The slide features a decorative layout of blue lines. A vertical line on the left and a horizontal line at the top meet at a corner marked with a small blue circle. Another horizontal line is positioned below the top one. A vertical line on the right and a horizontal line at the bottom meet at a corner marked with another small blue circle. The word "Sources" is written in a bold, purple font between the top and middle horizontal lines.

Sources

Advanced FLUKA Course

Overview

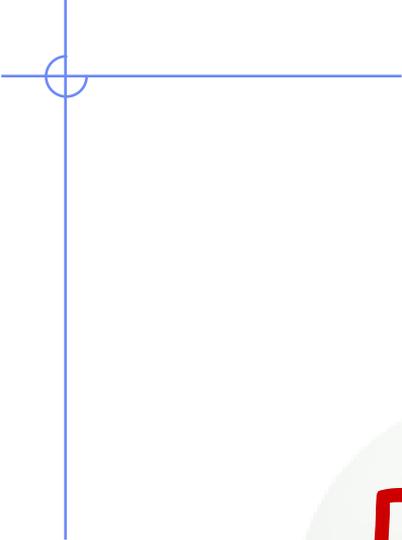
1. Built-in sources

- Beam definition
- Extended sources
- Sources for collider experiments

2. User-defined sources

- User routine SOURCE
- Useful auxiliary routines
- Sampling techniques
- Two-step methods

3. Example: point vs. extended source



Built-in sources

Beam definition - 1

Input card: **BEAM**

defines several *beam characteristics*:
type of particle, energy, divergence, profile

Example

```
* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .  
BEAM          3.5 -0.082425          -1.7          0.0          0.0          0.0 PROTON
```

- 3.5 GeV/c [**WHAT (1)**] proton beam [**SDUM**] with weight 1 [**WHAT (6)**]
- Gaussian momentum distribution: 0.082425 GeV/c FWHM [**WHAT (2)**]
- Gaussian angular distribution: 1.7 mrad FWHM [**WHAT (3)**]
- no beam width along x (point-like source) [**WHAT (4)**]
- no beam width along y (point-like source) [**WHAT (5)**]

Beam definition - 2

Input card: **BEAMPOS**

If **SDUM** = blank:

defines the **coordinates of the centre of the beam spot** (*i.e.*, the point from which transport starts) and the **beam direction**

Example

* . . . + 1 + 2 + 3 + 4 + 5 + 6 + 7 +
BEAMPOS 0.0 0.0 -0.1 0.0 0.0 0.0

- x-coordinate: 0.0 [**WHAT (1)**]
 - y-coordinate: 0.0 [**WHAT (2)**]
 - z-coordinate: -0.1 cm [**WHAT (3)**]
 - direction cosine with respect to the x-axis: 0.0 [**WHAT (4)**]
 - direction cosine with respect to the y-axis: 0.0 [**WHAT (5)**]
 - **WHAT (6)** is not used !
- beam points in the positive z-direction starting at (0.,0.,-0.1)

Beam definition - 3

Input card: **BEAMAXES**

defines the **beam reference frame** which all parameters defined with BEAM and BEAMPOS refer to (angular divergence, transverse profile, polarization, extended sources)

Example

* . . . + 1 + 2 + 3 + 4 + 5 + 6 + 7 +
BEAMAXES 1.0 0.0 0.0 0.0 0.7071068 0.7071068

- cosine of angle between x-axis of beam and x-axis of geometry frame [WHAT (1)]
- cosine of angle between x-axis of beam and y-axis of geometry frame [WHAT (2)]
- cosine of angle between x-axis of beam and z-axis of geometry frame [WHAT (3)]
(1.,0,0) → x-axes of beam and geometry frames are parallel

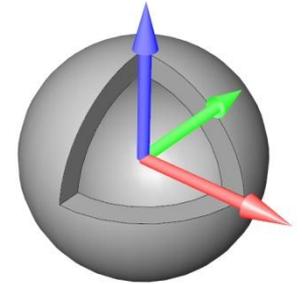
- cosine of angle between z-axis of beam and x-axis of geometry frame [WHAT (4)]
- cosine of angle between z-axis of beam and y-axis of geometry frame [WHAT (5)]
- cosine of angle between z-axis of beam and z-axis of geometry frame [WHAT (6)]
(0.,0.7071068,0.7071068) → z-axes of beam frame is at 45deg to both y- and z-axes of geometry frame

Extended sources - *Spherical shell source*

Input card: **BEAMPOS**

If **SDUM** = SPHE-VOL:

defines a spatially extended source in a **spherical shell**



Example

* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .							
BEAMPOS		0.0	0.0	0.0	0.0	0.0	0.0
BEAMPOS		0.0	1.0	0.0	0.0	0.0	0.0 SPHE-VOL

- radius (in cm) of the inner sphere shell: 0.0 cm [WHAT (1)]
- radius (in cm) of the outer sphere shell: 1.0 cm [WHAT (2)]
- **WHAT (3) - WHAT (6)** are not used !

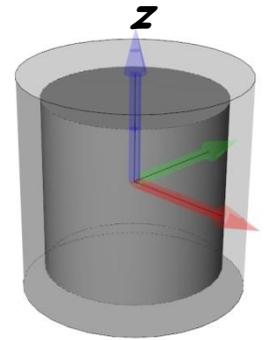
The shell is centred at the (x,y,z) point defined by another BEAMPOS card with **SDUM** = blank (or = **NEGATIVE**). The particle direction or angular distribution are those defined by BEAM, BEAMAXES and another BEAMPOS cards.

Extended sources - *Cylindrical shell source*

Input card: **BEAMPOS**

If **SDUM** = CYLI-VOL:

defines a spatially extended source in a **cylindrical shell** with the height parallel to the z-axis of the beam frame



Example

* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .							
BEAMPOS	0.0	0.0	0.0	0.0	0.0	0.0	0.0
BEAMPOS	0.0	1.0	0.0	1.0	0.0	0.0	0.0CYLI-VOL

- radius (in cm) of the inner cylinder defining the shell: 0.0 cm [WHAT (1)]
- radius (in cm) of the outer cylinder defining the shell: 1.0 cm [WHAT (2)]
- height (in cm) of the inner cylinder defining the shell: 0.0 cm [WHAT (3)]
- height (in cm) of the outer cylinder defining the shell: 1.0 cm [WHAT (4)]
- **WHAT (5) - WHAT (6)** are not used !

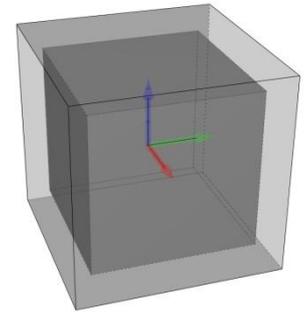
The shell is centred at the (x,y,z) point defined by another BEAMPOS card with **SDUM** = blank (or = **NEGATIVE**). The particle direction or angular distribution are those defined by BEAM, BEAMAXES and another BEAMPOS cards.

Extended sources - Cartesian shell source

Input card: **BEAMPOS**

If **SDUM** = CART-VOL:

defines a spatially extended source in a **Cartesian shell** with the sides parallel to the beam frame axes



Example

* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .						
BEAMPOS	0.0	0.0	0.0	0.0	0.0	0.0
BEAMPOS	0.0	1.0	0.0	1.0	0.0	1.0 CART-VOL

- length (in cm) of the x-side of the inner parallelepiped defining the shell: 0.0 cm [WHAT (1)]
- length (in cm) of the x-side of the outer parallelepiped defining the shell: 1.0 cm [WHAT (2)]
- length (in cm) of the y-side of the inner parallelepiped defining the shell: 0.0 cm [WHAT (3)]
- length (in cm) of the y-side of the outer parallelepiped defining the shell: 1.0 cm [WHAT (4)]
- length (in cm) of the z-side of the inner parallelepiped defining the shell: 0.0 cm [WHAT (5)]
- length (in cm) of the z-side of the outer parallelepiped defining the shell: 1.0 cm [WHAT (6)]

The shell is centred at the (x,y,z) point defined by another BEAMPOS card with **SDUM** = blank (or = **NEGATIVE**). The particle direction or angular distribution are those defined by BEAM, BEAMAXES and another BEAMPOS cards.

