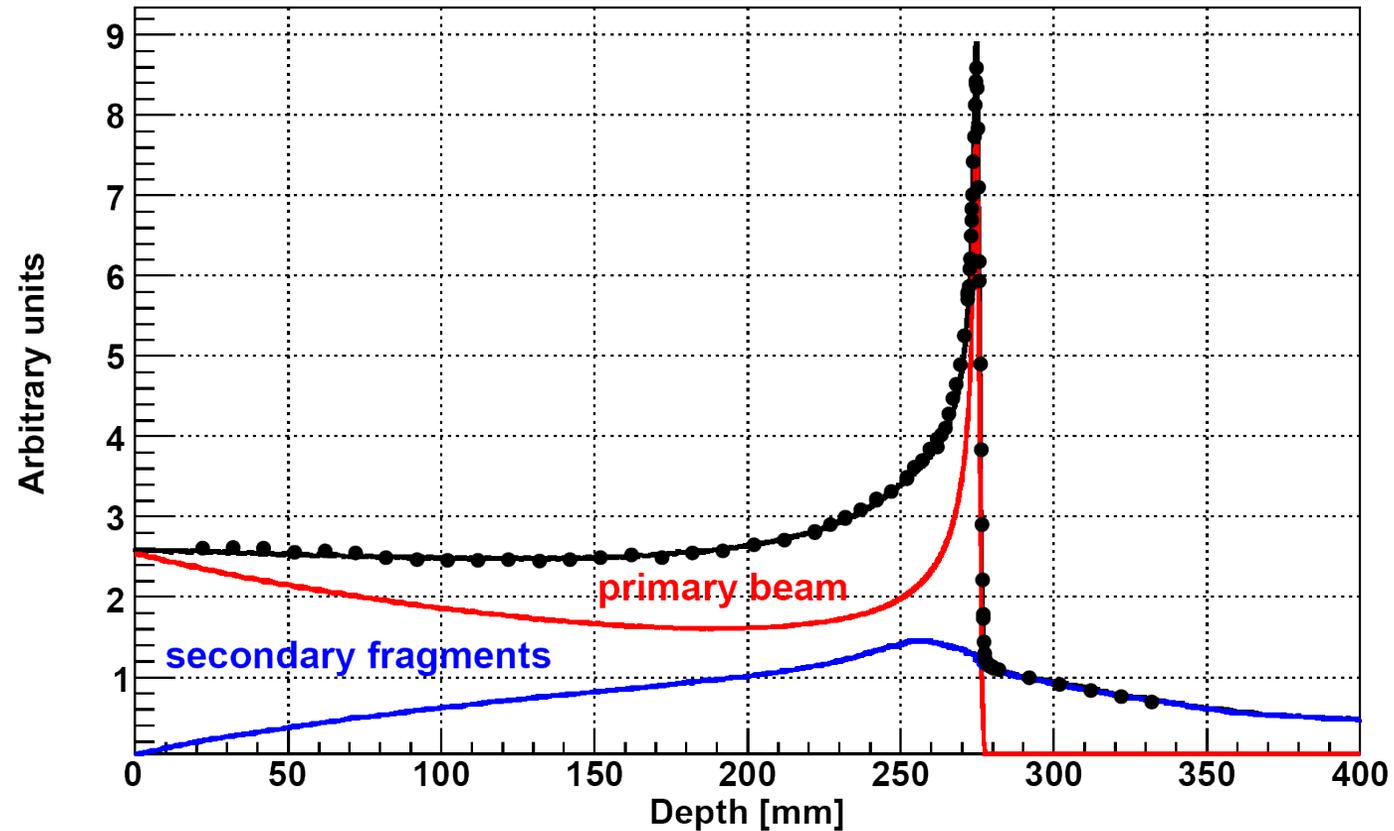


**In collaboration with CEA-Saclay*





Bragg peak in a water phantom: 400 MeV/A C beam:
The importance of fragmentation

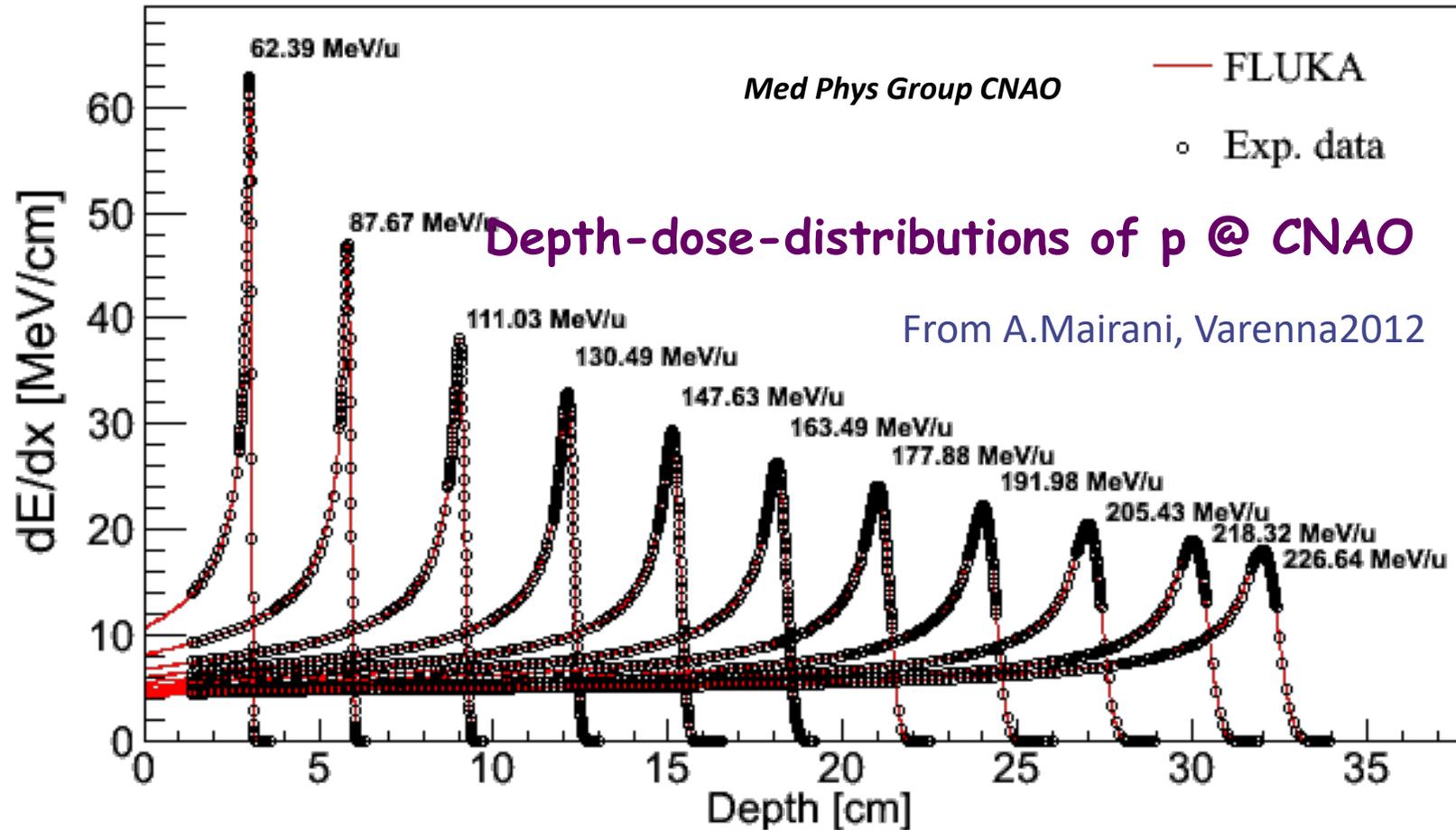


Exp. Data (points) from Haettner et al, Rad. Prot. Dos. 2006
Simulation: A. Mairani PhD Thesis, 2007, Nuovo Cimento C, 31, 2008

Fluka vs hadrontherapy, present: HIT, CNAO, ...

