

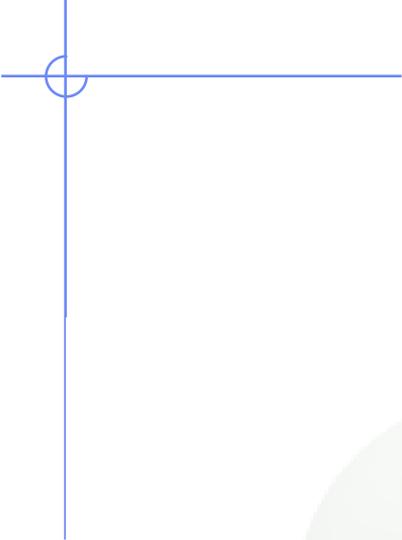


# FLUKA Advanced Scoring

FLUKA Advanced Course

# Built-In and User Scoring

- Several **pre-defined estimators** can be activated in FLUKA.
- One usually refers to these estimators as **"scoring"** capabilities
- Users have also the possibility to build their own scoring through user routines (some of which mentioned afterwards), HOWEVER:
  - **Built-in scoring** covers most of the **common needs**
  - **Built-in scoring** has been **extensively tested**
  - **Built-in scoring** takes BIASING **weights automatically into account**
  - **Built-in scoring** has **refined algorithms** for track subdivision
  - **Built-in scoring** comes with **utility programs** that allow to evaluate statistical errors
- Scoring can be geometry dependent AND/OR geometry independent  
FLUKA can score **particle fluences, current, track length, energy spectra, particle spectra, energy deposition...**
- Either integrated over the **"run"**, with proper normalization, OR **event-by event**
- Standard scoring can be weighted by means of **simple user routines**



# A Reminder on Flux/Fluence/Current

# Reaction Rate and Cross Section [1/3]

- We call **mean free path**  $\lambda[cm]$  the average distance travelled by a particle in a material before an interaction. Its inverse,  $\Sigma [cm^{-1}]$  is the probability of interaction per unit distance, and is called **macroscopic cross section**. Both  $\lambda$  and  $\Sigma$  depend on the material and on the particle type and energy.
- For  $N$  identical particles, the number of reactions  $R$  occurring in a given time interval will be equal to the total distance travelled  $l$  times the probability per unit distance  $\Sigma$ :  $R = l\Sigma$
- The reaction rate will be  $\dot{R} = dl/dt \Sigma = v\Sigma$ , where  $v$  is the average particle velocity.

# Reaction Rate and Cross Section [2/3]

- Assume now that  $n(\mathbf{r}, v) = dN/dV$  [ $cm^{-3}$ ] be the density of particles with velocity  $v = dl/dt$  [ $cm/s$ ], at a spatial position  $\mathbf{r}$ . The reaction rate inside the volume element  $dV$  will be:  $d\dot{R}/dV = n(\mathbf{r}, v)v\Sigma$
- The quantity  $\dot{\Phi}(\mathbf{r}, v) = n(\mathbf{r}, v)v$  is called **fluence rate** or **flux density** and has dimensions [ $cm^{-3} cm t^{-1}$ ] = [ $cm^{-2} t^{-1}$ ].
- The time integral of the flux density  $\Phi(\mathbf{r}, v) = n(\mathbf{r}, v)dl$  is the **fluence** [ $cm^{-2}$ ]
- Fluence is measured in **particles per  $cm^2$**  but in reality it describes the **density of particle tracks**
- The number of reactions inside a volume  $V$  is given by the formula:  $R = \Sigma\Phi V$  (where both  $\Sigma$  and  $\Phi$  are integrated over energy or velocity)

# Reaction Rate and Cross Section [3/3]

- Dividing the macroscopic cross section by  $N_0$ , the number of atoms per unit volume, one obtains the **microscopic cross section**  $\sigma$  [*barn* =  $10^{-24} \text{cm}^2$ ].

$$\frac{\text{probability/cm}}{\text{atoms/cm}^3} = \frac{\text{probability} \times \text{cm}^2}{\text{atom}} = \text{atom effective area}$$

- i.e., the **area of an atom weighted with the probability of interaction** (hence the name "cross section").
- But it can also be understood as the **probability of interaction per unit length, with the length measured in atoms/cm<sup>2</sup>** (the number of atoms contained in a cylinder with a 1 cm<sup>2</sup> base).
- In this way, both microscopic and macroscopic cross section are shown to have a similar physical meaning of "probability of interaction per unit length", with length measured in different units. Thus, the number of interaction can be obtained by both by multiplying by the corresponding particle track-length.

# Fluence estimation [1/2]

- Track length estimation:

$$\dot{\Phi}(v) dt = n(v) v dt = \frac{dN(v)}{dV} \frac{dl(v)}{dt} dt = \lim_{\Delta V \rightarrow 0} \frac{\sum_i l_i(v)}{\Delta V}$$

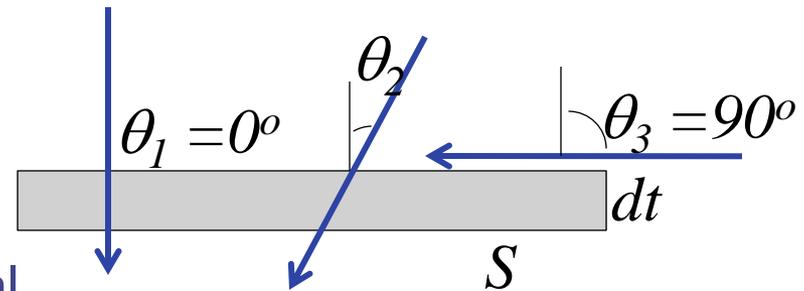
- Collision density estimation:

$$\dot{\Phi}(v) = \frac{\dot{R}(v)}{\sigma(v) N_o} = \frac{\dot{R}(v)}{\Sigma(v)} = \dot{R}(v) \lambda(v)$$

# Fluence estimation [2/2]

## Surface crossing estimation

- Imagine a surface having an infinitesimal thickness  $dt$ . A particle incident with an angle  $\theta$  with respect to the normal of the surface  $S$  will travel a segment  $dt/\cos\theta$ .