Extended sources - *Spherical surface source*

Input card: **BEAMPOS**

If **SDUM** = FLOOD:

defines a source distribution on a **spherical surface**

Example

* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .							
BEAMPOS	0.0	0.0	0.0	0.0	0.0	0.0	0.0
BEAMPOS	0.0	1.0	0.0	0.0	0.0	0.0	0.0FLOOD

- radius (in cm) of the sphere: 1.0 cm [**WHAT (1)**]
- **WHAT (2)** - **WHAT (6)** are not used !

The surface is centred at the (x,y,z) point defined by another BEAMPOS card with **SDUM** = blank (or = **NEGATIVE**). The particle direction or is distributed according to a diffusive distribution, so that to generate a uniform and isotropic fluence equal to $1/\pi R^2$ everywhere in the sphere (in absence of materials)

Extended sources - Example

Radioactive source of ^{60}Co (two main γ -emissions: 1332.5 keV and 1173.2 keV)

cylindrical shape, 2cm diameter, 2mm height along z, centre of base of cylinder at origin

```
*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...
BEAM          0.0                                     ISOTOPE
HI-PROPE      27.0          60.0
BEAMPOS       0.0          0.0          0.1          0.0          0.0          0.0
BEAMPOS       0.0          1.0          0.0          0.2          0.0          0.0CYLI-VOL
```

or

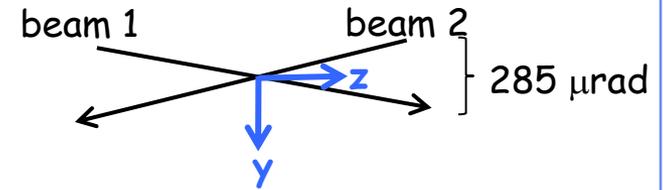
```
*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...
BEAM          1252.8E-6          10000.                PHOTON
BEAMPOS       0.0          0.0          0.1          0.0          0.0          0.0
BEAMPOS       0.0          1.0          0.0          0.2          0.0          0.0CYLI-VOL
```

If height along x (instead of z) add

```
*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...
BEAMAXES      0.0          0.0          -1.0          1.0          0.0          0.0
```

Special sources - hadron-nucleus collision

Input card: **SPECSOUR**



Example: LHC

7 TeV/c, full crossing angle of 285 μ rad in yz-plane

Momentum vectors of colliding proton beams: three possibilities

1) If **SDUM** = PPSOURCE:

SPECSOUR	0.	0.9975	6999.9999	0.0	0.9975-6999.9999	PPSOURCE
----------	----	--------	-----------	-----	------------------	----------

- x, y, z-components of lab momentum for proton beam 1 [WHAT (1-3)]
- x, y, z-components of lab momentum for proton beam 2 [WHAT (4-6)]

2) If **SDUM** = CROSSASY:

SPECSOUR	7000.	142.5E-6	90.0	7000.	142.5E-6	0.0	CROSSASY
----------	-------	----------	------	-------	----------	-----	----------

- lab momentum for proton beam 1 [WHAT (1)]
- polar angle (rad) between proton beam 1 and positive z-direction [WHAT (2)]
- azimuth angle (deg!) defining crossing plane [WHAT (3)]
- lab momentum for proton beam 2 [WHAT (4)]
- polar angle (rad) between proton beam 2 and positive z-direction [WHAT (5)]

Special sources - hadron-nucleus collision

3) If SDUM = CROSSSYM:

*...+...1...+...2...+...3...+...4...+...5...+...6...+...7...+...
SPECSOUR 7000. 142.5E-6 90.0 0.0 0.0 0.0CROSSSYM

- lab momentum for proton beams 1 and 2 [WHAT (1)]
- half crossing angle (rad) [WHAT (2)]
- azimuth angle (deg!) defining crossing plane [WHAT (3)]
- WHAT (4) - WHAT (6) are not used !

Interaction point of colliding proton beams (continuation card):

SPECSOUR 7000. 142.5E-6 90.0 0.0 0.0 0.0CROSSSYM
SPECSOUR 12.E-5 12.E-5 5.0 &

- sigma_x in cm for Gaussian sampling around XBEAM: 12 mm [WHAT (7)]
- sigma_y in cm for Gaussian sampling around YBEAM: 12 mm [WHAT (8)]
- sigma_z in cm for Gaussian sampling around ZBEAM: 5 cm [WHAT (9)]

(XBEAM,YBEAM,ZBEAM) defined with BEAMPOS card

Special sources - hadron-nucleus collision

BEAM	7000.0						HEAVYION
HI-PROPE	82.0	208.0					
...							
SPECSOUR	7000.	142.5E-6	90.0	0.0	0.0	0.0	CROSSSYM
SPECSOUR	12.E-5	12.E-5	5.0	-2.0	208.0	82.0	&

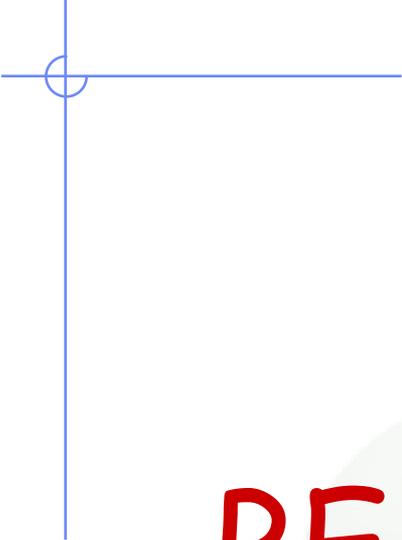
- id of beam particle 1 (default the one of BEAM) [WHAT (10)]
- mass of beam particle 2 (default 1) [WHAT (11)]
- charge of beam particle 2 (default 1) [WHAT (12)]

For collisions in the DPMJET energy range, don't forget the following cards

* . . . + . . . 1 . . . + . . . 2 . . . + . . . 3 . . . + . . . 4 . . . + . . . 5 . . . + . . . 6 . . . + . . . 7 . . . + . . .			
EVENTYPE		2.0	DPMJET
PHYSICS	8000.0		LIMITS

Where "8000.0" should be larger than $\frac{1}{2}$ of the centre-of-mass energy (the energy of whichever beam for a head-on collision in a symmetric collider)

WARNING: In the present release, EM-dissociation is not implemented at source collision level (e.g. Dominant contribution for LHC Pb-Pb)



BEAM Visualization



Within geometry viewer

For the moment the beam position and direction can be plotted with arrows inside the flair geometry editor.

- Add a #define to define the beam length

```
#define bl 50.0
```

- Add an !arrow card and set as whats the following functions:

```
x:      =c(BEAMPOS,0,1)
```

```
y:      =c(BEAMPOS,0,2)
```

```
z:      =c(BEAMPOS,0,3)
```

```
dx:     =bl*c(BEAMPOS,0,4)
```

```
dy:     =bl*c(BEAMPOS,0,5)
```

```
dz:     =bl*sqrt(1.0-c(BEAMPOS,0,4)**2-c(BEAMPOS,0,5)**2)
```

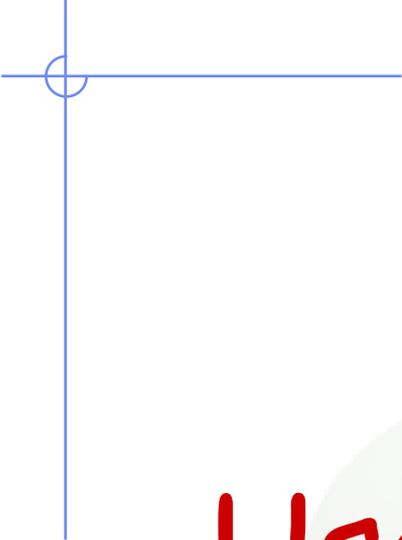
`c(BEAMPOS,n,m)` is a function that returns from the n^{th} (zero based) BEAMPOS card the m^{th} argument