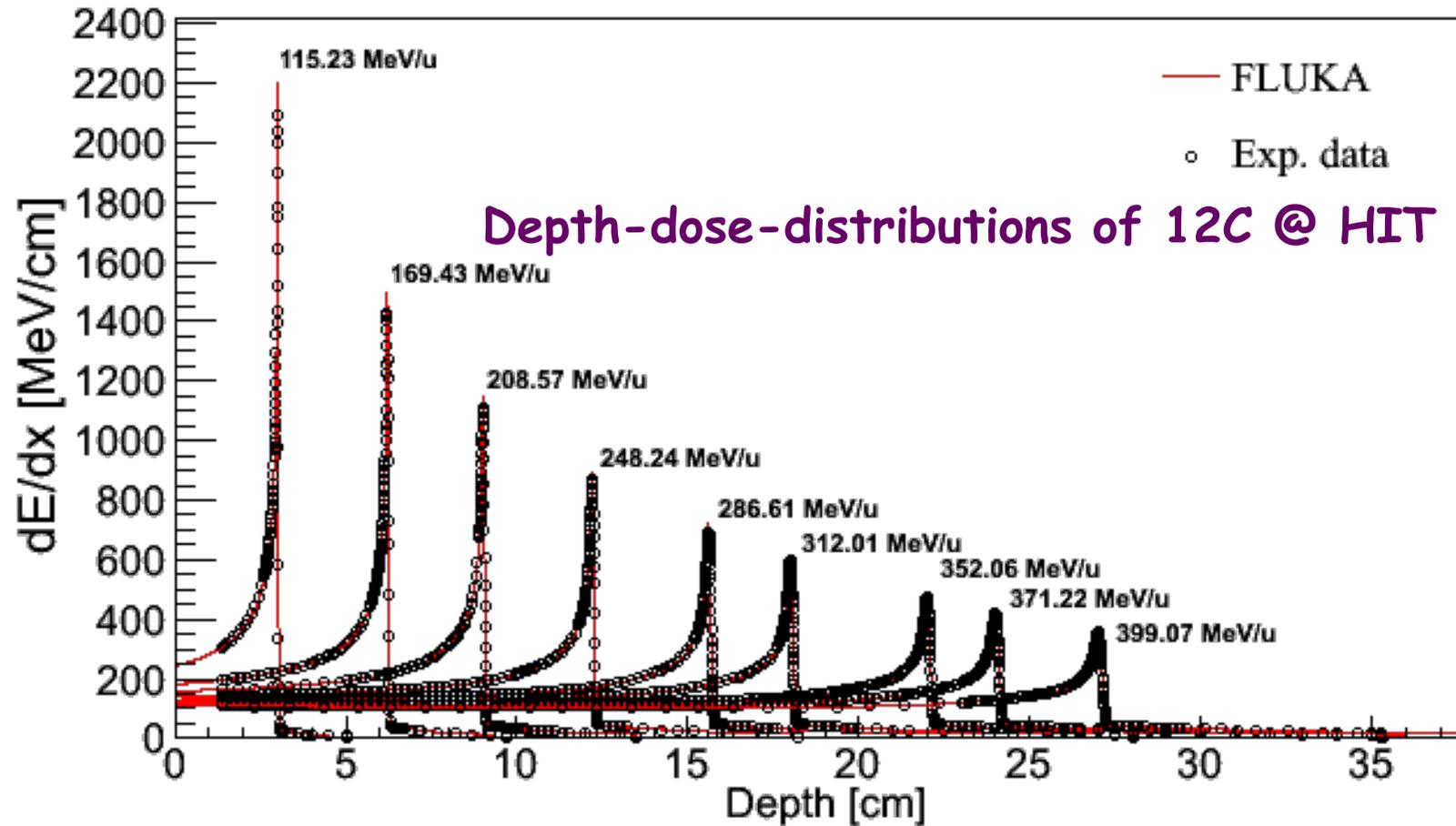


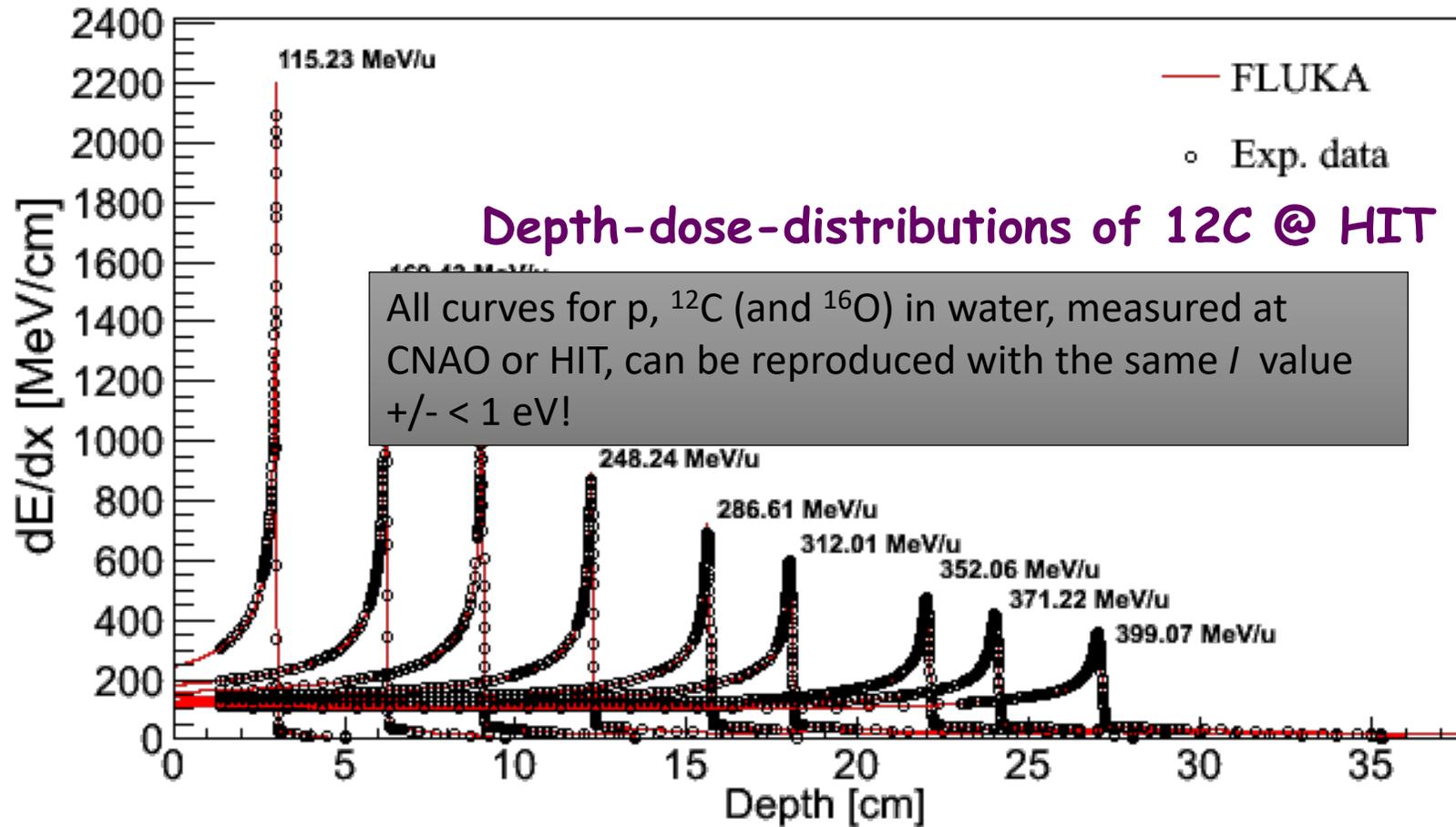
Used for generating p, ¹²C dose vs depth databases then used for TP