- Therefore, we can calculate an average surface fluence by adding  $dt/\cos\theta$  for each particle crossing the surface, and dividing by the volume  $S dt$

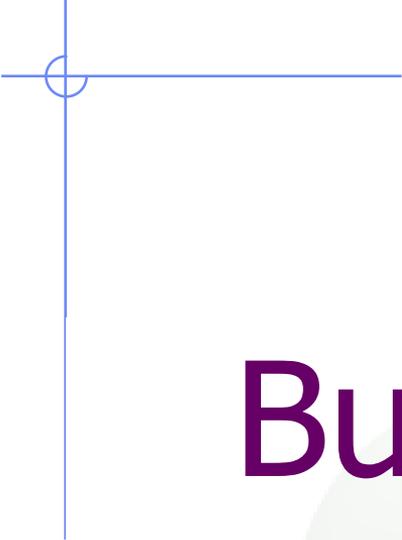
$$\Phi = \lim_{dt \rightarrow 0} \frac{\sum_i \frac{dt}{\cos \theta_i}}{S dt}$$

- While the **current**  $J$  will be to count the number of particles crossing the surface divided by the surface

$$J = dN/dS$$

The **fluence is independent** from the orientation of **surface**  $S$ ,  
while the **current is NOT!**

In an isotropic field can be easily seen that on a flat surface  $J = \Phi/2$



# Built-in Conversions and AUXSCORE



# Scoring Cards (see Beginner's Course)

- **SCORE** scores energy deposited (or star density) in all regions
- **USRTRACK**, **USRCOLL** score average  $d\Phi/dE$  (differential fluence) of a given type or family of particles in a given region
- **USRBDX** scores average  $d^2\Phi/dEd\Omega$  (double-differential fluence or current) of a given type or family of particles on a given surface
- **USRBIN** scores the spatial distribution of energy deposited, or total fluence (or star density, or momentum transfer) in a regular mesh (cylindrical or Cartesian) described by the user
- **USRYIELD** scores a double differential yield of particles escaping from a surface. The distribution can be with respect to energy and angle, but also other more "exotic" quantities
- **RESNUCLEi** scores residual nuclei in a given region

# Scoring cards

(most explained in the following slides)

- **EVENTBIN** is like **USRBIN**, but prints the binning output **after each event** instead of an average over histories
- **ROTPRBIN** sets the **storage precision** (single or double) and assigns **rotations/translations** for a given user-defined binning (**USRBIN** or **EVENTBIN**)
  - linked to lecture about the use of **LATTICE**
- **USERDUMP** defines the events to be written onto a **“collision tape”** file
  - see **mgdraw** part later in this lecture
- **AUXSCORE** defines **filters** and **conversion coefficients**
- **TCQUENCH** sets scoring **time cut-offs** and/or **Birks quenching** parameters for binnings (**USRBIN** or **EVENTBIN**) indicated by the user
- **DETECT** scores **energy deposition in coincidence** or anti-coincidence with a trigger, separately for each “event” (primary history)
  - dedicated post-processing routine is now available

# Lattice Related Scorings

**EVENTBIN** or **USRBIN** with **WHAT(1)=8** :

Special user-defined 3D binning. Two variables are discontinuous (*e.g.*, region number), the third one is continuous, but not necessarily a space coordinate.

Variable	Type	Default	Override Routine
1 <sup>st</sup>	integer	region number	MUSRBR
2 <sup>nd</sup>	integer	lattice cell number	LUSRBL
3 <sup>rd</sup>	float	Before used as $\eta$ , now set to zero*	FUSRBV

\* In the past it scored:  $n = -\ln(\tan(0.5 \arctan(\sqrt{x^2 + y^2})/z))$

**ROTPRBIN** can assign rotations/translations (as defined by ROTDEFI) for a given user-defined binning (**USRBIN** or **EVENTBIN**):

- this allows *e.g.*, defining a 'normal' scoring around a prototype and then 'replicating' the scoring to the respective lattices

# "FILTER" : AUXSCORE

**WARNING!!**  
all energy deposition  
ionization+NIEL by the selected  
particle

There is the possibility to filter the estimators, restricting the scoring to a selected subset of particles.

For instance: USRBIN energy deposition by muons only

USRBIN	11.0	ENERGY	-40.0	10.0	15.0	TargEne
USRBIN	0.0		-5.0	100.0	200.0	&
AUXSCORE	USRBIN	MUONS		TargEne	TargEne	

Assign the "muons" filter to the USRBIN estimator named TargEne

Another example: score the yield of 56-Iron ions (very useful: there is no separate name for each ion specie, except light ones. HEAVYION score all isotopes heavier than alpha's together!)

USRYIELD	124.0	ALL-PART	-87.	TARGS3	INAIR	1.0	Fe56
USRYIELD	180.0	0.0	18.	10.0	0.0	3.0	&
AUXSCORE	USRYIELD	-5602600.		Fe56	Fe56		

The requested ion is coded in what(2) according to its **A**, **Z** and (optionally) isomeric state **m**:

$\text{what}(2) = - (100 * \mathbf{Z} + 100000 * \mathbf{A} + \mathbf{m} * 100000000)$

with 0==all , i.e. 2600 == all Iron isotopes

# Dose-Equivalent (not Dose)

For some quantities, there is the possibility to get built-in conversions, without the need for user routines: done with generalized particles, the most commonly used is dose equivalent (ambient dose equivalent or effective dose):

**DOSE-EQ**                      Dose Equivalent [pSv]

**DOSEQLET**                  Dose Equivalent LET (Q(LET) according to ICRP60)

*!!!! Different to !!!!:*

( **DOSE**                      total absorbed dose in (obviously...) GeV/g!  
**DOSE-EM**                dose as above (electromagnetic part only) )

The set of conversion coefficients used to calculate DOSE-EQ can be selected by the user among a list (see manual) with AUXSCORE:

USRBIN	11.0	DOSE-EQ	-40.0	10.0	15.0	TargDEQ
USRBIN	0.0		-5.0	100.0	200.0	&
AUXSCORE	USRBIN			TargDEQ	TargDEQ	AMB74

Scores equivalent dose by folding the particle fluences with the "AMB74" conversion coefficients

**WARNING : DOSE-EQ no coefficients available for heavy ions (ok for DOSEQLET) !!!**

# Available Conversion Coefficients

The following dose conversion coefficients sets are available:

- 1) Effective dose sets from ICRP74 and Pelliccioni data calculated with ICRP radiation weighting factors  $W_r$ 
  - (a) **EAP74** : Anterior-Posterior irradiation
  - (b) **ERT74** : Rotational irradiation geometry
  - (c) **EWT74** : WORST possible geometry for the irradiation
- 2) Effective dose sets from ICRP74 and Pelliccioni data calculated with the Pelliccioni radiation weighting factors  $W_r$ 
  - (a) **EAPMP** : Anterior-Posterior irradiation
  - (b) **ERTMP** : Rotational irradiation geometry
  - (c) **EWTMP** : WORST possible geometry for the irradiation
- 3) Ambient dose equivalent from ICRP74 and Pelliccioni data
  - (a) **AMB74** : **[Default]**
- 4) Ambient dose equivalent with old "GRS"-conversion factors
  - (a) **AMBGS**

**(see backup slides for details)**

# Radiation Damage to Electronics

- All important quantities to estimate risks of damage to electronics can be directly scored in FLUKA (*see Materials lecture*):

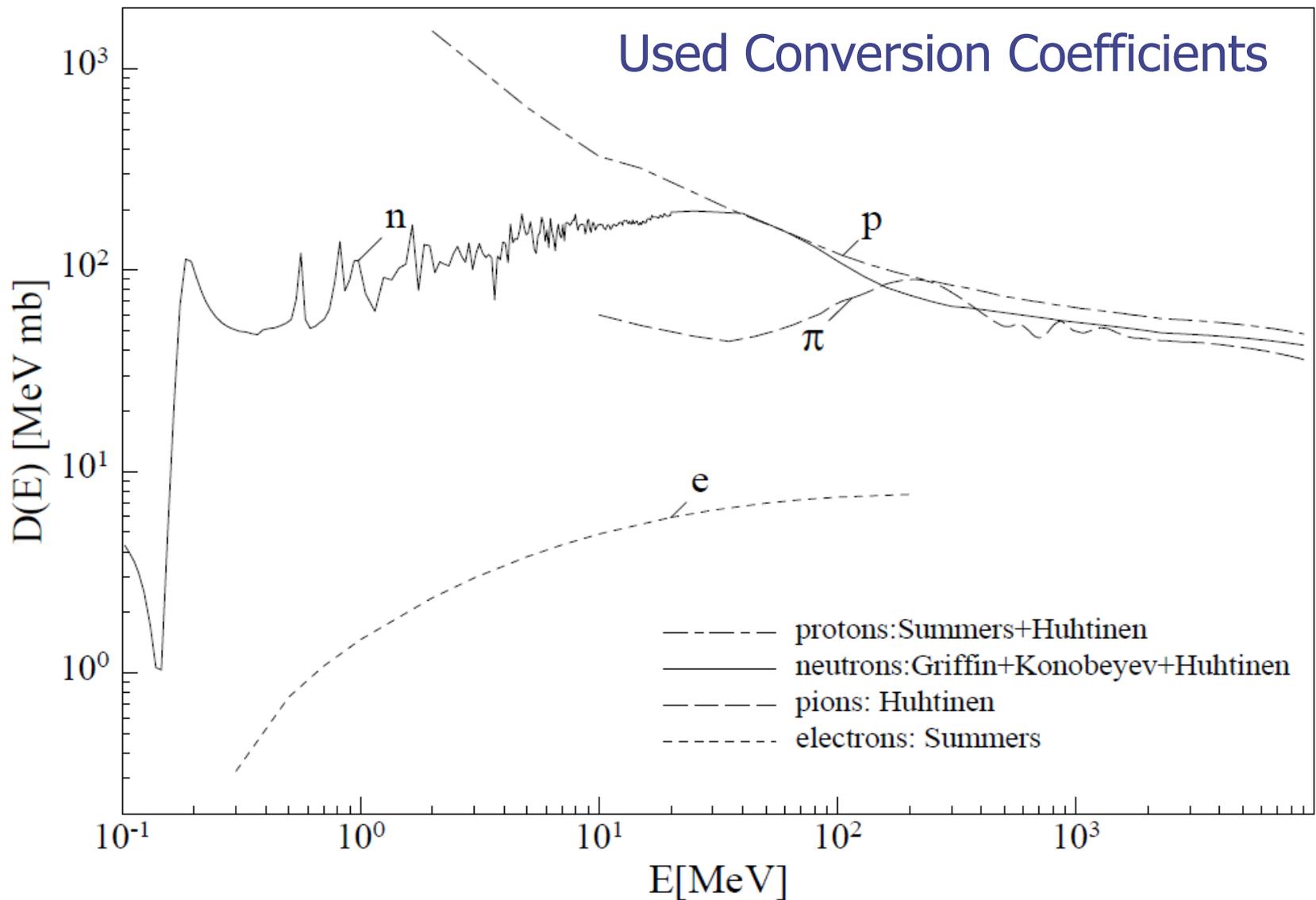
## **Cumulative damage:**

- Energy deposition (total dose) by scoring DOSE with any 'energy deposition like estimator' (*e.g.*, USRBIN)
- Si Lattice displacement (1-MeV neutron equivalent particle fluxes) with any 'fluence like estimator' (*e.g.*, USRTRACK)

## **Stochastic failures (SEU):**

- "high" energy hadron fluences ("E>20 MeV") with any 'fluence like estimator' (*e.g.*, USRTRACK)  
**(the option of special threshold functions [user defined] is currently in development and will be included in the next release together with the scoring related to the "damage by thermal neutrons")**
- The powerful FLUKA scoring options together with the analysis of particle energy spectra allows a detailed study in order to select best possible locations for electronics or efficiently design shielding implementations