USRBIN

- Create a usrbins covering the beam position (preferentially X-Y-Z)
- Set as scoring particle the BEAMPART
- Setting all materials to vacuum (to speed up calculation)
- Make one run of 1 cycle
- Visualize the results:
 - in flair as USRBIN plot
 - in the geometry editor as a custom USRBIN layer (don't forget to set properly the colorband)

With USERDUMP

- Add a USERDUMP card selecting ONLY Source particles
- Create a USERDUMP plot in flair
- Select the “Source” tab
- You have the ability to make
 - 1D histogram plots of any of the source quantities
 - 2D scattered plots for any of the quantities with even the possibility to overlay on an geometry image



User-defined sources



Source routine - 1

- Allows the **definition of primary particle properties** (in space, energy, time, direction or mixture of particles) which cannot be described with built-in sources
- Activated with **input card SOURCE**. The parameter list of that card (two continuation cards possible!) allows the user to pass on up to 18 numerical values **WHASOU (1-18)** and one 8-character string **SDUSOU** via **COMMON /SOURCM/**
- At each call, one (or more) particle(s) must be loaded onto **COMMON /FLKSTK/** (particle bank) before returning control. These values can be read from a file, generated by some sampling algorithm, or just assigned.
- **Argument list**: if **NOMORE=1** (output variable) the run will be terminated after exhausting the primary particles loaded onto the stack in the present call. The history number limit set with card **START** will be overridden.

Source routine - 2

```
...  
    LOGICAL LFIRST  
*  
    SAVE LFIRST  
    DATA LFIRST / .TRUE. /  
...  
    NOMORE = 0  
* +-----*  
* | First call initializations:  
* | IF ( LFIRST ) THEN  
* | *** The following 3 cards are mandatory ***  
* |     LFIRST = .FALSE.  
* |     TKESUM = ZERZER  
* |     LUSSRC = .TRUE.  
* | *** User initialization ***
```

Any **first-time initialization** can be inserted here, for example

- setting up parameters passed on via SOURCE card
- reading spectra from data files

```
END IF
```

```
...
```

Source routine - 3

```
...
  NPFLKA = NPFLKA + 1
* Wt is the weight of the particle
  WTFLK (NPFLKA) = ONEONE
  WEIPRI = WEIPRI + WTFLK (NPFLKA)
* Particle type (1=proton.....). Ijbeam is the type set by the BEAM
* card
* +-----*
* | (Radioactive) isotope:
  IF ( IJBEAM .EQ. -2 .AND. LRDBEA ) THEN
    IARES = IPROA
    IZRES = IPROZ
    IISRES = IPROM
    CALL STISBM ( IARES, IZRES, IISRES )
    IJHION = IPROZ * 1000 + IPROA
    IJHION = IJHION * 100 + KXHEAV
    IONID = IJHION
    CALL DCDION ( IONID )
    CALL SETION ( IONID )
* |
* +-----*
* | Heavy ion:
  ELSE IF ( IJBEAM .EQ. -2 ) THEN
    IJHION = IPROZ * 1000 + IPROA
    IJHION = IJHION * 100 + KXHEAV
    IONID = IJHION
    CALL DCDION ( IONID )
    CALL SETION ( IONID )
    ILOFLK (NPFLKA) = IJHION
* | Flag this is prompt radiation
  LRADC (NPFLKA) = .FALSE.
* | Group number for "low" energy neutrons, set to 0 anyway
  IGROUP (NPFLKA) = 0
* |
* +-----*
* | Normal hadron:
  ELSE
    IONID = IJBEAM
    ILOFLK (NPFLKA) = IJBEAM
* | Flag this is prompt radiation
  LRADC (NPFLKA) = .FALSE.
* | Group number for "low" energy neutrons, set to 0 anyway
  IGROUP (NPFLKA) = 0
  END IF
* |
* +-----*
...
```

increase pointer in FLKSTK

weight of particle (if ≠ 1 biased source)
total weight of primaries (don't change)

Definition of particle type

- The template sets the type of particle equal to the one defined by the BEAM card (and HI-PROPE, if used).

- Whichever valid particle type can be set inside the source (may be different event by event)

Source routine - 4

```
...
* Particle age (s)
  AGESTK (NPFLKA) = +ZERZER
  AKNSHR (NPFLKA) = -TWOTWO
* Kinetic energy of the particle (GeV)
  TKEFLK (NPFLKA) = SQRT ( PBEAM**2 + AM (IONID)*:
* Particle momentum
  PMOFLK (NPFLKA) = PBEAM
* Cosines (tx,ty,tz)
  TXFLK (NPFLKA) = UBEAM
  TYFLK (NPFLKA) = VBEAM
  TZFLK (NPFLKA) = WBEAM
*   TZFLK (NPFLKA) = SQRT ( ONEONE - TXFLK (NPFLKA)*
    & - TYFLK (NPFLKA)**2 )
* Particle coordinates
  XFLK (NPFLKA) = XBEAM
  YFLK (NPFLKA) = YBEAM
  ZFLK (NPFLKA) = ZBEAM
* Polarization cosines:
  TXPOL (NPFLKA) = -TWOTWO
  TYPOL (NPFLKA) = +ZERZER
  TZPOL (NPFLKA) = +ZERZER
...
```

momentum and energy

- here: taken from BEAM card (PBEAM in COMMON /BEAMCM/)
- the set can set any momentum or energy here (either from file or sampled)
- **NOTE:** BEAM card is still mandatory for initialization purposes. Momentum and energy set here must not be larger than those defined with the BEAM card.

direction cosines and coordinates

- here: taken from BEAMPOS card (PBEAM in COMMON /BEAMCM/)
- ensure proper normalization of cosines!

polarization

- TXPOL=-2 flag for "no polarization"

Source routine - 5

* User dependent flag:

```
LOUSE (NPFLKA) = 0
```

* User dependent spare variables:

```
DO 100 ISPR = 1, MKBMX1
```

```
    SPAREK (ISPR,NPFLKA) = ZERZER
```

```
100 CONTINUE
```

* User dependent spare flags:

```
DO 200 ISPR = 1, MKBMX2
```

```
    ISPARK (ISPR,NPFLKA) = 0
```

```
200 CONTINUE
```



Variables that allow to store additional information in COMMON /FLKSTK/, such as **information on ancestors** of a certain particle

Auxiliary routines - *Random numbers*

... = **FLRNDM** (XDUMMY)

returns a **64-bit random number [0-1)**

NOTE: Fundamental for SOURCE! No other external random generators must be used, otherwise the history reproducibility will be lost.

CALL FLNRRN (RGAUSS)

returns a **normally distributed random number** RGAUSS

CALL FLNRR2 (RGAUS1, RGAUS2)

returns an **uncorrelated pair of normally distributed random numbers** RGAUS1 and RGAUS2

CALL SFECFE (SINT, COST)

returns SINT and COST, sine and cosine of a **random azimuth angle**
 $SINT^{**2} + COST^{**2} = 1.D+00$

CALL RACO (TXX, TYY, TZZ)

returns a **random 3D direction** (TXX, TYY, TZZ) such that:
 $TXX^{**2} + TYY^{**2} + TZZ^{**2} = 1.D+00$

Auxiliary routines - *Name/number conv.*

Conversion of **region name to number**

CALL GEON2R (REGNAM, NREG, IERR)

Input variable:

Regnam = region name (CHAR*8)

Output variables:

Nreg = region number

Ierr = error code (0 on success, 1 on failure)

Conversion of **region number to name**

CALL GEOR2N (NREG, REGNAM, IERR)

Input variable:

Nreg = region number

Output variables:

Regname = region name (CHAR*8)

Ierr = error code (0 on success, 1 on failure)

Auxiliary routines - Others

CALL OAUXFI ('file' , LUN , 'CHOPT' , IERR)

to **open an auxiliary file** (to read data or parameters) looking automatically for the file in some default locations (temporary directory, working directory)

CALL FLABRT ('name' , 'message')

this allows to force a **FLUKA abort on user request**: it might be useful to perform a debugging (using gdb for instance)

CALL SFLOOD (XXX , YYY , ZZZ , UXXX , VYYY , WZZZ)

returns in XXX, YYY, ZZZ a **random position ON the surface of a sphere** of radius 1 and centre 0 (multiply XXX, YYY, ZZZ by the actual radius and add the centre coordinates) and UXXX, VYYY, WZZZ are random cosines distributed so as to generate a uniform and isotropic fluence inside the sphere numerically given by $1/(\pi R^2)$, R being the sphere radius.