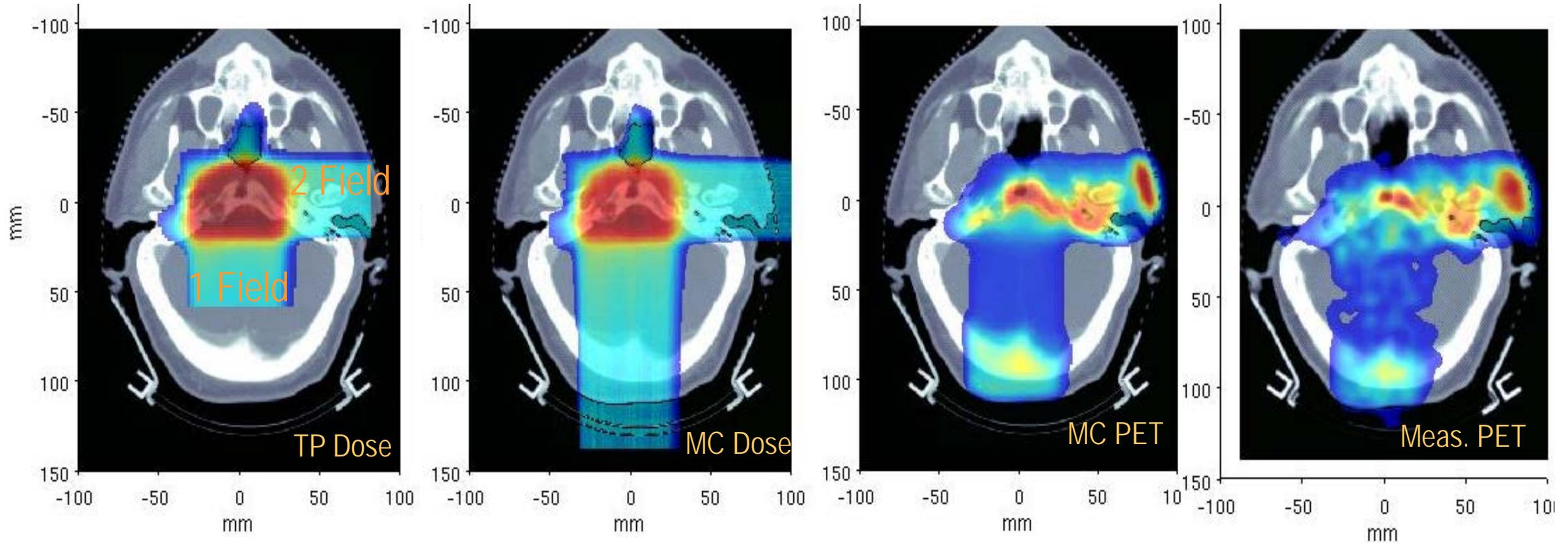
in water wo/with RiFi for the 147 energies in the initial phase of the operation

Fluka vs hadrontherapy, present: HIT, CNAO, ...