# 1MeV Neutron Equivalent



# Electronic Damage - Related Scoring

<b>DOSE</b>	total absorbed dose in (obviously...)	GeV/g!
<b>SI1MEVNE</b>	Silicon 1 MeV-neutron equivalent fluence	
<b>HADGT20M</b>	Hadrons fluence with energy > 20 MeV	

- **USRTRACK** scores average  $d\Phi/dE$  (differential fluence) in a given region (SI1MEVNE, HADGT20M or any particle type)
- **USRBDX** scores for the same quantities average  $d^2\Phi/dEd\Omega$  (double-differential fluence or current) on a given surface (between two regions)
- **USRBIN** scores the spatial distribution either of deposited dose, or fluence (1MeV or 20MeV) in a regular mesh (cylindrical or Cartesian) described by the user
- **USRBIN** also scores the same quantities on a region basis

## \* 1) high-energy hadron fluence spectrum

```
USRTRACK      -1.  HADGT20M      -31.  RADMON1      125.  170.Ust20MeV
USRTRACK      1D3    1D-14
&
```

## \* 2) displacement damage spectrum

```
USRBDX       98.  SI1MEVNE      -41.  TAIR  RADMON1  150.Usx1MeV
USRBDX       1D3    1D-14      170.  &
```

## \* 3) dose distribution in a regular mesh through the geometry

```
USRBIN       10.  DOSE          -21.  100.  20.  200.UsbDose
USRBIN      -100. -20.  -100.  100.  20.  150.&
```

## \* 4) integrated high-energy hadron fluence on a region basis

```
USRBIN       18.0 HADGT20M      -37.0  LSTREG  300.0  10000.0UsbReg20
USRBIN      FSTREG  0.0 -10000.0  1.0  1.0  1.0  &
```

# Displacement Damage / Charge Collection

For all charged particles and Heavy Ions FLUKA calculates the recoil as a normal particle. During transport it calculates the restricted and unrestricted nuclear stopping power, allowing to score dpa's and non-ionizing energy loss (NIEL):

**NIEL-DEP**      Non Ionizing Energy Loss deposition

**DPA-SCO**      Displacements per atoms

(details see **Material's Lecture**)

In addition (not necessarily linked to displacement damage) the following can be useful in order to get the net charge deposition in a given region:

**NET-CHRG**      Net Charge

# Activation Scoring (Reminder!)

Input card: **RADDECAY**

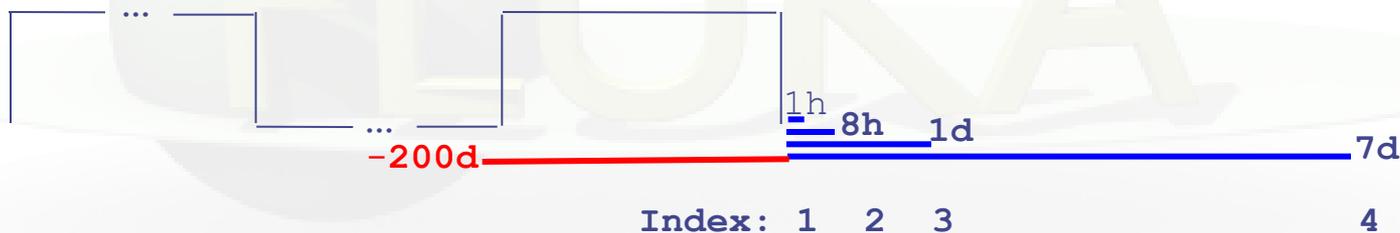
requests simulation of decay of produced radioactive nuclides and allows to modify biasing and transport thresholds (defined with other cards) for the transport of decay radiation

Input card: **IRRPROFI**

definition of an irradiation profile (irradiation times and intensities)

Input card: **DCYTIMES**

definition of decay (cooling ) times



Input card: **DCYSCORE**

associates scoring detectors (radio-nuclides, fluence, dose) with different cooling times

Input card: **AUXSCORE**

allows to associate scoring estimators with dose equivalent conversion factors or/and to filter them according to (generalized) particle identity

# Before: Change Of 'Geometry'

## -> Two-Step Method

Two separate FLUKA simulations were necessary in order to change the geometry between the prompt and the decay part (see *e.g.*, CERN FLUKA Course for explanation of use [S. Roesler]):

### 1<sup>st</sup> step

- simulation of production of radioactive nuclides and of their build-up and decay for a certain irradiation pattern and different cooling times
- write-out of all information on produced radio-nuclides at each cooling time into external file via user-routine **usrrnc.f**
- uses the analytical solution of the Bateman equation in FLUKA (*i.e.*, radioactive build-up and decay identical to 1-step method)

### 2<sup>nd</sup> step

- simulation(s) of radioactive decay and transport of decay radiation
- information on radio-nuclides read in from file created in 1<sup>st</sup> step via user-routine **source.f**
- individual simulations for each requested cooling time

# Now: Change Of 'Geometry'

The latest FLUKA version (this course) contains the possibility of selectively changing regions to vacuum/blackhole and/or switching on/off possible fields) when transporting radioactive decay products. Radioactive decay products originating from regions switched to vacuum/blackhole are ignored. This is helpful for situations where the emissions of an activated object in a complex environment have to be evaluated standalone.