Sampling from a distribution - *Discrete*

1) From the cumulative distribution

- Suppose to have a *discrete* random variable x , that can assume values $x_1, x_2, \dots, x_n, \dots$ with probability $p_1, p_2, \dots, p_n, \dots$
- Assume $\sum_i p_i = 1$, or normalize it
- Divide the interval $[0,1)$ in n subintervals, with limits

$$y_0 = 0, y_1 = p_1, y_2 = p_1 + p_2, \dots$$

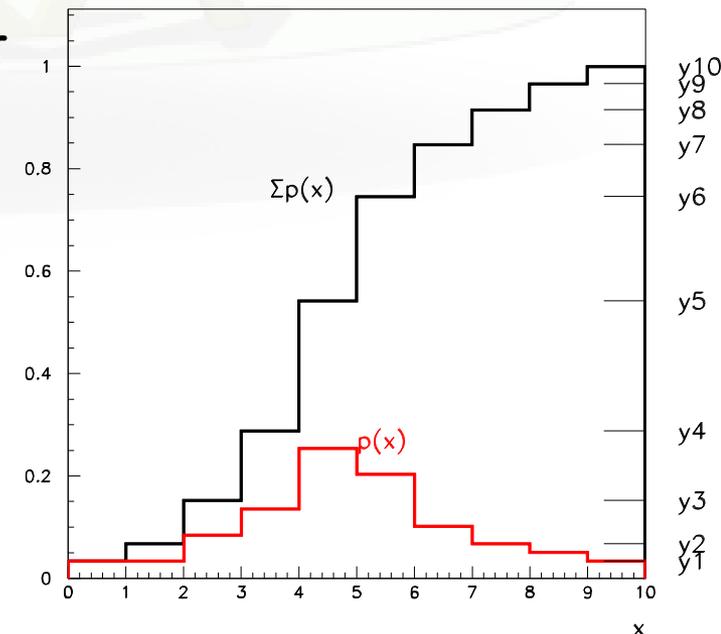
- Generate a uniform pseudo-random number ξ
- Find the interval i^{th} y -interval such that

$$y_{i-1} \leq \xi < y_i$$

- Select $X = x_i$ as the sampled value

Since ξ is uniformly random:

$$P(x_i) = P(y_{i-1} \leq \xi < y_i) = y_i - y_{i-1} = p_i$$



Sampling from a distribution - *Discrete*

2) By adjusting weights

- Suppose to have an fluence energy spectrum Φ given in N discrete energy bins between E_0 and E_N : Φ_1, \dots, Φ_N
- Generate a uniform pseudo-random number ξ
- Find the interval i^{th} energy bin such that
$$E_{i-1} \leq \xi (E_N - E_0) < E_i$$
- Generate another uniform pseudo-random number $\xi \in [0,1)$ and sample an energy uniformly within the i^{th} energy bin
- assign a weight Φ_i to that primary particle

Note: This method is often used for spectra steeply decreasing with energy (e.g., $\Phi \sim 1/E$), where the result depends significantly on the particle cascades cause by high energy primaries, as it ensures faster convergence to the mean value.

Example Sampling from a histogram - 1

```
PARAMETER (NMAX=1000)
DIMENSION ERG(NMAX), CUM(NMAX)
CHARACTER*250 LINE
SAVE N, ERG, CUM

IF ( LFIRST ) THEN
...
  LUNRD = NINT(WHASOU(1))
  N = 0
  SUM = ZERZER
  EPREV = ZERZER
10  CONTINUE
    READ (LUNRD, '(A)', ERR=9999 END=20 ) LINE
    READ (LINE, *, ERR=10) E, H
    N = N + 1
    IF (N .GT. NMAX)
&      CALL FLABRT('SOURCE', 'Please increase NMAX')
    IF (N .EQ. 1 .AND. ABS(H) .GT. AZRZRZ)
&      CALL FLABRT(
&        'SOURCE', 'ZERO was was expected as first value')
*** Create cumulative sum of events! dN=dE*V
    SUM = SUM + H*(E-EPREV)
    EPREV = E
    ERG(N) = E
    CUM(N) = SUM

    GO TO 10
20  CONTINUE
    CLOSE (LUNRD)
END IF
```

Logical unit from input file
Use **OPEN** input card to open the file
Pass the unit with as what1 in SOURCE
The file contains
pairs Energy Value
WARNING first value should have be 0
in order to define the lower energy limit

Build cumulative sum (as a histogram
NOT as flux)

Example Sampling from a histogram - 2

* From this point

*** Select a random energy interval

```
C = CUM(N) * FLRNDM(C)
```

*** Find interval (no need to check first interval CUM=0)

```
DO I=2,N
```

```
  IF (CUM(I) .GT. C) THEN
```

*** Found interval I, select a random energy inside

```
  E = ERG(I-1) + (ERG(I)-ERG(I-1))*FLRNDM(C)
```

```
  GO TO 90
```

```
END IF
```

```
END DO
```

Select a random position

Find interval in cumulative sum

Select a random Energy inside the interval

Sampling from a distribution - *Continuous*

1) By integration

- Integrate the distribution function $f(x)$, analytically or numerically, and normalize to 1 to obtain the **normalized cumulative distribution**

$$F(\xi) = \frac{\int_{x_{\min}}^{\xi} f(x)dx}{\int_{x_{\min}}^{x_{\max}} f(x)dx}$$

- Generate a uniform pseudo-random number $\xi \in [0,1)$
- Get the desired result by finding the **inverse value** $X = F^{-1}(\xi)$, **analytically** or most often numerically, i.e. by **interpolation** (table look-up)

Since ξ is uniformly random:

$$P(a < x < b) = P(F(a) \leq \xi < F(b)) = F(b) - F(a) = \int_a^b f(x)dx$$

Sampling from a distribution - *Continuous*

Example

Take $f(x) = e^{-\frac{x}{\lambda}}$, $x \in [0, \infty)$

Cumulative distribution:

$$F(t) = \int_0^t e^{-\frac{x}{\lambda}} dx = \lambda \times \left(1 - e^{-\frac{t}{\lambda}} \right)$$

Normalized:

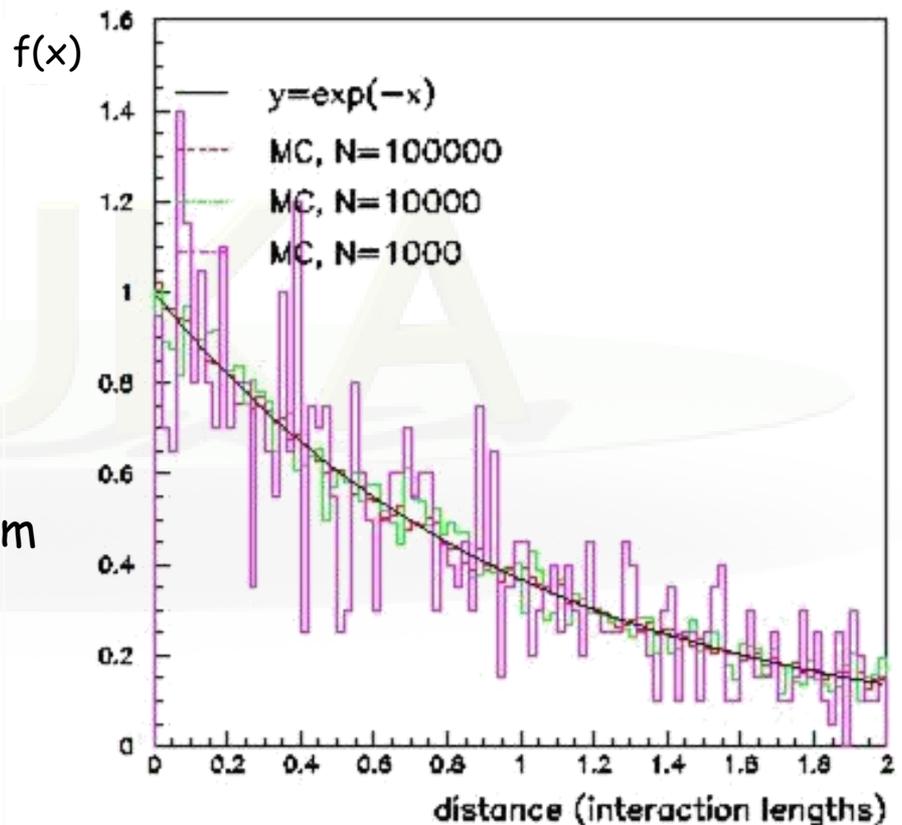
$$F'(t) = \int_0^t \frac{e^{-\frac{x}{\lambda}}}{\lambda} dx = 1 - e^{-\frac{t}{\lambda}}$$

