

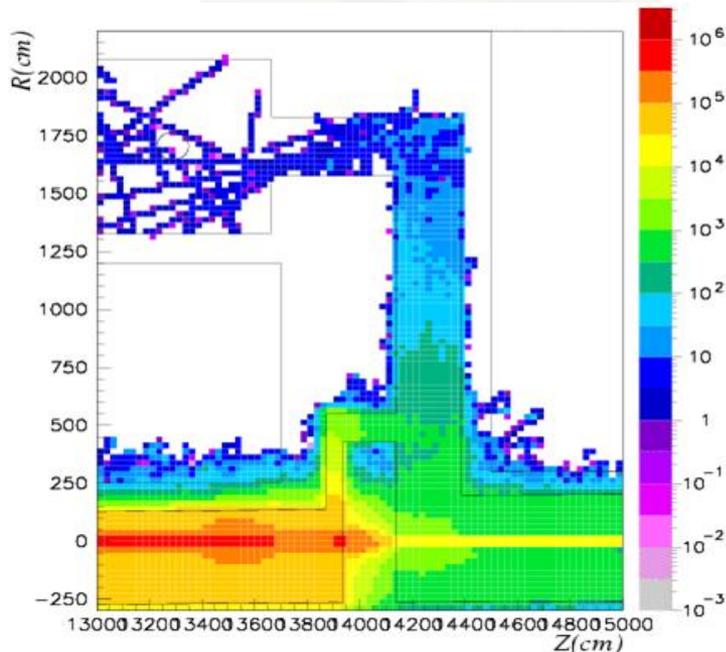


# Advanced Biasing

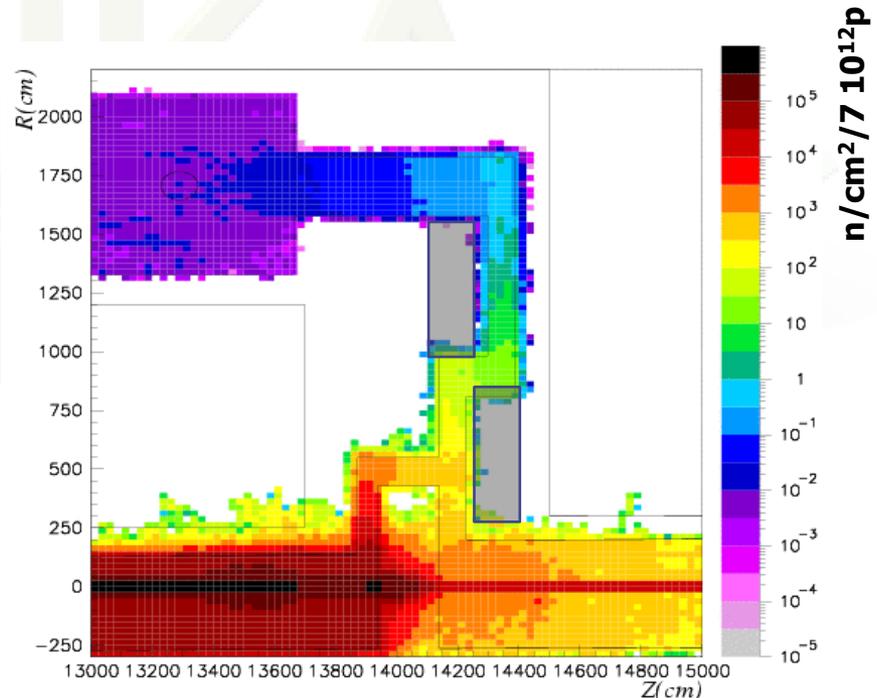
Advanced FLUKA Course

# Concept

- Variance reduction techniques in Monte Carlo calculations **reduce the computer time** *or the opposite* to obtain results of sufficient precision in the phase-space region of interest.
- **Remember:** that precision is not the only requirement for a Good Monte Carlo calculation. Even a zero variance calculation cannot accurately predict natural behavior if **other sources of error** are not minimized.



No Bias and no maze



Region Biasing + maze

# Monte Carlo Flavors

	<b>Microscopic Analog</b>	<b>Microscopic Biased</b>	<b>Macroscopic Analog</b>
Physics Models	Theoretical	Theoretical	Parameterizations
PDF sampling	Physical processes	Artificial distributions	Fits & Data
Predict Average	Yes	Yes	Yes
Predict Higher Moments	Yes	-	-
Preserves Correlations	Yes	-	-
Reproduces Fluctuations	Yes	-	-
Rare events	-	Yes	-
Predictability	Yes	Yes	-
Convergence	Slow	Fast privileged regions	Fast
Safe	Yes	Almost	-

# Figure of Merit

- Computer cost of an estimator

$$FOM = \sigma^2 \times t$$

$$\sigma^2 = \text{Variance} \propto 1/N, \quad t = \text{CPU time} \propto N$$

- some biasing techniques are aiming at reducing the  $\sigma^2$ , others at reducing  $t$
- often reducing  $\sigma^2$  increase  $t$ , and vice versa
- therefore, minimizing  $\sigma^2 \times t$  means to reduce  $\sigma$  at a faster rate than  $t$  increases, or vice versa
- $\Rightarrow$  the choice depends on the problem, and sometimes a combination of several techniques is most effective
- **bad judgment, or excessive "forcing"** on one of the two variables, can have catastrophic consequences on the other one, making computer cost "explode"