Through Input card: **ASSIGNMA**

(a (single-element or compound) material is assigned to each geometry region)

## Example

```
*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...  
ASSIGNMA      GOLD      REG1      REG2      1.0      0.0
```

MATERIAL from REGION to REGION in steps of

**WHAT(5) = 1/2/3:** a magnetic/electric/both field(s) is(are) present in the region(s) defined by WHAT(2), (3), and (4), for **both prompt and radioactive decay products**

= **4/5/6:** same as above, but for **prompt products only**

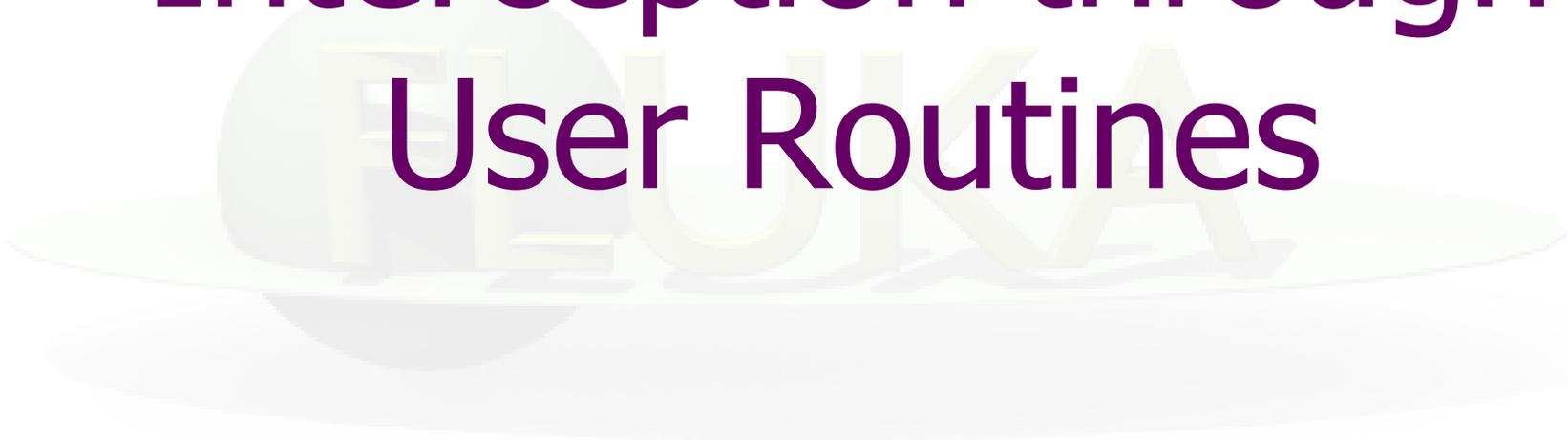
= **7/8/9:** same as above, but for **radioactive decay products only**

**Electr/magn. field is present, either in both/prompt or decay part (material set to vacuum in the other case!**

*Note: so far distinction between lattices (all instances are affected)*



# Interception through User Routines



# Routines Linked to the FLUKA Output

## in association to FLUKA output

- `comscw.f` ... weighting energy deposition or star production
- `fluscw.f` ... weighting fluence, current and yield
- `mgdraw.f` ... general scoring tracking interface
- `usrrnc.f` ... isotope production and possible conversion
- `endscp.f` ...
- `fldscp.f` ...
- `musrbr.f` ...
- `lusrbl.f` ...
- `fusrbv.f` ...

### Possibly Related: (Intercepting Particle Stack)

- `mdstck.f`
- `stupre.f`
- `stuprf.f`

**See Lecture  
User-Routines**

# comscw.f

## (weighting energy deposition or star production)

Argument list (all variables are input only)

```
IJ      : particle type (1 = proton, 8 = neutron, etc.: see code in 5.1)
XA, YA, ZA : current particle position
MREG    : current geometry region
RULL    : amount to be deposited (unweighted)
LLO     : particle generation
ICALL   : internal code calling flag (not for general use)
```

Activated by option **USERWEIG** with **WHAT(6) > 0.0**. Energy and stars obtained via **SCORE**, **USRBIN** and **EVENTBIN**, and production of residual nuclei obtained via **RESNUCLEi** are multiplied by the value returned by this function. The user can implement any desired logic according to the argument list (particle type, position, region, amount deposited, particle generation), or information available in **COMMON SCOHLP** (binning number, type of scored quantity). The scored quantity is given by the flag **ISCRNG** (in **SCOHLP**):

**ISCRNG = 1** → Energy density binning

**ISCRNG = 3** → Residual nuclei scoring

**ISCRNG = 5** → Activity density binning

**ISCRNG = 2** → Star density binning

**ISCRNG = 4** → Momentum transfer

**ISCRNG = 6** → Net charge density

# comscw.f – Tips & Tricks

The binning/detector number is given by **JSCRNG** (in **SCOHLP**) and is printed in output between the estimator type and the detector name.

Note that a detector of residual nuclei can have the same **JSCRNG** number as a binning (use the value of **ISCRNG** to discriminate).

Further information can be obtained including **COMMON TRACKR** (for instance particle's total energy, direction cosines, age). **TRACKR** contains also special user variables (both integer and in double precision) which can be used to save information about particles which have undergone some particular event.

If data concerning the current material are needed, it can be accessed as **MEDIUM(MREG)** if **(FLKMAT)** is included.

# fluscw.f

## (weighting fluence, current and yield)

Argument list (all variables are input only)	
IJ	: particle type
PLA	: particle momentum (if > 0.0) or -PLA = kinetic energy (if < 0.0)
TXX, TYY, TZZ	: particle current direction cosines
WEE	: particle weight
XX, YY, ZZ	: particle position
NRGFLK	: current region (after boundary crossing)
IOLREG	: previous region (before boundary crossing). Useful only with boundary crossing estimators (for other estimators it has no meaning)
LLO	: particle generation
NSURF	: internal code calling flag (not for general use)

Similar to **COMSCW**. Function **FLUSCW** is activated by option **USERWEIG**, with **WHAT(3) > 0.0**. Yields obtained via **USRyield**, fluences calculated with **USRBDX**, **USRTRACK**, **USRCOLL**, **USRBIN**, and currents calculated with **USRBDX** are multiplied by the value returned by this function.