Generate a uniform pseudo-random number $\xi \in [0,1)$

Sample t by inverting

$$t = -\lambda \ln(1 - \xi)$$

Repeat N times



Sampling from a distribution - *Continuous*

2) By rejection

- Let be $f'(x)$, a normalized distribution function, which cannot be sampled by integration and inversion
- Let be $g'(x)$, a normalized distribution function, which can be sampled, and such that $Cg'(x) \geq f'(x)$, $\forall x \in [x_{\min}, x_{\max}]$
- Sample X from $g'(x)$, and generate a uniform pseudo-random number $\xi \in [0, 1)$
- Accept X if $\xi < f'(X)/Cg'(X)$, if not repeat the previous step
- The overall efficiency (accepted/rejected) is given by:

$$R = \int \frac{f'(x)}{Cg'(x)} g'(x) dx = \frac{1}{C}$$

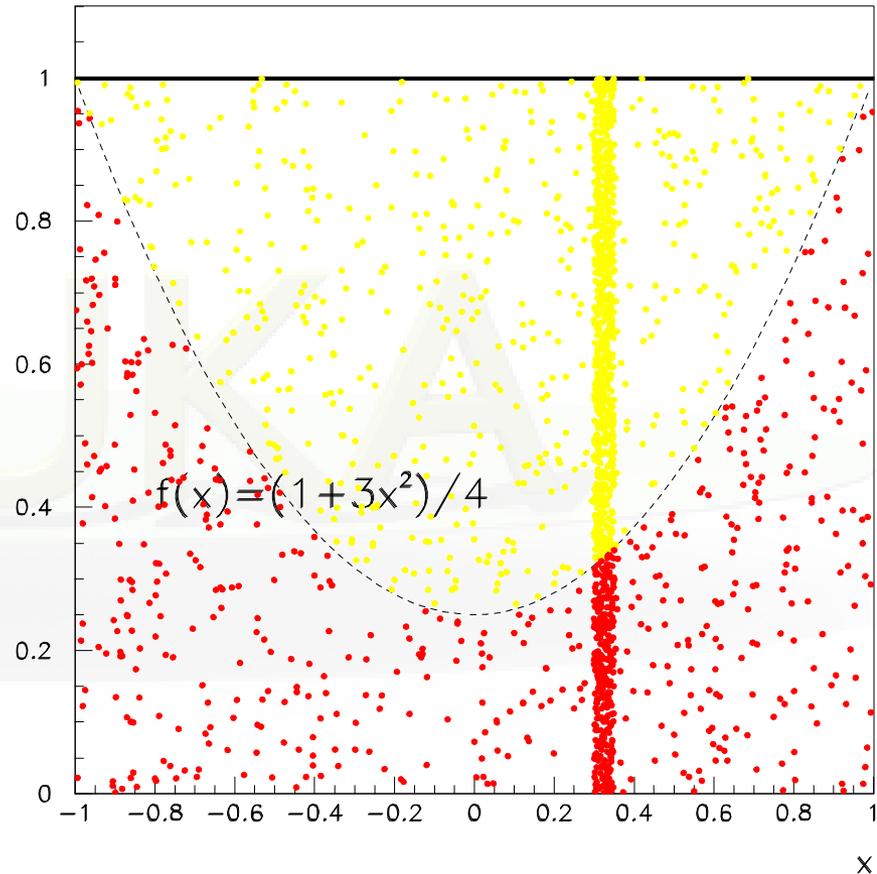
- and the probability that X is accepted is unbiased:

$$P(X)dX = \frac{1}{R} g'(X)dX \times \frac{f'(X)}{Cg'(X)} = f'(X)dX$$

Sampling from a distribution - *Continuous*

Example

- Let be $f(x) = (1+3x^2)/4$,
 $x \in [-1,1]$,
- Take $g(x) = 1/2$, $C=2$
- Generate two uniform pseudo-random numbers
 $\xi_1, \xi_2 \in [0,1]$
- Accept $X = 2\xi_1 - 1$ if
 $\xi_2 < (1+3X^2)/4$, if not
repeat



Sampling from a distribution - *Continuous*

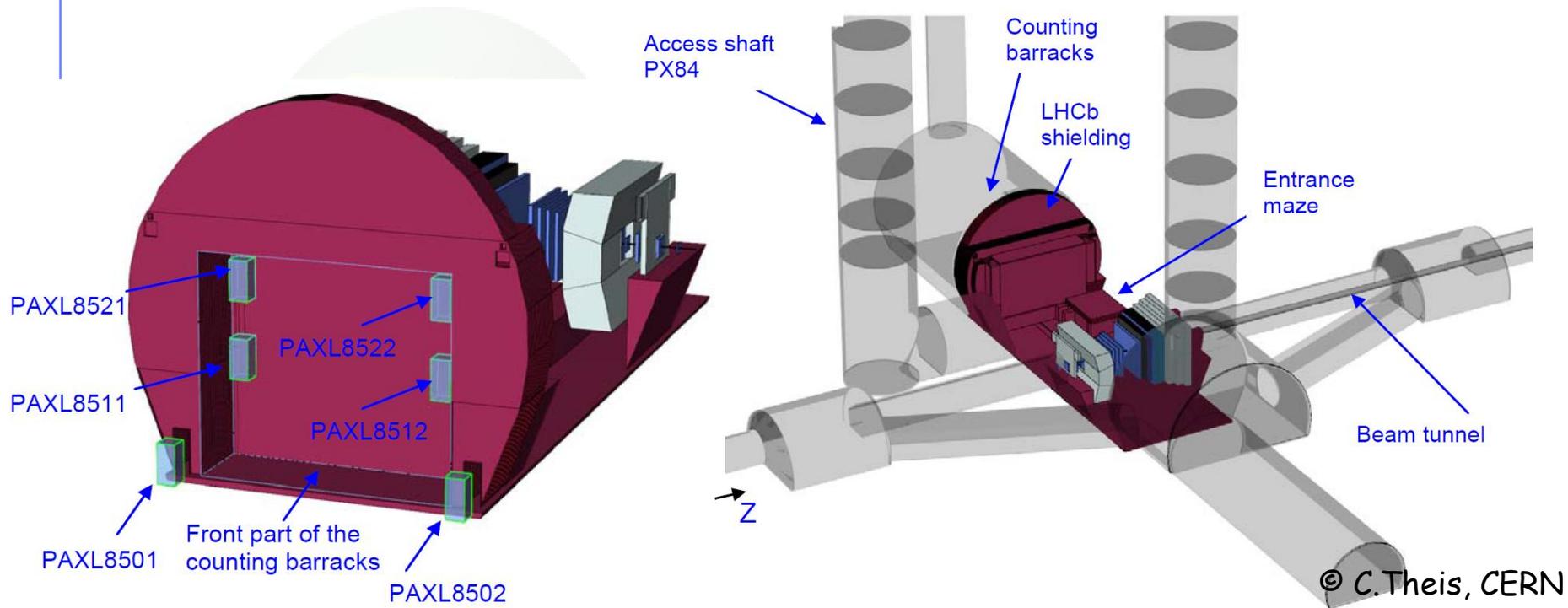
3) By adjusting weights

- Suppose to have a fluence energy spectrum $\Phi(E)$ given in between E_0 and E_1
- Generate a uniform pseudo-random number $\xi \in [0,1)$ and calculate the sampled energy $E = E_0 + \xi (E_1 - E_0)$
- Assign a weight $\Phi(E)$ to that primary particle