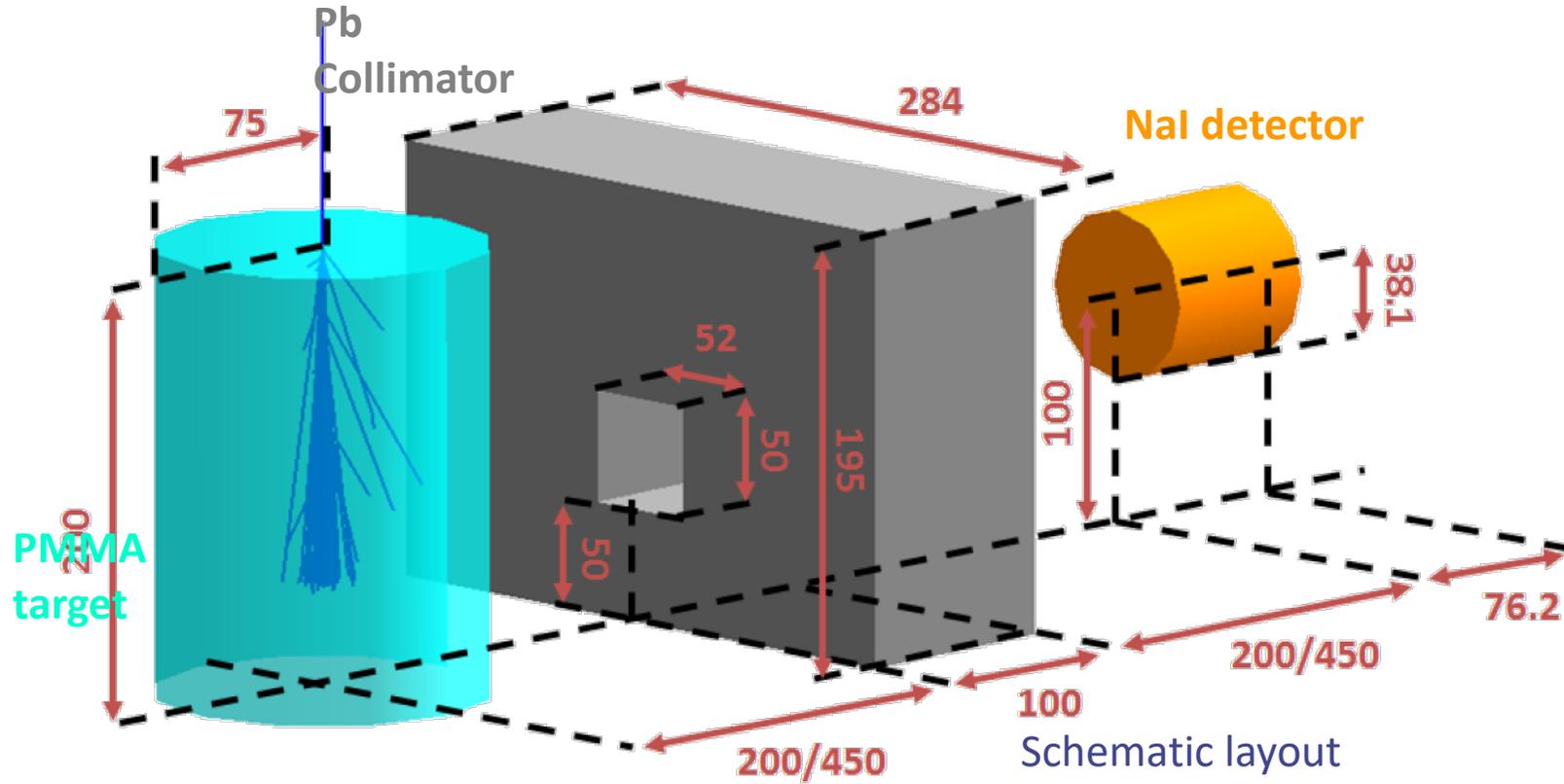


Clival Chordoma, 0.96 GyE / field, $\Delta T_1 \sim 26$ min, $\Delta T_2 \sim 16$ min