# fluscw.f - Tips & Tricks

The user can implement any desired logic according to the argument list (particle type, energy, direction, weight, position, region, boundary, particle generation), or information available in **COMMON SCOHLP** (binning or detector number, estimator type). The estimator type is given by the flag **ISCRNG** (in **COMMON SCOHLP**):

**ISCRNG** = 1 → Boundary crossing estimator

**ISCRNG** = 2 → Track-length binning

**ISCRNG** = 3 → Track-length estimator

**ISCRNG** = 4 → Collision density estimator

**ISCRNG** = 5 → Yield estimator

# Useful (*e.g.*, in fluscw/comscw)

When interception of regions is required, thus the conversion of **region name to number** (or opposite, but rarely) might be important:

```
CALL GEON2R ( REGNAM, NREG, IERR )
```

Input variable:

Regnam = region name (CHAR\*8)

Output variables:

Nreg = region number

Ierr = error code (0 on success, 1 on failure)

## Conversion of **region number to name**

```
CALL GEOR2N ( NREG, REGNAM, IERR )
```

Input variable:

Nreg = region number

Output variables:

Regname = region name (CHAR\*8)

Ierr = error code (0 on success, 1 on failure)

# mgdraw.f

## (General interface to FLUKA transport and scoring)

```
Argument list (all variables are input only)
ICODE : FLUKA physical compartment originating the call
        = 1: call from subroutine KASKAD (hadrons and muons)
        = 2: call from subroutine EMFSCO (e-, e+ and photons)
        = 3: call from subroutine KASNEU (low-energy neutrons)
        = 4: call from subroutine KASHEA (heavy ions)
        = 5: call from subroutine KASOPH (optical photons)
MREG  : current region
```

See Lecture  
Code-Structure

Subroutine **MGDRAW**, activated by option **USERDUMP** with **WHAT(1) ≥ 100.0**, usually writes a "collision tape", *i.e.*, a file where all or selected transport events are recorded. The default version (unmodified by the user) offers several possibilities, selected by **WHAT(3)** in **USERDUMP**.

# mgdraw.f

The different **ENTRY** points of **MGDRAW**

Additional flexibility is offered by a user entry **USDRAW**, interfaced with the most important physical events happening during particle transport.

The user can modify also any other entry of this subroutine:

**BXDRAW** called at boundary crossings,

**EEDRAW** called at event end,

**MGDRAW** called at each step, for trajectory drawing and dE/dx energy deposition events,

**ENDRAW** for recording of point energy deposition events,

**SODRAW** for recording of source events

# mgdraw.f: the SODRAW entry

## Argument list

No arguments

**SODRAW** writes by default, for each source or beam particle:

- NCASE:** (in **COMMON CASLIM**, with a minus sign to identify SODRAW output) number of primaries followed so far
- NPFLKA:** (in **COMMON FLKSTK**) stack pointer
- NSTMAX:** (in **COMMON FLKSTK**) highest value of the stack pointer encountered so far
- TKESUM:** (in **COMMON SOURCM**) total kinetic energy of the primaries of a user written source, if applicable. Otherwise = 0.0
- WEIPRI:** (in **COMMON SUMCOU**) total weight of the primaries handled so far

NPFLKA times: (all variables in COMMON FLKSTK)	ILOFLK: TKEFLK + AM: WTFK: XFLK, YFLK, ZFLK: TXFLK, TYFLK, TZFLK:	type of source particle total particle energy (kinetic+mass) source particle weight source particle position source particle direction cosines
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# mgdraw.f: the MGDRAW entry

- MTRACK:** number of energy deposition events along the track
- JTRACK:** type of particle
- ETRACK:** total energy of the particle
- WTRACK:** weight of the particle
- NTRACK:** values of **XTRACK**, **YTRACK**, **ZTRACK**: end of each track segment
- MTRACK:** values of **DTRACK**: energy deposited at each deposition event
- CTRACK:** total length of the curved path

Other variables are available in **TRACKR** (but not written by **MGDRAW** unless the latter is modified by the user: particle momentum, direction cosines, cosines of the polarisation vector, age, generation, etc. see a full list in the comment in the **INCLUDE** file).

# mgdraw.f: the ENDRAW entry

Called at  
point-like Energy  
Deposition dumps

(for example:  
stopping particles,  
photoelectric eff.,  
etc.)

## Argument list (all variables are input only)

```
ICODE : type of event originating energy deposition
ICODE = 1x: call from subroutine KASKAD (hadrons and muons);
          = 10: elastic interaction recoil
          = 11: inelastic interaction recoil
          = 12: stopping particle
          = 14: particle escaping (energy deposited in blackhole)
ICODE = 2x: call from subroutine EMFSCO (electrons, positrons and photons)
          = 20: local energy deposition (i.e. photoelectric)
          = 21 or 22: particle below threshold
          = 23: particle escaping (energy deposited in blackhole)
ICODE = 3x: call from subroutine KASNEU (low-energy neutrons)
          = 30: target recoil
          = 31: neutron below threshold
          = 32: neutron escaping (energy deposited in blackhole)
ICODE = 4x: call from subroutine KASHEA (heavy ions)
          = 40: ion escaping (energy deposited in blackhole)
ICODE = 5x: call from subroutine KASOPH (optical photons)
          = 50: optical photon absorption
          = 51: optical photon escaping (energy deposited in blackhole)
MREG   : current region
RULL   : energy amount deposited
XSCO, YSCO, ZSCO : point where energy is deposited
```

# mgdraw.f: the BXDRAW entry

Called at Boundary Crossings

Argument list (all variables are input only)

ICODE : physical compartment originating the call, as in the MGDRAW entry  
MREG : region from which the particle is exiting  
NEWREG : region the particle is entering  
XSCO, YSCO, ZSCO : point where the boundary crossing occurs

# mgdraw.f: the USDRAW entry

USDRAW is called  
after each  
particle interaction  
(requested by the  
user with option  
USERDUMP,  
WHAT(4) ≥ 1.0)

```
Argument list (all variables are input only)
ICCODE : type of event
ICCODE = 10x: call from subroutine KASKAD (hadron and muon interactions);
           = 100: elastic interaction secondaries
           = 101: inelastic interaction secondaries
           = 102: particle decay secondaries
           = 103: delta ray generation secondaries
           = 104: pair production secondaries
           = 105: bremsstrahlung secondaries
ICCODE = 20x: call from subroutine EMFSCO (electron, positron and photon interactions)
           = 208: bremsstrahlung secondaries
           = 210: Møller secondaries
           = 212: Bhabha secondaries
           = 214: in-flight annihilation secondaries
           = 215: annihilation at rest secondaries
           = 217: pair production secondaries
           = 219: Compton scattering secondaries
           = 221: photoelectric secondaries
           = 225: Rayleigh scattering secondaries
ICCODE = 30x: call from subroutine KASNEU (low-energy neutron interactions)
           = 300: neutron interaction secondaries
ICCODE = 40x: call from subroutine KASHEA (heavy ion interactions)
           = 400: delta ray generation secondaries

MREG   : current region
XSCO, YSCO, ZSCO : interaction point
```