# Biasing Techniques

FLUKA offers the following possibilities for biasing

- Importance Biasing (**BIASING**)
- Weight window (**WW-FACTOr**, **WW-THRESH**, **WW-PROFIle**)
- Leading Particle Biasing (**EMF-BIAS**)
- Multiplicity Tuning (**BIASING**)
- Biased down scattering for neutrons, *only for experts* (**LOW-DOWN**)
- Non analogue absorption (**LOW-BIAS**)
- Biasing Mean free paths (**LAM-BIAS**)
- User defined biasing (**usbset.f**, **usimbs.f**)

# Importance Biasing

- Importance biasing combines two techniques:
  - **Surface splitting**: Reduces  $\sigma$  but increases  $t$
  - **Russian roulette**: which does the opposite
- It is the **simplest, most safe** and easiest to use of all biasing
- The user assigns a **relative importance** to each geometry region (the absolute value doesn't matter) based on:
  - **expected fluence attenuation** with respect to other regions
  - **probability of contribution** to score by particles entering the region
- Importance biasing is commonly used to **maintain a constant particle population**, compensating for attenuation due to absorption or distance.
- In FLUKA it can be **tuned per type of particle**

# Importance Biasing

## Surface Splitting

- If a particle crosses a region boundary coming from a region of importance  $I_1$  and enters a region of higher importance  $I_2 > I_1$ :
  - the particle is replaced on average by  $n = I_2/I_1$  identical particles with the same characteristics
  - the weight of the daughter is multiplied by  $I_1/I_2$
- If  $I_2/I_1$  is too large, excessive splitting may occur, which will be prevented by FLUKA

## Russian Roulette

- RR acts in the opposite direction: from higher importance  $I_1$  to lower importance  $I_2 < I_1$ 
  - The particle is submitted a random survival test with a chance of  $I_2/I_1$  the particle survives with its weight increased by a factor  $I_1/I_2$
  - with chance  $1 - I_2/I_1$  the particle is killed

# Importance Biasing Problems

- Although important biasing is relatively easy and safe to use, there are a few cases where caution is recommended

<b>I=?</b>    E	I=8	D
	I=4	C
	I=2	B
	I=1	A

Which importance shall we give to region E? Whatever value we choose we will get an inefficient splitting/RR at some of the boundaries.

- Another case is that of **splitting in vacuum** (or air). Splitting daughters are **strongly correlated**. It must be made sure that their further histories are differentiated enough to forget their correlation.
- The above applies in part also to **muons** (the differentiation provided by **multiple scattering** and by **Landau  $dE/dx$**  fluctuations is not always sufficient).

# Leading Particle Biasing

- Leading Particle Biasing is available only for  $e^+$ ,  $e^-$  and photons
- It is generally used to avoid the geometrical increase with energy of the number of particles in the electromagnetic shower
- It is a characteristic of EM interactions that 2 particles are present in the final state
- if LPB is activated only one is randomly retained and its weight is adjusted so as to conserve weight  $\times$  probability
- The most energetic of the two particles is kept with higher probability
- LPB is reducing  $t$ , but increases  $\sigma$  by introducing large weight fluctuations. Therefore it should nearly always be backed up by WW
- Very useful for shielding calculations at both electron and proton accelerators

# Weight Window

- The WW technique is a combination of splitting and RR, but it is based on the **absolute value** of the weight of each **individual particle**, rather than on relative region importance
- The user sets **an upper and a lower weight limit**, generally as a function of region, energy and particle
- Particles having a **weight larger than the upper limit are split**, those with weight **smaller than the lower limit are submitted to RR**  $\Rightarrow$  killed or **put back** "inside the window"
- WW is a more **powerful** biasing tool than Importance Biasing, but it requires also more experience and patience to set it up correctly  
*"It is more an art than a science"* (From MCNP Manual)
- Use of the WW is essential whenever other biasing techniques generate large weight fluctuations in a given phase space region.
- Works with all particles with the exception of **Low Energy neutrons** (see **WW-PROFILE**)

## Weight Windows - 2

Killing a particle with a very low weight (with respect to the average for a given phase space region) decreases  $t$  but has very little effect on the score (and therefore on  $\sigma$ )