K. Parodi et al., PMB52, 3369 (2007)

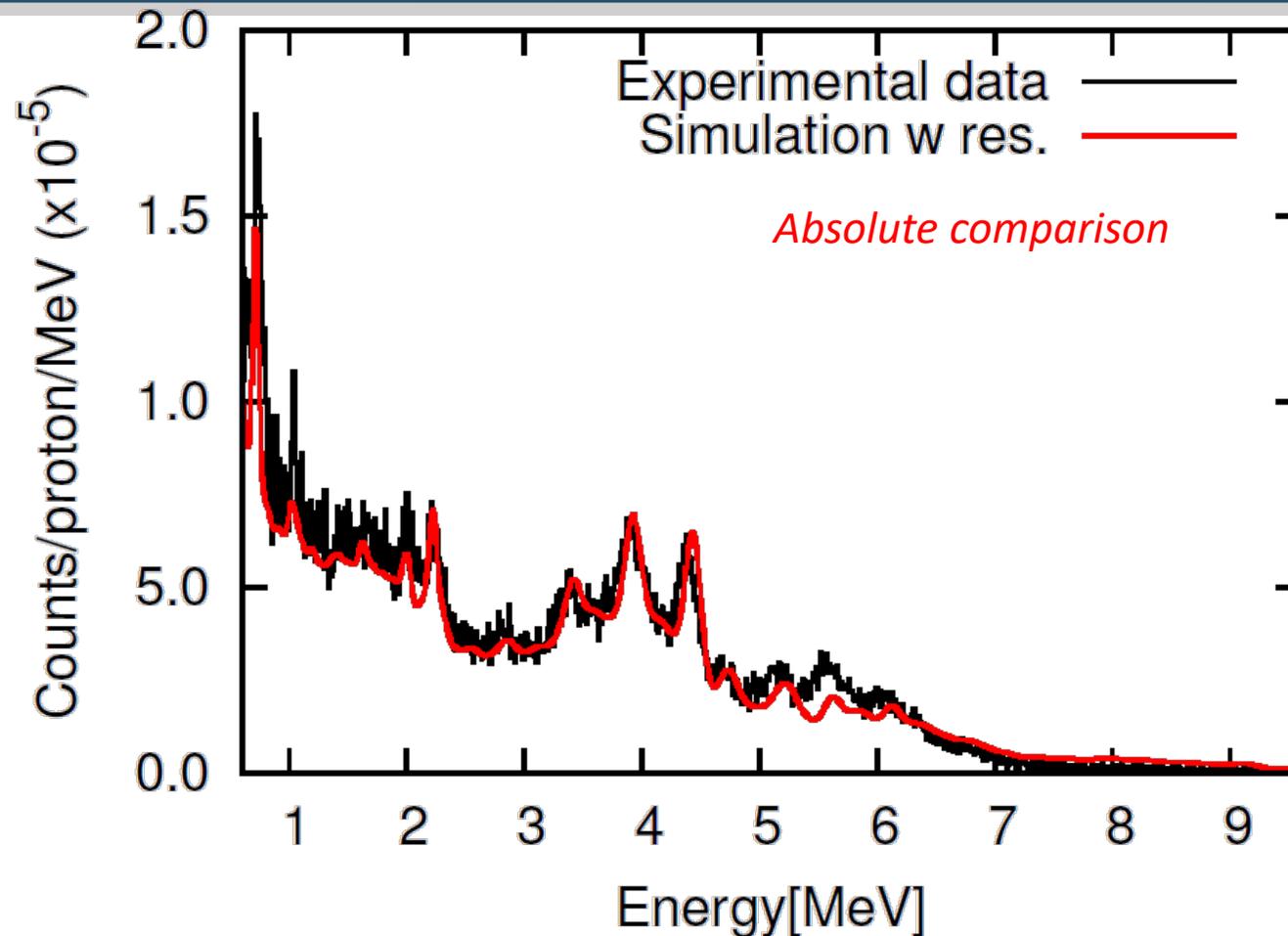
Photon yields by 160 MeV p in PMMA



Schematic layout
(dimensions mm)
from J.Smeets et al., IBA



Photon yields by 160 MeV p in PMMA



Energy spectrum of “photons” after background subtraction (collimator open – collimator closed) for 160 MeV p on PMMA. FLUKA **red line**, data **black line** (J.Smeets et al., IBA, ENVISION WP3)



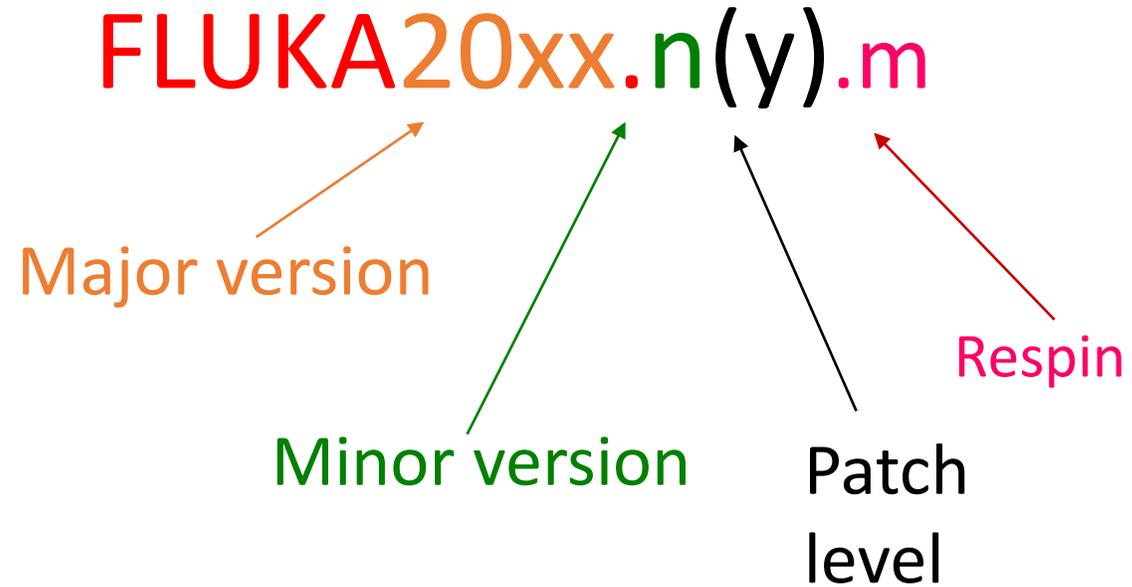
This course is intended to provide users with the basic (and possibly more than basic!) knowledge of:

- a) The most relevant FLUKA instructions and options
- b) The physics models adopted in FLUKA
- c) The different scoring options embedded in FLUKA
- d) The different running options
- e) The tools to plot results
- f) The right approach to the existing documentation
- g) The procedures to overcome difficulties and problems and related debugging tools
- h) etc. etc.