Two-step methods

Example:

predict reading of a (small) radiation detector at a remote location in a huge geometry, *e.g.*, LHCb experiment at CERN



Problem: direct calculation in one step highly inefficient due to the small affected phase-space

Two-step methods

- Solution:** split simulation into two steps
- 1) Calculation of radiation field at detector location
 - 2) Simulation of detector reading

Two options:

Directional dependence of detector reading is negligible

- calculate **average fluence energy spectra**, separately for different particle types, at the detector location
- simulate reading of detector with user-defined source which reads in the calculated spectra and samples particle type, energy and direction (*e.g.*, isotropic incidence)
- **important:** results of the second step have to be **normalized to the integrated particle fluence** obtained in the first step

Directional dependence of detector reading is important

- replace detector by 'blackhole' and write all information on particles entering it (type, energy, position, direction) into an **external file**
- simulate reading of detector (if possible with the original geometry now containing the detector) with user-defined source which reads in the particles from the external file
- **important:** **pick entries randomly** from external file to avoid going through identical sequence of particles if several runs are performed

Two-step example - Dumping particles

You can dump the particles with several ways e.g:

- `mgdraw.f` activated with `USERDUMP`
- `fluscw.f` activated with `USERWEIGHT`

The following example is using the `fluscw.f` activated with `USERWEIGHT` and coupled with the first `USRBDX` scoring

```
* Activate with USERWEIGHT Use FLUSCW+ WHAT(3)>2 to
* Couple scoring with the first Boundary crossing estimator
  IF (ISCRNG.EQ.1 .AND. JSCRNG.EQ.1) THEN
    IF (LFIRST) THEN
      WRITE (99,*)
&   '# 1.IJ  2.X 3.Y 4.Z  5.TX 6.TY 7.TZ   8.E 9.W'
      LFIRST = .FALSE.
    END IF
    WRITE (99, '(I3,8(1X,F22.14))')
&
&   IJ,XX,YY,ZZ,TXX,TYY,TZZ,-PLA,WEE
  END IF
```

Two-step example - Sampling particles - 1

```
PARAMETER (NMAX=1000000)
SAVE LFIRST
DATA LFIRST / .TRUE. /
CHARACTER*250 LINE
INTEGER      NNN, IJ(NMAX)
DIMENSION   XXX(NMAX), YYY(NMAX), ZZZ(NMAX)
DIMENSION   UUU(NMAX), VVV(NMAX), WWW(NMAX)
DIMENSION   ERG(NMAX), WGT(NMAX)
SAVE XXX, YYY, ZZZ
SAVE UUU, VVV, WWW
SAVE ERG, WGT
```

```
IF ( LFIRST ) THEN
  LUNRD = NINT(WHASOU(1))
  NNN = 0
```

10

```
CONTINUE
  READ( LUNRD, '(A)', ERR=9999, END=20 ) LINE
  READ (LINE,*,ERR=10) I, X, Y, Z, U, V, W, E, WG
  NNN = NNN + 1
  IF (NNN.GT.NMAX) CALL FLABRT('SOURCE', 'Increase NMAX')
  ...
```

Logical unit from input file
Use **OPEN** input card to open the file
Pass the unit with as what1 in SOURCE

Two-step example - Sampling particles - 2

```
      IJ(NNN) = I
      XXX(NNN) = X
      YYY(NNN) = Y
      ZZZ(NNN) = Z
* | Normalize direction to 1.0
      UVW = SQRT(U**2 + V**2 + W**2)
      UUU(NNN) = U / UVW
      VVV(NNN) = V / UVW
      WWW(NNN) = W / UVW
      ERG(NNN) = E
      WGT(NNN) = WG
GOTO 10
20 CONTINUE
IF (NNN.EQ.0) CALL FLABRT('SOURCE','Error reading file')
WRITE (LUNOUT,*)
WRITE (LUNOUT,*) '*** rdsorce: ',NNN,' particles loaded'
WRITE (LUNOUT,*)
END IF
```

Normalize direction

Two-step example - Sampling particles - 3

```
RNDSIG = FLRNDM (RNDSIG)
```

```
N = INT (NNN*RNDSIG)+1
```

* Wt is the weight of the particle

```
WTFLK (NPFLKA) = WGT (N)
```

```
ILOFLK (NPFLKA) = IJ (N)
```

Choose a random particle

Has the benefit of reusing the recorded particles and all results will be normalized per recorded particle

* Kinetic energy of the particle (GeV)

```
* TKEFLK (NPFLKA) = SQRT ( PBEAM**2 + AM (IONID)**2 ) - AM  
(IONID)
```

```
TKEFLK (NPFLKA) = ERG (N)
```

* Particle momentum

```
PMOFLK (NPFLKA) = SQRT ( TKEFLK (NPFLKA) * ( TKEFLK (NPFLKA)  
& + TWOTWO * AM (IONID) ) )
```

Push particle into stack

* Cosines (tx,ty,tz)

```
TXFLK (NPFLKA) = UUU (N)
```

```
TYFLK (NPFLKA) = VVV (N)
```

```
TZFLK (NPFLKA) = WWW (N)
```

* Particle coordinates

```
XFLK (NPFLKA) = XXX (N) !+ XBEAM
```

```
YFLK (NPFLKA) = YYY (N) !+ YBEAM
```

```
ZFLK (NPFLKA) = ZZZ (N) !+ ZBEAM
```

Two-step: Normalization and Errors

- The dumped particles represent only a fraction of the full shower → therefore the **second step consists only of a subset of the full simulation**
- Thus the results of the second step should be multiplied (normalized) with the recorded weight of the **first step**

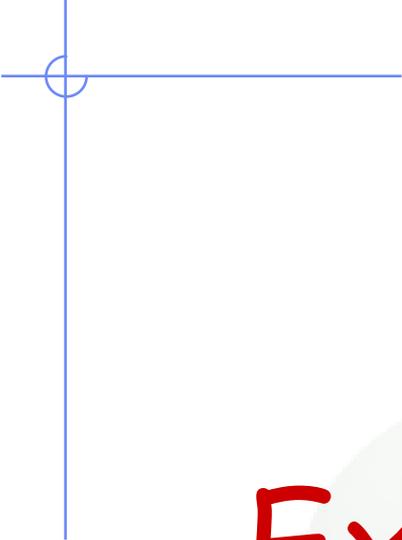
$$\text{Normalization} = \frac{\sum \text{weights of recorded particles}}{\sum \text{weights of source particles}}$$

WARNING:

- verify that the recorded particles contains **ALL** the possible ones that contribute to the effect under study.
You didn't miss any other that could have an impact on the results