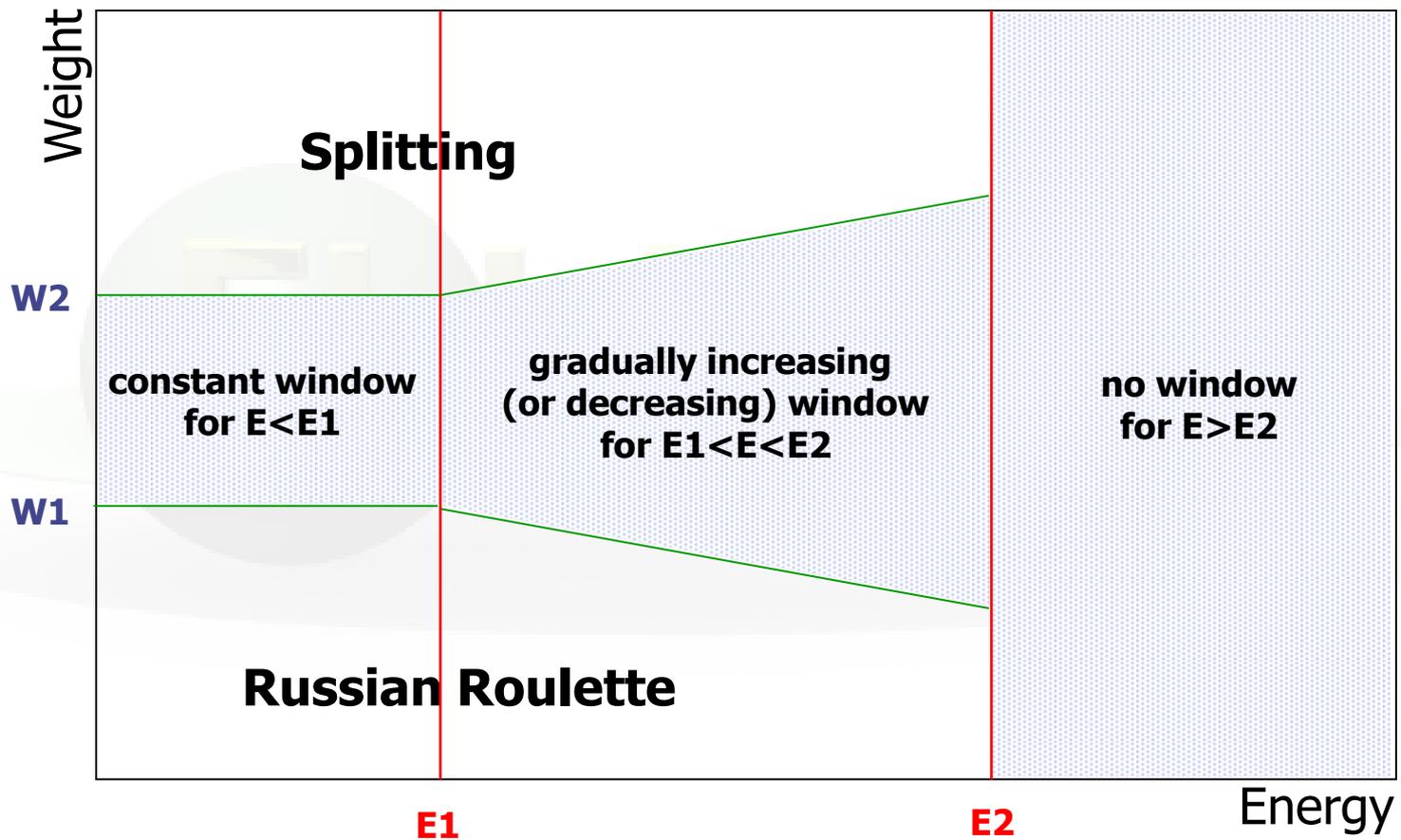
Splitting a particle with a large weight *increases*  $t$  (in proportion to the number of additional particles to be followed) but at the same time *reduces*  $\sigma$  by avoiding large fluctuations in the contributions to scoring.



The global effect is to reduce  $\sigma^2 t$

A too wide window is of course ineffective, but also too narrow windows should be avoided. Otherwise, too much CPU time would be spent in repeated splitting / Russian Roulette. *A typical ratio between the upper and the lower edge of the window is about 10.* It is also possible to do Russian Roulette without splitting (setting the upper window edge to infinity) or splitting without Russian Roulette (setting the lower edge to zero)

# Weight Windows - 3



# Weight Windows - 4

WW-FACTO	13.0	120.0	1.5	27.0	31.0	2.0
----------	------	-------	-----	------	------	-----

## Defines Weight Windows in selected regions

**WHAT(1)  $\geq 0.0$  : Window "bottom" weight**

**$< 0.0$  : resets to -1.0** (no Russian Roulette, Default)

Weight below which *Russian Roulette* is played at the lower energy threshold (set by WW-THRES).

**WHAT(2)  $> 1.7 * \text{WHAT}(1)$  : Window "top" weight**

**$= 0.0$  : ignored**

**$\leq 1.7 * \text{WHAT}(1)$  : resets to infinity**

(no Splitting, Default)

Weight above which *Splitting* is applied at the lower energy threshold (set by WW-THRES).

**WHAT(3)  $> 0.0$  : Multiplicative factor** (Default: 1.0)

**$= 0.0$  : ignored**

**$< 0.0$  : resets to 1.0**

Factor to be applied to the two energy thresholds for Russian Roulette / Splitting (set by WW-THRES)

# Weight Windows - 5

WW-FACTO	13.0	120.0	1.5	27.0	31.0	2.0
----------	------	-------	-----	------	------	-----

**WHAT(4)** = lower bound of the region indices (Default = 2.0)

**WHAT(5)** = upper bound of the region indices (Default = WHAT(4))

**WHAT(6)** = step length in assigning indices (Default = 1.0)

**SDUM** : a number from 1.0 to 5.0 in any position, indicating the **low-energy neutron weight-window profile** to be applied in the regions selected (see WW-PROFI). (Default = 1.0)  
 = blank, zero or non numerical: ignored  
 < 0.0 : resets to 1.0

**Attention:** Option WW-FACTO alone is not sufficient to define a weight window. One or more WW-THRES cards are also necessary in order to activate the window.