Possible problems:



- People here are not all at the same level of FLUKA knowledge. There are those who already have some experience.
- However we need to start from scratch.
- We apologize to the experienced people and beg them to be patient: it's not excluded a priori that they can learn something new also concerning the very basic elements!
- FLUKA is written in **fortran**. No knowledge of fortran or other languages is needed in this course, however some of the terminology used might be derived from fortran. If this happens and it causes problems, please **ask**!
- FLUKA runs in a **Linux/Mac OSX** environment. A basic knowledge of most common Linux/Unix commands is required, as well as the capability to use a text editor (emacs, vi, gedit..). If some of you has troubles with this, please tell us



In this course we are using: **FLUKA2024.1.0**

The FLUKA license (it is not GPL):



❑ Standard download: **binary library + user routines.**

- FLUKA can be used freely for scientific and academic purposes, *(ad-hoc agreement for commercial purposes)*
- It cannot be used for weapon related applications

❑ It is possible, by explicit signing of license, to download the **source** for researchers of scientific/academic Institutions.

- FLUKA can neither be copied into other codes (not even partially), nor translated into another language without permission

❑ For **commercial use**, trial version (limited in time and random seeds) available.

Please register on www.fluka.organd read the license!



- fluka-users@fluka.org

Users are automatically subscribed here when registering on the web site. It is used to communicate the availability of new versions, patches, etc.

- fluka-discuss@fluka.org

Users are encouraged to subscribe at registration time, but can uncheck the relevant box. It is used to have user-user and user-expert communication about problems, bugs, general inquiries about the code and its physics content

Users are strongly encouraged to keep this subscription

In the next months the FLUKA mailing lists will migrate to a web-based forum... keep in touch to be ready for the transition



Platform: Linux with gfortran (on 64 bit computers) and g77 (on 32 and 64 bit computers)
Mac OSX (both Intel and Apple Silicon) with gfortran

The code should only be compiled/run using operating systems, compilers (and associated) options tested and approved by the development team

Standard Input:

- Command/options driven by “data cards” (ascii file)
A powerful Graphical interface is available
- Standard Geometry (“Combinatorial geometry”): input by “data cards”

Standard Output and Scoring:

- Highly flexible and powerful, sufficient for most purposes
- Output processing and plotting interface available



- ❑ A good FLUKA user **is not one that only masters technically** the program
- ❑ BUT a user that:
 - Indeed masters technically the code;
 - Know its limitations and capabilities;
 - Can tune the simulation to the specific requirements and needs of the problem under study;
 - and most of all
 - **Has a critical judgment on the results**
- ❑ Therefore in this course we will equally focus on:
 - **The technical aspects of the code**
[building your input, geometry, scoring, biasing, extracting results...]
 - as well as
 - **The underlying physics and MC techniques**

The course team



Anna Ferrari



Konstantin Batkov



Stefan Mueller



Alfredo Ferrari



Paola R. Sala

Agenda: Monday



08:00

Welcome

Lanzhou University, Lanzhou, China

08:30 - 08:40

09:00

Introduction to FLUKA

This lecture

Lanzhou University, Lanzhou, China

08:40 - 09:30

10:00

Statistics and sampling

MonteCarlo explained, and the importance of random numbers.

Lanzhou University, Lanzhou, China

09:30 - 10:30

Coffee break

Lanzhou University, Lanzhou, China

10:30 - 11:00

11:00

Installing and running

The very first step, do it together.
Learn how fluka runs and about output files

Lanzhou University, Lanzhou, China

11:00 - 12:00

12:00

FLUKA manual and basic input

Our first commands, and **THE MANUAL!!!!**

Lanzhou University, Lanzhou, China

12:00 - 13:00

13:00

Lunch break

14:00

Lanzhou University, Lanzhou, China

13:00 - 14:30

15:00

FLAIR

Introduction to graphical interface, try it with us

Lanzhou University, Lanzhou, China

14:30 - 15:30

16:00

Exercise: Compound materials

First exercise, all by yourself..