# mgdraw.f: the EEDRAW entry

Called at Event End

Argument list (all variables are input only)

**ICODE** : physical compartment originating the call, as in the MGDRAW entry

# mgdraw.f - Tips & Tricks

- An interesting aspect of the routine is that the six entries (all of which, if desired, can be activated at the same time by setting **USERDUMP** with **WHAT(3) = 0.0** and **WHAT(4) ≥ 1.0**) constitute a complete interface to the entire Fluka transport.
  - Therefore, MGDRAW can be used not only to write a collision tape, but to do any kind of complex analysis. (*e.g.*, event by event output as used in HEP applications).
- When mgdraw should better not be used
  - When biasing is requested (non-analogue run)
  - Whenever low-energy neutrons ( $E < 20$  MeV) are used, unless one has a deep knowledge of the peculiarities of their transport and quantities (*i.e.*, kerma, etc)



# (User) Bugs & Problems (Reminder)

# Scoring: USRBIN/EVENTBIN

## Error Message:

\*\*\* Activity/fission/neutron balance binnings cannot be track-length!!!

USRBIN scoring method:

- Track-length quantities: that can be distributed along a track, e.g. fluence, energy deposition...  
**WHAT(1) >= 10**
- Point-wise quantities: that have to be scored on a point, or in the middle of the step!  
e.g. Activity, Fission, Neutron balance...  
**WHAT(1) < 10**

# Other Error in Scoring Definitions

- Example: badly defined USRBIN limits

## Error Message:

```
***** Fluka stopped in Usrbin: "usr/eventbin" n. 1 *****  
***** with zero width 0.000 for axis R *****
```

- Never use unit numbers smaller than 20 or higher than 99  
<20 They are reserved by FLUKA  
>99 Depends on fortran
- **Never mix** the output of different scoring cards in the same unit

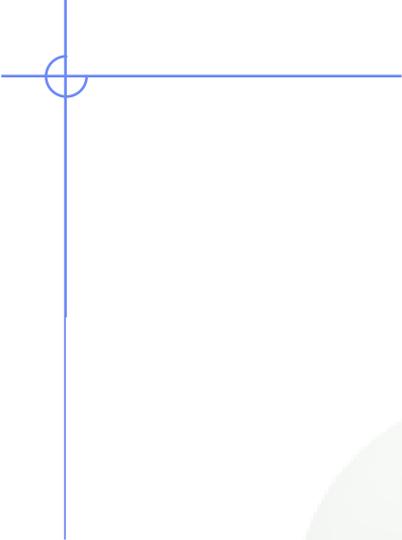
**automatically checked by FLAIR**

# Merging Cycles from Different Inputs

## Error Message:

none

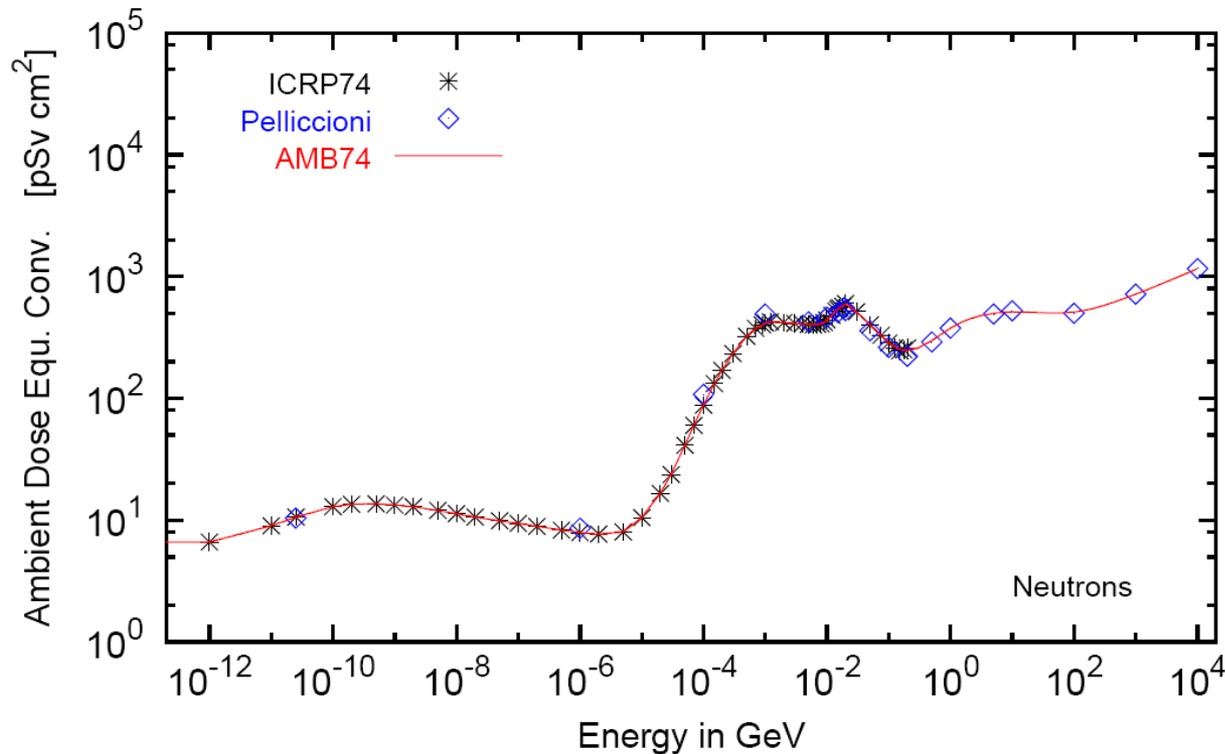
- Verify that you didn't merge cycles from different runs that the input has been modified.
- It's a good habit to clean the files before starting a new run with a modified input file.
- Flair offers this possibility from the "Output Files" frame.
- It's good to develop the habit to clean the output files from test runs.
- ....change the name of the input file for every new problem!