# Weight Windows - 6

WW-THRES	2.0	0.05	2.4	3.0	7.0	0.0
----------	-----	------	-----	-----	-----	-----

**Defines the energy limits and particle-dependent modification factors**

- WHAT(1) > 0.0: upper kinetic energy threshold (GeV)**  
 Low-energy neutrons: lower group number (included)  
 = 0.0: ignored  
 < 0.0: any previously selected threshold is cancelled
- WHAT(2) >= 0.0 and < WHAT(1): lower kinetic energy threshold (GeV)**  
 Low-energy neutrons: upper group number (included)  
 < 0.0 or > WHAT(1): WHAT(2) is set = WHAT(1)
- WHAT(3) > 0.0: amplification factor** to define the weight window width at the higher energy threshold represented by WHAT(1).  
 The weight window at the higher energy threshold is obtained by multiplying by WHAT(3) the upper weight limit and by dividing by the same factor the lower weight limit. (Default = 10.0)
- < 0.0: |WHAT(3)| multiplication factor** for the lower and upper weight limits for the particles selected by WHAT(4-6) (Default = 1.0)

# Weight Windows - 7

WW-THRES	2.0	0.05	2.4	3.0	7.0	0.0
----------	-----	------	-----	-----	-----	-----

**WHAT(4) = lower bound of the particle indices** (Default = 1.0)

Note that particle index 40 indicates low-energy neutrons (for this purpose only!). Particle index 8 indicates neutrons with energy > 19.6 MeV.

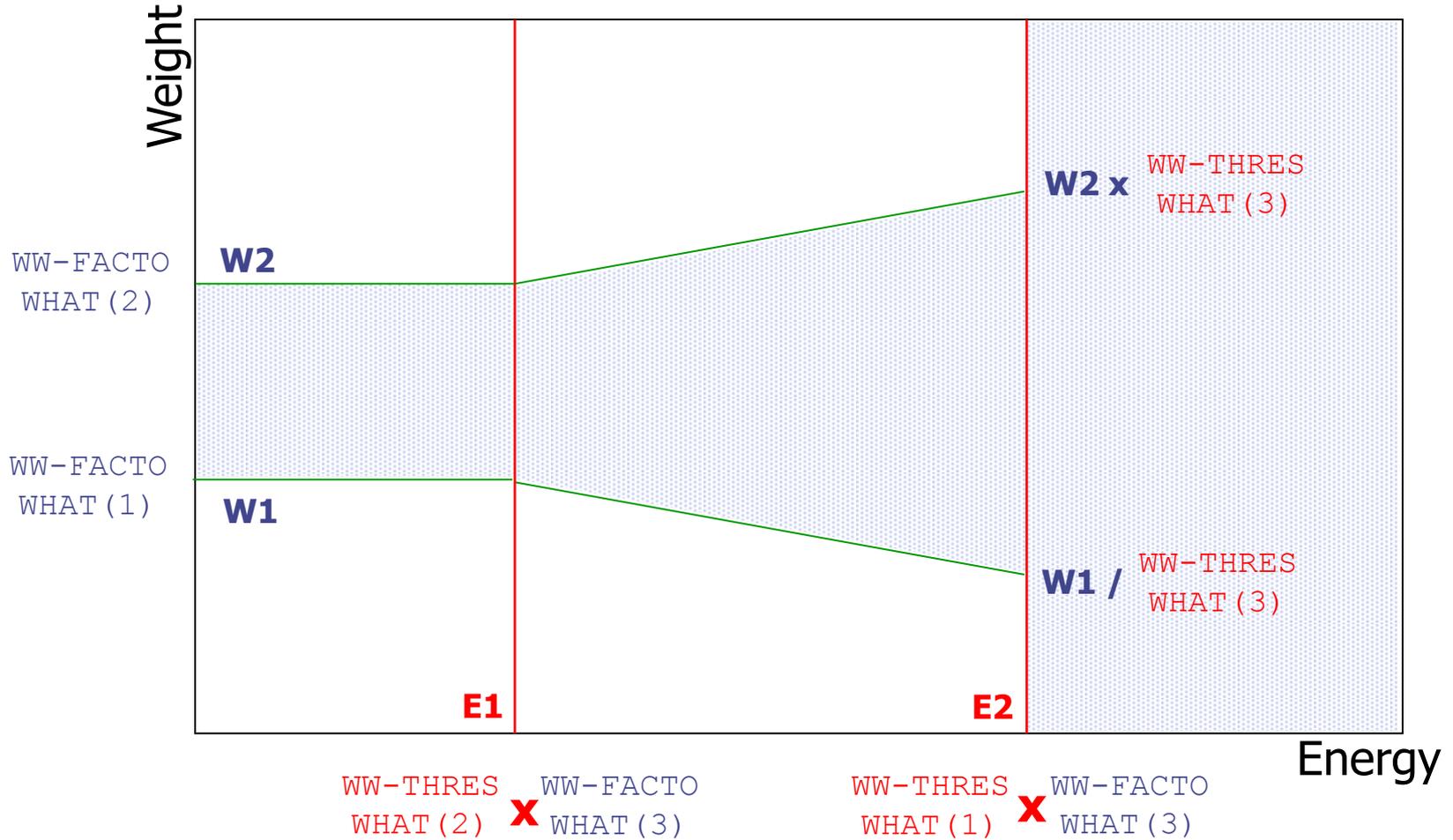
**WHAT(5) = upper bound of the particle indices**  
(Default = WHAT(4) if WHAT(4) > 0, all particles otherwise)