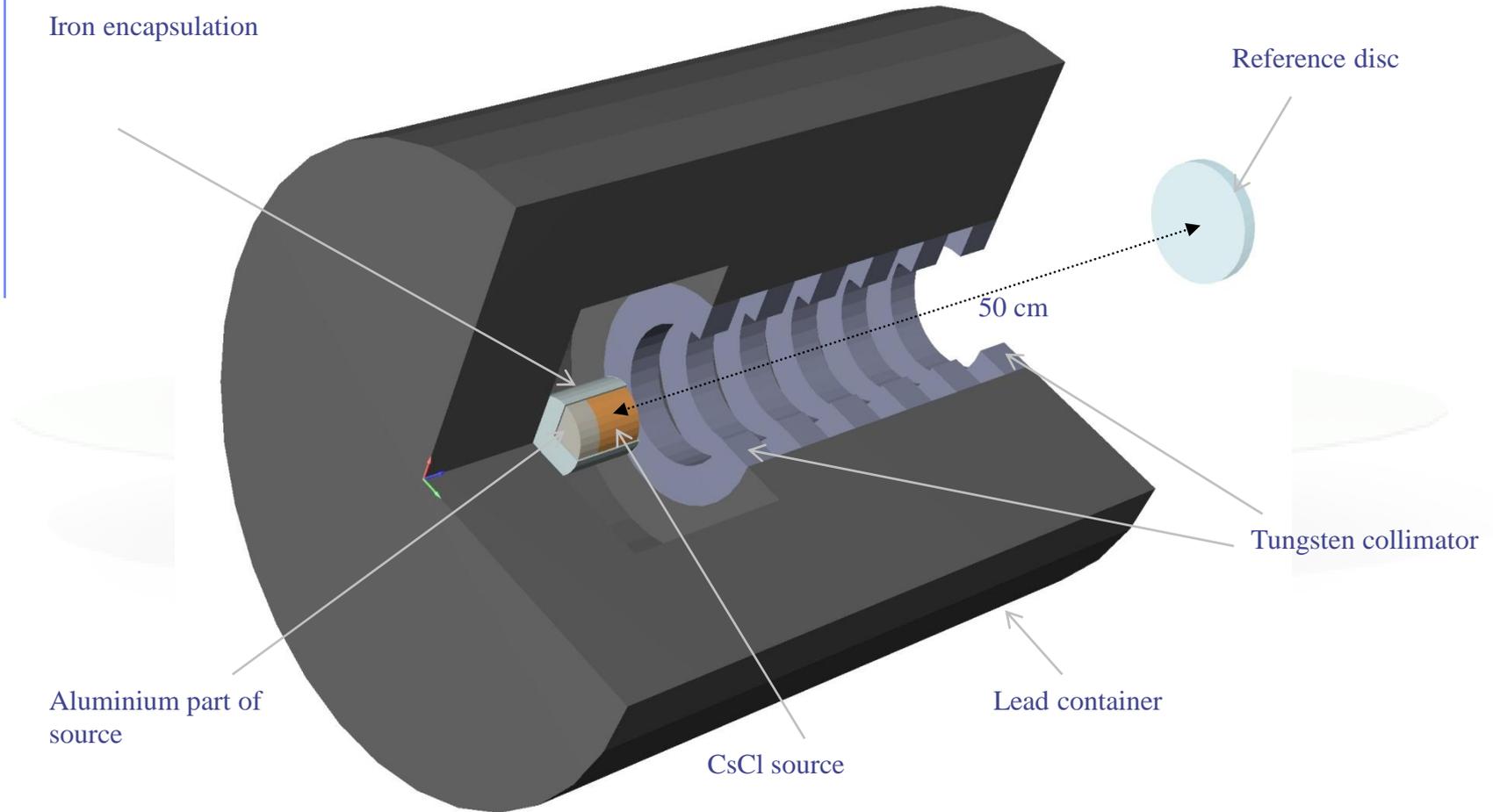
Two-step: Things to remember

1. **Sample randomly** the recorded particles on the second step! It has many benefits: *i)* you don't have to go through the entire list sequentially (especially if enormous), *ii)* you can reuse particles *iii)* results will be normalized per record particle
2. **Verify your NORMALIZATION**
optionally you can make a full run to compare the results between the two step and full run
3. Like in a biasing run the purpose of a two step approach is to **keep the mean but to reduce the error or time or study different configurations.**
4. A more honest two step approach will be to record several, cycles (e.g. the typical 5 cycles) independently from the 1st step, and run separate 2nd steps one for each cycle.
Merging the results will provide a more honest estimation of the variance
5. Verify that no other source of particles could contribute to your results (or at least is insignificant)

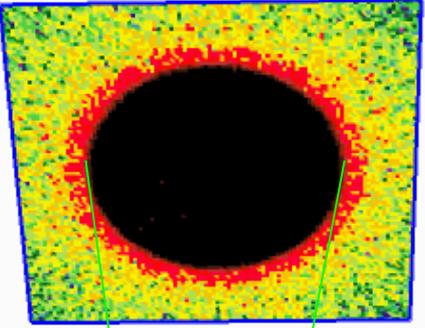
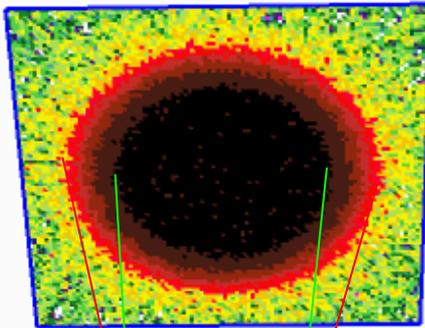


Example:
point vs. extended source

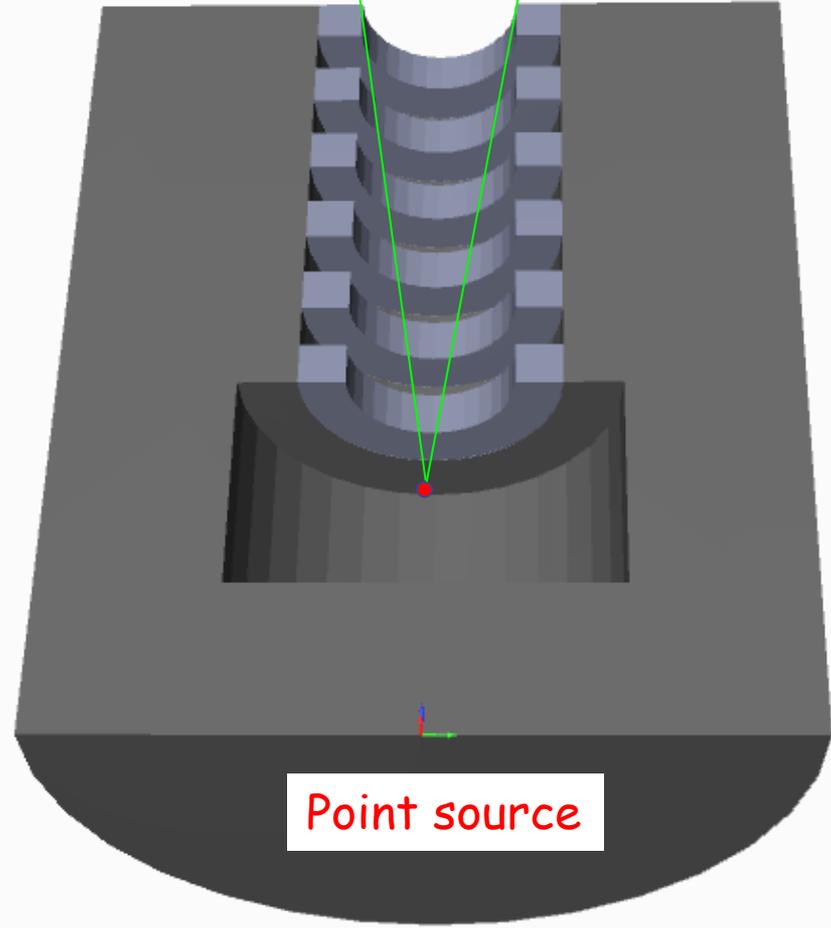
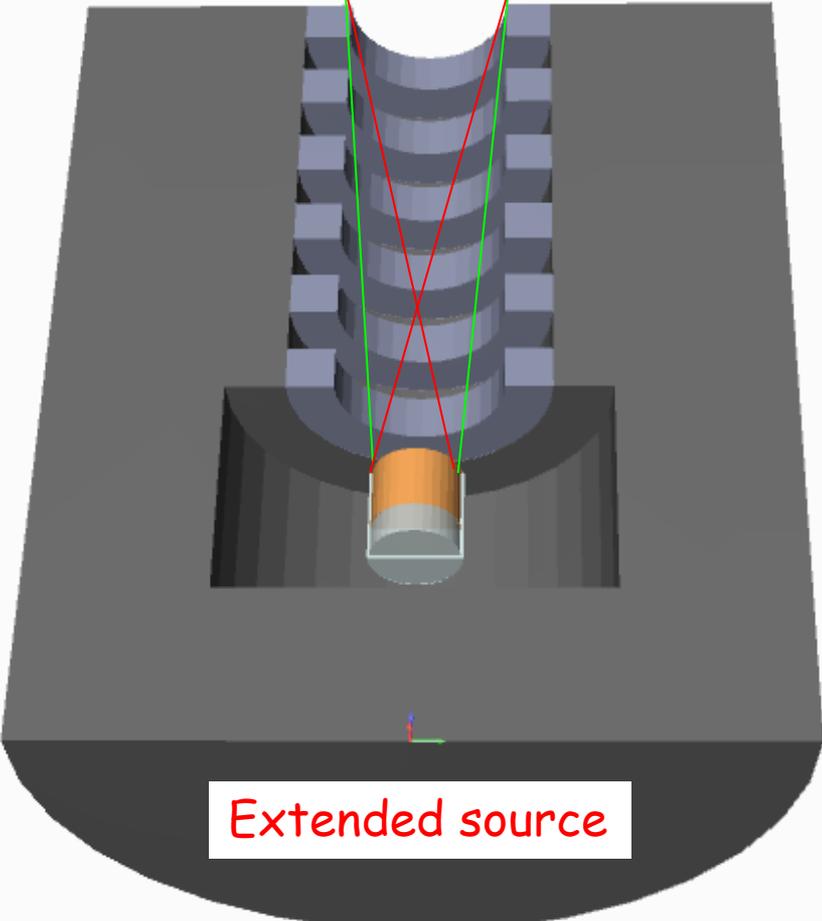
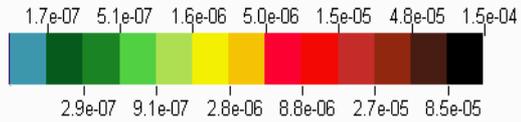
Example - *Cs* irradiator