Lanzhou University, Lanzhou, China

15:30 - 16:30

Coffee break

Lanzhou University, Lanzhou, China

16:30 - 17:00

17:00

Exercise: Compound materials (continued)

Please ask, we are here to help

Lanzhou University, Lanzhou, China

17:00 - 18:00

18:00

Agenda: Tuesday



09:00	Geometry Learn how to build a geometry in FLUKA <i>Lanzhou University, Lanzhou, China</i>	08:30 - 10:00
10:00	Exercise: Geometry Apply what you learned <i>Lanzhou University, Lanzhou, China</i>	10:00 - 10:30
	Coffee break <i>Lanzhou University, Lanzhou, China</i>	10:30 - 11:00
11:00	Exercise. Geometry (continued) Ask!! <i>Lanzhou University, Lanzhou, China</i>	11:00 - 12:00
12:00	Physics models I: Hadronic interactions Protons, neutrons, and other particles hit friends: see what comes out <i>Lanzhou University, Lanzhou, China</i>	12:00 - 13:00
13:00	Mount Wuquan visit, and Banquet	



Agenda: Wednesday



08:00

Scoring, and example

How to get results out of Fluka, and plot them, with examples

Lanzhou University, Lanzhou, China

08:30 - 09:30

09:00

Low energy neutrons

Neutrons below 20 MeV: how to deal with them

Lanzhou University, Lanzhou, China

09:30 - 10:30

10:00

Coffee break

Lanzhou University, Lanzhou, China

10:30 - 11:00

11:00

Exercise: Low energy neutrons

Apply what you learned about neutrons

Lanzhou University, Lanzhou, China

11:00 - 13:00

12:00

13:00

Lunch break

14:00

Lanzhou University, Lanzhou, China

13:00 - 14:30

Biasing

How to improve speed, and discover hidden results

Lanzhou University, Lanzhou, China

14:30 - 15:30

15:00

Exercise: Biasing

Experience by yourself the power of biasing

Lanzhou University, Lanzhou, China

15:30 - 16:30

16:00

Coffee break

Lanzhou University, Lanzhou, China

16:30 - 17:00

17:00

Exercise: Biasing (continued)

Need help? ASK!!!

Lanzhou University, Lanzhou, China

17:00 - 18:00

18:00

Agenda: Thursday



08:00	
09:00	Physics models II:EM interactions Learn how FLUKA deals with electrons and photons <i>Lanzhou University, Lanzhou, China</i> 08:30 - 09:30
10:00	Ionization and transport Learn about charged particle transport, and how to optimize your simulation <i>Lanzhou University, Lanzhou, China</i> 09:30 - 10:30
	Coffee break <i>Lanzhou University, Lanzhou, China</i> 10:30 - 11:00
11:00	Exercise: Cutoffs Work with particle thresholds and more <i>Lanzhou University, Lanzhou, China</i> 11:00 - 13:00
13:00	Lunch break

14:00	<i>Lanzhou University, Lanzhou, China</i> 13:00 - 14:30
15:00	Heavy ions How ion interact among themselves, and which options you need <i>Lanzhou University, Lanzhou, China</i> 14:30 - 15:30
16:00	Medical applications I Processes and input cards for medical applications <i>Lanzhou University, Lanzhou, China</i> 15:30 - 16:30
	Coffee break <i>Lanzhou University, Lanzhou, China</i> 16:30 - 17:00
17:00	Medical applications II Guided application to a therapy-like example <i>Lanzhou University, Lanzhou, China</i> 17:00 - 18:00
18:00	

Agenda: Friday



08:00

Radioactivity

How to calculate activation of materials,
residual doses, inventory, time evolution

Lanzhou University, Lanzhou, China

08:30 - 09:30

09:00

Exercise: Activation

Apply

Lanzhou University, Lanzhou, China

09:30 - 10:30

10:00

Coffee break

Lanzhou University, Lanzhou, China

10:30 - 11:00

11:00

Exercise: Activation (continued)

Lanzhou University, Lanzhou, China

11:00 - 12:00

12:00

Advanced topics

What you can learn in the advanced
course...

Lanzhou University, Lanzhou, China

12:00 - 13:00

13:00

Lunch break

14:00

Lanzhou University, Lanzhou, China

13:00 - 14:30

Handling of errors and crashes

15:00

Examples of common errors and their solution

Lanzhou University, Lanzhou, China

14:30 - 15:30

User cases, Q&A

16:00

Free discussion, bring your examples,

Lanzhou University, Lanzhou, China

15:30 - 16:30

Coffee break

Lanzhou University, Lanzhou, China

16:30 - 17:00

17:00

User cases, Q&A

Doubts, And wishes

Lanzhou University, Lanzhou, China

17:00 - 18:00

18:00

Closing

Lanzhou University, Lanzhou, China

18:00 - 18:10



- All the material for this course is available in the usb stick you received
 - Lectures : the presentations
 - Exercises: subdirectories with
 - Instruction
 - Input files
 - Flair files
- Solutions are also there, try not to look at them until the end of the dedicated time.
- Do NOT work on the usb: always create a new directory for each example/exercise



- Exercises will be on your own
- With a description to guide you
- And an input file to start with
- Teachers and support will be around, call us if you need help

- Also during lectures, if you have doubts please do not be shy, ***ask!***

Thanks for your attention!