**WHAT(6) = step length in assigning indices** (Default = 1.0)

**SDUM = PRIMARY:** the weight window applies also to primary particles (default)  
**= NOPRIMARY:** the weight window doesn't apply to primaries

# Weight Windows - 8

region dependent  
particle type dependent



# Selecting Weight Windows - 1

BIASING                    0.0                    0.0                    4.64                    8.                    18.                    2. PRINT

*FLUKA output file:*

Hadron importance RR/Splitting counters

Reg. #	N. of RR	<Wt> in	<Wt> kil	Reg. #	N. of RR	<Wt> in	<Wt> kil	Reg. #	N. of RR	<Wt> in	<Wt> kil
1	0.00E+00	0.00E+00	0.00E+00	2	0.00E+00	0.00E+00	0.00E+00	3	1.15E+05	9.31E-01	4.70E-02
Reg. #	N. of Sp	<Wt> in	<Wt> out	Reg. #	N. of Sp	<Wt> in	<Wt> out	Reg. #	N. of Sp	<Wt> in	<Wt> out
1	0.00E+00	0.00E+00	0.00E+00	2	0.00E+00	0.00E+00	0.00E+00	3	0.00E+00	0.00E+00	0.00E+00
Reg. #	N. of RR	<Wt> in	<Wt> kil	Reg. #	N. of RR	<Wt> in	<Wt> kil	Reg. #	N. of RR	<Wt> in	<Wt> kil
4	1.36E+04	4.66E-01	1.47E-01	5	8.97E+03	3.22E-01	1.06E-01	6	6.03E+03	2.16E-01	7.10E-02
Reg. #	N. of Sp	<Wt> in	<Wt> out	Reg. #	N. of Sp	<Wt> in	<Wt> out	Reg. #	N. of Sp	<Wt> in	<Wt> out
4	1.01E+05	9.99E-01	7.64E-01	5	9.25E+04	6.80E-01	5.23E-01	6	8.24E+04	4.65E-01	3.55E-01

"N. of RR"    --> Number of FLUKA particles entering a region and which are not split (i.e., particles undergoing Russian Roulette as well as neither Russian Roulette nor splitting)

"<Wt> in"    --> Average weight of these particles

"<Wt> kil"    --> Average weight of particles killed after being submitted to Russian Roulette

"N. of Sp"    --> Number of FLUKA particles entering the region and which are split

"<Wt> in"    --> Average weight of these particles

"<Wt> out"    --> Average weight of particles after being submitted to splitting

# Selecting Weight Windows - 2

where

$$\begin{aligned} A &= \text{"N. of RR"} + \text{"N. of Sp"} \\ &= \textit{total number of particles entering the region} \end{aligned}$$

$$\begin{aligned} B &= (\text{"<Wt> in"}_{\text{RR}} * \text{"N. of RR"}) + (\text{"<Wt> in"}_{\text{Sp}} * \text{"N. of Sp"}) \\ &= \textit{total weight of the particles entering the region} \end{aligned}$$

$$B/A = \textit{average weight of the particles entering the region}$$

*Note -1:* RR and splitting arising from Weight-Window biasing (options WW-FACTOR, WW-THRESH, WW-PROFI) or from multiplicity biasing (WHAT(2) in option BIASING) are not accounted for in the counters.

*Note - 2:* Separate counters are printed for hadrons/muons, electrons/photons and low-energy neutrons (referring to importance biasing requested by BIASING, respectively, with WHAT(1) = 1.0, 2.0 and 3.0, or = 0.0 for all).

# Selecting Weight Windows - 3

## Strategy:

1. run without any biasing and print counter, *e.g.*,

```
BIASING      0.0      1.0      1.0      1.      9.      PRINT
```

2. analyze counter and adjust region importance biasing, *e.g.*, according to the inverse of the attenuation in shielding, add other biasing, *e.g.*, leading particle biasing run and print counter again

```
BIASING      0.0      1.0      1.0      1.      9.      PRINT
BIASING      0.0      1.0      1.47     4.
BIASING      0.0      1.0      2.15     5.
BIASING      0.0      1.0      3.16     6.
BIASING      0.0      1.0      4.64     7.
BIASING      0.0      1.0      4.64     8.
```

3. analyze counter, select Weight Windows (WW-THRES, WW-FACTO) around average weights and perform final (high-statistics) run