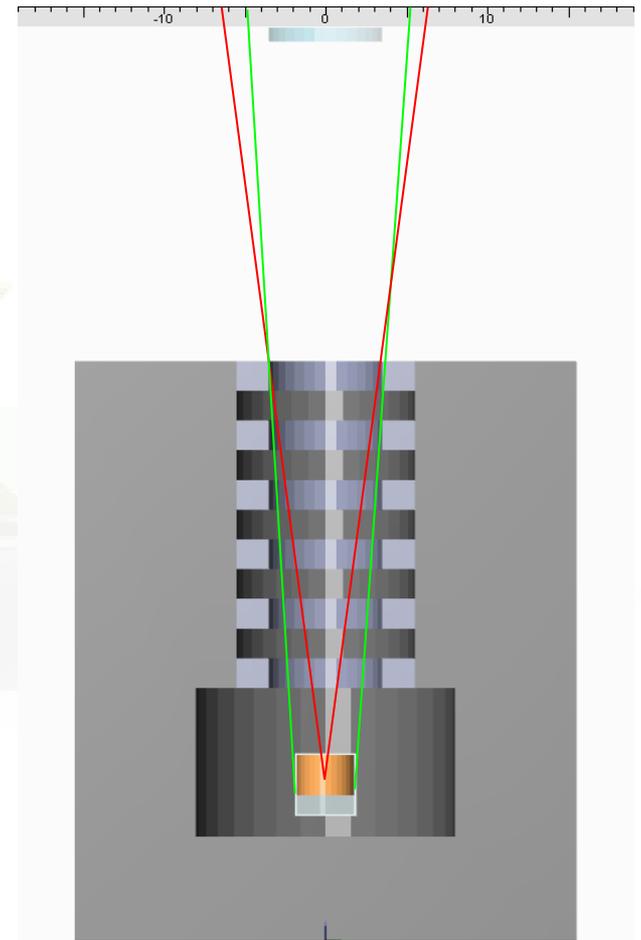
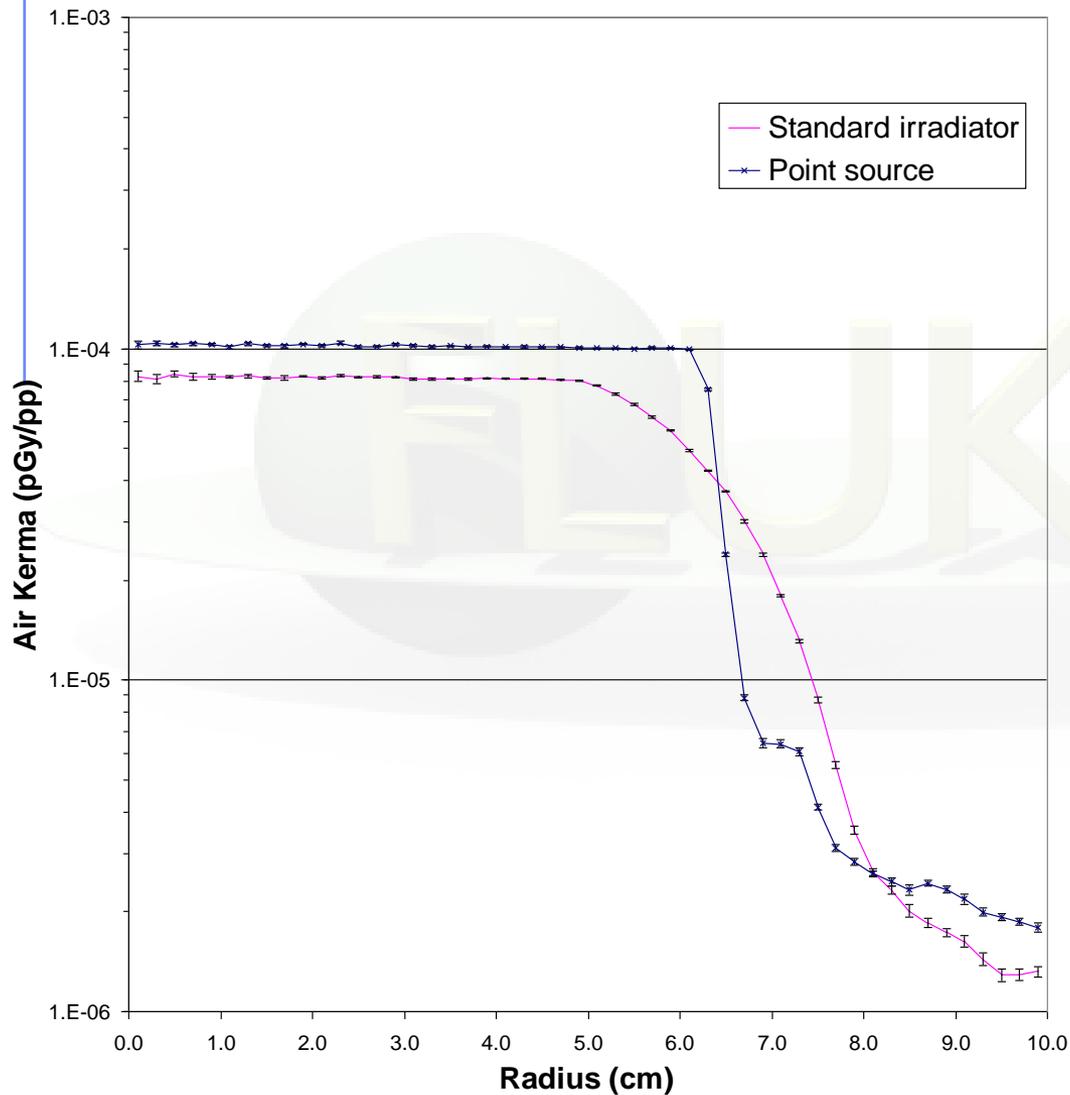
© H.Vincke, CERN



pGy/primary



Example - Cs irradiator



© H.Vincke, CERN