# Backup

# Conversion Coefficients

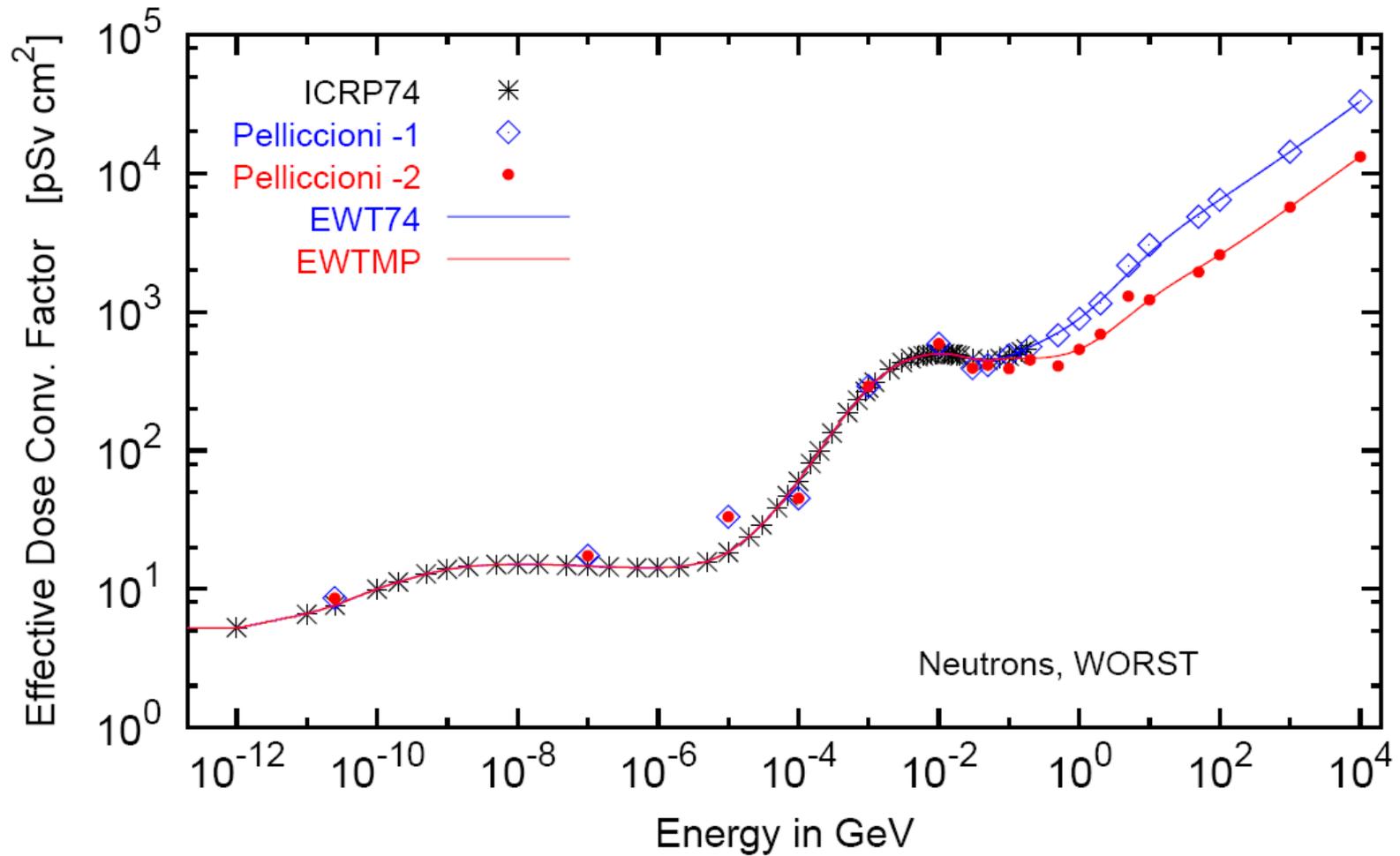
Conversion coefficients from fluence to ambient dose equivalent are based on ICRP74 values and values calculated by M.Pelliccioni. They are implemented for **protons, neutrons, charged pions, muons, photons, electrons** (conversion coefficients for other particles are approximated by these). AMB74 is the default choice for dose equivalent calculation.



# Fluence to effective dose coefficients

- Conversion coefficients from fluence to effective dose are implemented for three different irradiation geometries:
  - ◆ anterior-posterior
  - ◆ rotational
  - ◆ WORST (“Working Out Radiation Shielding Thicknesses”) is the maximum coefficient of anterior-posterior, posterior-anterior, right-lateral and left-lateral geometries. It is recommended to be used for shielding design.
- Implemented for radiation weighting factors recommended by ICRP60 (e.g., **SDUM=ETW74**) and recommended by M.Pelliccioni (e.g., **SDUM=EWTMP**). The latter anticipate the 2007 recommendations of ICRP.
- Implemented for **protons, neutrons, charged pions, muons, photons, electrons** (conversion coefficients for other particles are approximated by these)
- **Zero** coefficient is applied to all **heavy ions**

# Fluence to effective dose coefficients



# USRYIELD

- Scores a **double-differential particle yield** around an extended or a point target.
- “Energy-like” quantities

Kinetic energy , total momentum , total energy , longitudinal momentum in the lab frame ,  
longitudinal momentum in the c.m.s. frame LET

- “Angle-like” quantities (in degrees or radians)

Rapidity in the lab frame , rapidity in the c.m.s. frame , pseudorapidity in the lab frame ,  
pseudorapidity in the c.m.s. frame , Feynman-x in the lab frame ,  
Feynman-x in the c.m.s. frame , transverse momentum , transverse mass ,  
polar angle (\*) in the lab frame , polar angle (\*) in the c.m.s. frame ,  
square transverse momentum , charge , weighted angle in the lab frame ,  
weighted transverse momentum