# Biassing Mean Free Paths

## Multiplicity Tuning

## BIASING

- Multiplicity tuning is meant to be to **hadrons** what **LPB** is for electrons and photons.
- A hadronic nuclear interaction at LHC energies can end in **hundreds of secondaries**. Except for the leading particle, many secondaries are of the same type and have **similar energies** and other characteristics
- The user **can tune the average multiplicity** in different regions

## Interaction Length

## LAM-BIAS

- **Mean life / average decay length** of unstable particles can be artificially shortened
- Can **increase generation rate** of decay products without discarding the parent
- For **hadrons the mean free path for nuclear inelastic interactions** can be artificially decreased. Useful for very thin targets, and also for photonuclear reactions where the cross section is relatively small

# Non analogue absorption

- Implemented in most **low-energy** neutron transport codes, where the user must choose between two options:
  - At each neutron collision either **analog scattering or absorption** according to the actual physical probability  $\sigma_s/\sigma_\tau$  and  $1 - \sigma_s/\sigma_\tau$ .
- In FLUKA there is a third choice: the user **can force** the neutron absorption probability to take an **arbitrary value**, pre-assigned on a region-by-region basis as a function of energy.

## When and How

- A **small survival probability** is often assigned to **thermal neutrons** to limit the number of scatterings in non-absorbing media
- Also very useful in materials with **unusual scattering properties** (e.g. Iron)
- Survival probabilities **too small** with respect to the physical one  $\sigma_s/\sigma_\tau$  may introduce large weight fluctuations due to the very different number of collisions suffered by individual neutrons. In these cases a WW should be applied.
- Card: **LOW-BIAS**      *Also called survival biasing*

# Card: RADDECAY [1/2]

\* 1) request radioactive decays

```
RADDECAY      1.0      0      3.0      0000099999      0
```

<b>RADDECAY</b>	Decays: Active ▼	Patch Isom: ▼	Replicas: 3.0
h/μ Int: ignore ▼	h/μ LPB: ignore ▼	h/μ WW: ignore ▼	e-e+ Int: ignore ▼
e-e+ LPB: ignore ▼	e-e+ WW: ignore ▼	Low-n Bias: ignore ▼	Low-n WW: ignore ▼
	decay cut: 0.0	prompt cut: 99999.0	Coulomb corr: ▼

- WHAT(1)** = 1 **radioactive decays activated for requested cooling times**  
Decays: Active "activation study case": time evolution calculated analytically for *fixed* (cooling) times. Daughter nuclei as well as associated radiation is considered at these (fixed) times
- > 1 **radioactive decays activated in semi-analogue mode**  
Semi-Analogue each radioactive nucleus is treated like all other unstable particles (random decay time, daughters and radiation), all secondary particles/nuclei carry time stamp ("age")
- WHAT(2)** > 0 **isomer "production" activated**  
Patch Isom: On
- WHAT(3)** # **number of "replicas" of the decay of each individual nucleus**  
Replicas:

# Card: RADDECAY [2/2]

## RADDECAY

h/μ Int: ignore ▼  
e-e+ LPB: ignore ▼

Decays: Active ▼  
h/μ LPB: ignore ▼  
e-e+ WW: ignore ▼  
decay cut: 0.0

Patch Isom: ▼  
h/μ WW: ignore ▼  
Low-n Bias: ignore ▼  
prompt cut: 99999.0

Replicas: 3.0  
e-e+ Int: ignore ▼  
Low-n WW: ignore ▼  
Coulomb corr: ▼

## WHAT(4)

h/m Int .. Low-n WW

**switch for applying various biasing features only to prompt radiation or only to particles from radioactive decays**

9 digits, each responsible for a different biasing

Example:

5th digit, e+/e-/gamma leading particle biasing applied  
000010000 to prompt radiation only  
000020000 to decay radiation only  
000030000 to both

Default: 11111111 (or blank as above)

## WHAT(5)

decay cut: #  
prompt cut: #

**multiplication factors to be applied to transport cutoffs**  
10 digits, first five for decay radiation, second five for prompt radiation (see manual)

Special cases:

0000099999 kill EM cascade for prompt radiation  
9999900000 kill EM cascade for residual radiation

# User written biasing

FLUKA offers the following routines for **user-written biasing**

- **ubsset.f: User Biasing SETting**
  - called after reading in the input file and before the first event
  - allows to alter almost any biasing weight on a region-dependent basis
- **usimbs.f: USer defined IMportance Biasing**
  - if activated, called at every particle step
  - allows to implement any importance biasing scheme based on region number and/or phase space coordinates
- **udcdrl.f: User defined DeCay DiRection biasing and Lambda**
  - only for neutrinos emitted in decays: bias the direction of emitted neutrino

**Not biasing by itself, but it could be used for generating biased runs**

- **source.f: User written source**
  - to sample primary particle properties from distribution in space, energy, time...

# User Defined Importance Biasing

- Typical problem: Spend a lot of time to write the problem input, geometry and on the first runs you realize that statistics are not good
- First method (and safest) is to introduce region importance biasing. In FLUKA you can introduce it with two ways:
  - 1<sup>st</sup> Manually slice the geometry and increase the number of regions. Modifying an existing geometry to introduce biasing can be a very cumbersome process
  - 2<sup>nd</sup> Introduce the “importance biasing” information with a user fortran routine **independent of the regions defined in the geometry**
- Routine: **usimbs.f**

## USER defined IMportance BiaSing

Allows to implement any importance biasing scheme based on region number and/or phase space coordinates

# User Defined Importance Biasing

- Enable the call to **USIMBS** routine with the **BIASING** card:
  - WHAT(1) Particles to be biased
  - **WHAT(2) and WHAT(3)  $\neq$  1.0** (Any value  $\neq$  1.0)
  - WHAT(4) Lower bound of region
  - WHAT(5) Upper bound of region
  - WHAT(6) Step
  - SDUM = **USER**



## Remember:

- If WHAT(3)=1 for a region, the routine will not be called during tracking of particles inside that region
- Cannot have both normal importance BIASING with cards and the routine at the same time

# usimbs.f – Routine

The routine is called on every particle step!

**WARNING:** can slow down the execution speed! Use with caution.

## Input:

- Region information at the **beginning** and **end** of the step
- X,Y,Z coordinates through the **TRACKR** common.  
Beginning: X, Y, Ztrack(0)  
End: X, Y, Ztrack(Ntrack)
- **Particle** type and **Energy** could be used for even more advanced biasing schemes

## Output:

The routine must return the importance ratio to the new position (end/beginning) in the variable **FIMP**

## Entry: USIMST

Split the particles step between interactions in half (or any other user defined value)

```
SUBROUTINE USIMBS ( MREG, NEWREG, FIMP )

INCLUDE '(DBLPRC)'
INCLUDE '(DIMPAR)'
INCLUDE '(IOUNIT)'

*
*-----*
* User defined IMportance BiaSing: *
* Created on 02 july 2001 by Alfredo Ferrari & Paola Sala *
*                               Infn - Milan *
* Last change on 09-jul-01 by Alfredo Ferrari *
* Input variables: *
*     Mreg = region at the beginning of the step *
*     Newreg = region at the end of the step *
* (thru common TRACKR): *
*     Jtrack = particle id. (Paprop numbering) *
*     Etrack = particle total energy (GeV) *
*     X,Y,Ztrack(0) = position at the beginning of the step *
*     X,Y,Ztrack(Ntrack) = position at the end of the step *
* Output variable: *
*     Fimp = importance ratio (new position/original one) *
*-----*

INCLUDE '(TRACKR)'
FIMP = ONEONE
RETURN
END

*====*
* Entry USIMST: *
* Input variables: *
*     Mreg = region at the beginning of the step *
*     Step = length of the particle next step *
* Output variable: *
*     Step = possibly reduced step suggested by the user *
*====*

ENTRY USIMST ( MREG, STEP )
IF ( STEP .GT. ONEONE ) STEP = HLFHLF * STEP
RETURN

*==== End of subroutine Usimbs *====*
END
```

# usimbs.f – Important Notes

- The routine has **only a relative effect** on the weight of the particle  
⇒ Beam particles can have any weight.
- Importance ratio will be limited by **WW-THRES** card
- The Russian Roulette / Splitting will take place at the middle (defined by the ENTRY) of the step, **and not on a fixed region boundary.**
- The biasing position will be a **little fuzzy** depending on the particle step. This has a visible effect when it is applied to **low density materials** (i.e. air)
- Results will **similar but not the same** as with the manual region biasing.
- Is a great time saver for complex geometries, as well different biasing schemes
- Combined with the particle type and energy could become very powerful

# usimbs.f: Simple Example

- Biasing a factor of 2 every 50 cm on the Z direction from 100cm to 500cm

```
ZSTART = ZTRACK(0)
IF (ZSTART .LT. 100.0D0) THEN
    FSTART = ONEONE
ELSE IF (ZSTART .GT. 500.0D0) THEN
    FSTART = TWOTWO ** NINT((500.0D0-100.0D0)/50.0D0)
ELSE
    FSTART = TWOTWO ** NINT((ZSTART-100.0D0)/50.0D0)
ENDIF
```

Initial position

```
ZEND = ZTRACK(NTRACK)
```

Final position

\* Similarly calculate the FEND from ZEND

...

```
FIMP = FEND / FSTART
```

Importance Ratio

# usimbs.f: Function example

- Introduce an importance biasing assuming an exponential law of attenuation in the R direction  $\exp(-\lambda \times R)$ , for  $R > 1\text{cm}$

```
RSTART = SQRT(XTRACK(0)**2 + YTRACK(0)**2 + ZTRACK(0)**2)
```

```
REND = SQRT(XTRACK(NTRACK)**2 + YTRACK(NTRACK)**2 +  
            ZTRACK(NTRACK)**2)
```

```
IF (RSTART .LT. ONEONE) THEN
```

```
    FSTART = ONEONE
```

```
ELSE
```

```
    FSTART = EXP(-ALAMBDA * RSTART)
```

```
ENDIF
```

```
IF (REND .LT. ONEONE) THEN
```

```
    FEND = ONEONE
```

```
ELSE
```

```
    FEND = EXP(-ALAMBDA * REND)
```

```
ENDIF
```

```
FIMP = FEND / FSTART
```

# usimbs.f: Sampling from array [1/4]

- Create a concentric cylindrical biasing with the weights sample from an array with various radii

- Define the variables:

```
PARAMETER (NBIAS=5)
```

```
PARAMETER (Xcenter=ZERZER)
```

```
PARAMETER (Ycenter=ZERZER)
```

```
DOUBLE PRECISION BIASR(NBIAS), BIASF(NBIAS)
```

\* Radius

```
DATA BIASR / 200.0, 250.0, 300.0, 400.0, 500.0 /
```

\* Biasing factor

```
DATA BIASF / NBIAS * 2.0 /
```

```
LOGICAL LFIRST
```

```
DATA LFIRST / .TRUE. /
```

```
SAVE LFIRST, BIASR, BIASF
```

# usimbs.f: Sampling from array [2/4]

- Initialization – build the cumulative importance factor

```
IF (LFIRST) THEN
```

```
  WRITE(LUNOUT,*) "*** User defined biasing ***"
```

```
  PREVBIAS = 1.0
```

```
  DO N=1,NBIAS
```

```
    PREVBIAS = PREVBIAS * BIASF(N)
```

```
    BIASF(N) = PREVBIAS
```

```
    WRITE(LUNOUT,*) "Bias cylinder: ",N, BIASR(N), BIASF(N)
```

```
*      convert the radius to square to avoid sqrt
```

```
    BIASR(N) = BIASR(N)**2
```

```
  ENDDO
```

```
  LFIRST = .FALSE.
```

```
ENDIF
```

# usimbs.f: Sampling from array [3/4]

- Calculate the importance biasing

- \* Find square of radius for starting/ending position

$$Rold = (Xtrack(0)-Xcenter)**2 + (Ytrack(0)-Ycenter)**2$$

$$Rnew = (Xtrack(Ntrack)-Xcenter)**2 + (Ytrack(Ntrack)-Ycenter)**2$$

- \* Search index of the starting position

$$Nold = NBinSearch(Rold, NBIAS, BIASR)$$

IF (Nold.EQ.0) THEN

$$BIASOLD = 1.0$$

ELSE

$$BIASOLD = BIASF(Nold)$$

ENDIF

- \* Search index of the new position

$$Nnew = NBinSearch(Rnew, NBIAS, BIASR)$$

IF (Nnew.EQ.0) THEN

$$BIASNEW = 1.0$$

ELSE

$$BIASNEW = BIASF(Nnew)$$

ENDIF

$$FIMP = BIASNEW / BIASOLD$$

# usimbs.f: Sampling from array [4/4]

Perform a binary search  
Converge in  $\log_2(N)$  steps

```
INTEGER FUNCTION NBinSearch(x, N, VEC)
INCLUDE '(DBLPRC)'
DOUBLE PRECISION VEC(N)
NLOW = 1
NHIGH = N
NBinSearch = 0
IF (X.GE.VEC(NLOW) .AND. X.LE.VEC(NHIGH)) THEN
10    CONTINUE
      MID = (NLOW+NHIGH)/2
      IF (MID.EQ.NLOW) THEN
          NBinSearch = MID
          RETURN
      ELSEIF (X .GT. VEC(MID)) THEN
          NLOW = MID
      ELSEIF (X .LT. VEC(MID)) THEN
          HIGH = MID
      ELSE
          NBinSearch = MID
          RETURN
      ENDIF
      GOTO 10
END IF
END
```

# Neutrino Decay Biasing

- There is a special routine `udcdrl.f` where one can bias the direction of the emitted neutrino in decays.

DOUBLE PRECISION **FUNCTION** `UDCDRL`( IJ, KPB, NDCY, UDCDRB, VDCDRB, WDCDRB)

Input variables:

IJ                   decaying particle  
KPB                  outgoing neutrino

Output variables:

U,V,W DCDRB       preferential outgoing direction for the neutrino  
UDCDRL            Lambda for direction biasing (1-cos( $\theta$ ))

The biasing expression is of the form:

$$e^{-(1-\cos\theta/\lambda)}$$

- Useful for neutrino applications like CNGS, Beta Beams...
- For a fixed direction the `LAM-BIAS` card with `SDUM=DCY-DIRE` could be used instead

# Direction Biasing Example (n\_TOF)

- Bias the direction of the neutrino in the pion decay so the daughter muon to be directed to the exp. area @(-120,0,18500). The direction is given in the lab-frame, and in this example the energies we are dealing are small, so safely we can assume that also in the lab-frame, the neutrino and muon go in opposite directions. The lambda is 1/4 wide enough to cover the whole exp area.

```
INCLUDE '(TRACKR)'  
*  
UDCDRB = XTRACK (NTRACK) + 120.D+00  
VDCDRB = YTRACK (NTRACK) - ZERZER  
WDCDRB = ZTRACK (NTRACK) - 18500.D+00  
HLPHLP = SQRT ( UDCDRB**2 + VDCDRB**2 + WDCDRB**2 )  
UDCDRB = UDCDRB / HLPHLP  
VDCDRB = VDCDRB / HLPHLP  
WDCDRB = WDCDRB / HLPHLP  
UDCDRL = ONEFOU  
*
```

Direction in the lab frame

Width [1-cos( $\theta$ )]